



Probing freeze-in scalar dark matter through decay

Johannes Herms

Technical University of Munich

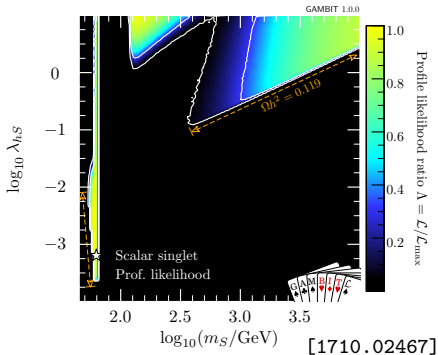
September 14th
Scalars2019, Warsaw

in collaboration with Alejandro Ibarra

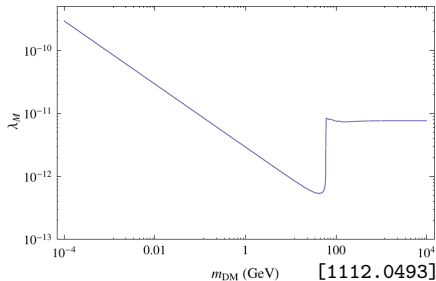


$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 - \frac{1}{4!} \lambda_\phi \phi^4 - \frac{1}{2} \lambda_{\phi h} (H^\dagger H) \phi \phi$$

WIMP DM

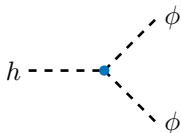


FIMP DM



Relic density from

$$\dot{n}_\phi + 3Hn_\phi = g_h \int \frac{d^3p_h}{(2\pi)^3} \frac{f_h \Gamma_{h \rightarrow \phi X}}{\gamma_h} + 2 \rightarrow 2 \text{ processes}$$



$$\Rightarrow \Gamma_{h \rightarrow \phi\phi} = \frac{\lambda_{\phi h}^2 v^2}{16\pi m_h} \sqrt{1 - \frac{4m^2}{m_h^2}},$$

correct relic density for

$$\lambda_{\phi h} = 9 \cdot 10^{-12} \left(\frac{m}{\text{GeV}} \sqrt{1 - \frac{4m^2}{m_h^2}} \right)^{-1/2}$$

(assuming initial abundance negligible)

\Rightarrow *how can we probe this coupling?*

- Feeble coupling g to heavy bath particle Σ
- Production requires

$$\Gamma_{\text{production}} \sim \frac{g^2}{16\pi} m_{\Sigma} \sim 10^{-25} m \left(\frac{m_{\Sigma}}{m} \right)^2$$

- Decay constraints typically

$$\Gamma_{\text{decay}} \sim \frac{g^2}{16\pi} \frac{m^5}{m_{\Sigma}^4} < 10^{-25} \text{ s}^{-1} \sim 10^{-50} \text{ GeV}$$

$$\Rightarrow m \lesssim 10^{-6} (m_{\Sigma}^3 \text{ GeV})^{1/4}$$

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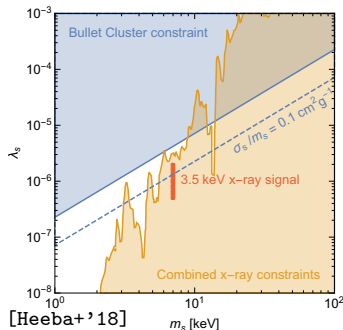
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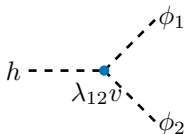
\Rightarrow For γ ray signals at higher energies from FIMP dark matter, need additional suppression of DM-decay.

\rightarrow Felix Kahlhoefer's talk Thursday



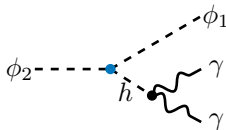
Two scalar singlets $\phi_{1,2}$, odd under a stabilising \mathbb{Z}_2 symmetry:

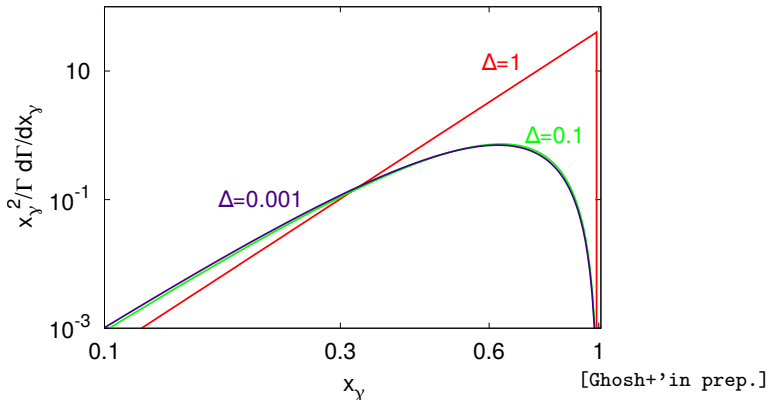
$$\mathcal{L}_{\text{scalar}} = \frac{1}{2} \partial_\mu \phi_i \partial^\mu \phi_i - \frac{1}{2} m_{ij} \phi_i \phi_j - \frac{1}{4} \lambda_{ijkl} \phi_i \phi_j \phi_k \phi_l - \frac{1}{2} \lambda'_{ij} (H^\dagger H) \phi_i \phi_j$$



$$\lambda_{12\text{DM}} \simeq 1.2 \cdot 10^{-11} \left(\frac{m_1 + m_2}{\text{GeV}} \right)^{-1/2} \sqrt{\frac{\Omega_{\lambda_{12}}}{\Omega_{\text{DM}}}}$$

Decay induced by λ_{12} :





$$\Gamma_{\phi_2 \rightarrow \phi_1 \gamma \gamma} = \frac{1}{26880\pi^3} \left(\frac{\lambda_{12} c_{\gamma\gamma}}{m_h^2} \right)^2 m_2^5 \Delta^7 {}_2F_1(3, 4, 8; \Delta),$$

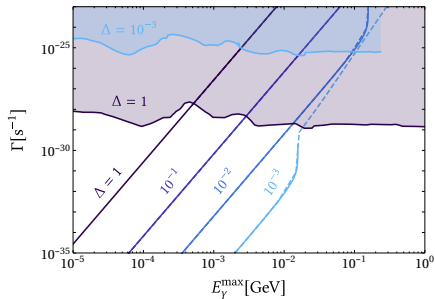
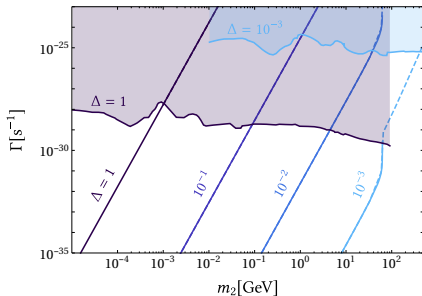
with $c_{\gamma\gamma} = -2.03 \cdot 10^{-3}$ and ${}_2F_1$ a hypergeometric function and mass splitting $\Delta = 1 - m_1^2/m_2^2$. [Ghosh, Ibarra, Mondal, Mukhopadhyaya in prep.]



Multicomponent scalar FIMP in γ -rays

Maximal attainable (ie. $\lambda_{11,22} = 0$) effective decay rate

$$\Gamma_{\text{eff}} = \Gamma_{\phi_2 \rightarrow \phi_1 \gamma \gamma} \cdot \Omega_2 / \Omega_{\text{DM}}$$

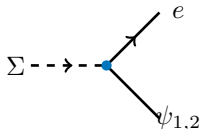


\Rightarrow multicomponent real scalar FIMP can be probed by MeV γ -rays

$$\mathcal{L} \supset (\mathcal{D}_\mu \Sigma)^\dagger (\mathcal{D}^\mu \Sigma) + m_\Sigma \Sigma^\dagger \Sigma + \frac{1}{2} \bar{\psi}_i i \not{\partial} \psi_i - \frac{1}{2} m_i \bar{\psi}_i^c \psi_i + g_{ei} \bar{l}_e P_L \psi_i \Sigma + \text{h.c.}$$

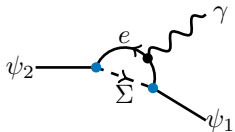
single- $\psi \rightarrow$ Laura Lopez Honorez' talk

- Freeze-in production, dominantly through Σ -decays *



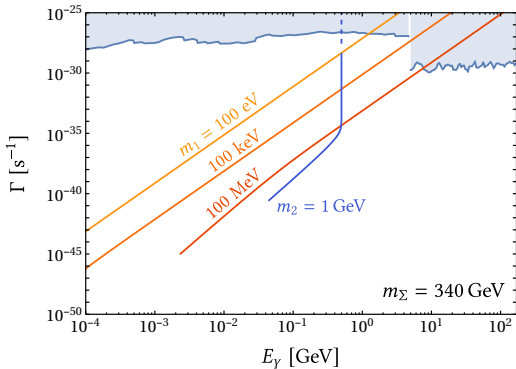
$$g_{ei} \sim 4 \cdot 10^{-12} \sqrt{\frac{m_\Sigma}{m_2} \frac{\Omega_{\text{FIMP}i}}{\Omega_{\text{DM}}}}$$

- Gamma ray lines from dark matter decay

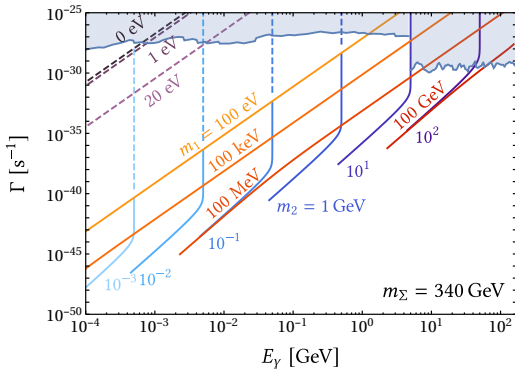


$$\Gamma_{\psi_2 \rightarrow \psi_1 \gamma} = \frac{e^2 (g_{e1} g_{e2})^2}{29 \pi^5} \frac{m_2^5}{m_\Sigma^4} \left(1 - \frac{m_1^2}{m_2^2} \right)^5$$

*2 \rightarrow 2 production relevant, dominates for $m_2 \sim m_\Sigma$. [Garny, Heisig'18], [Junius+'19]



- Gamma ray line limits [Essig+'13], [Fermi-LAT'13]
- $m_\Sigma > 339 \text{ GeV}$ charged track search [CMS'13]
- Fraction of non-cold dark matter and dark radiation [Diamanti+'17]
[Kamada+'19]



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- Multicomponent freeze-in dark matter can be probed by DM-decay
- Portal coupling responsible for abundance controls gamma ray signals
- Real scalar singlet
 - (softened) sawtooth-like γ -ray spectral feature in the MeV range from decay via off-shell Higgs
- Leptophilic fermion
 - γ -ray lines in MeV to TeV range

