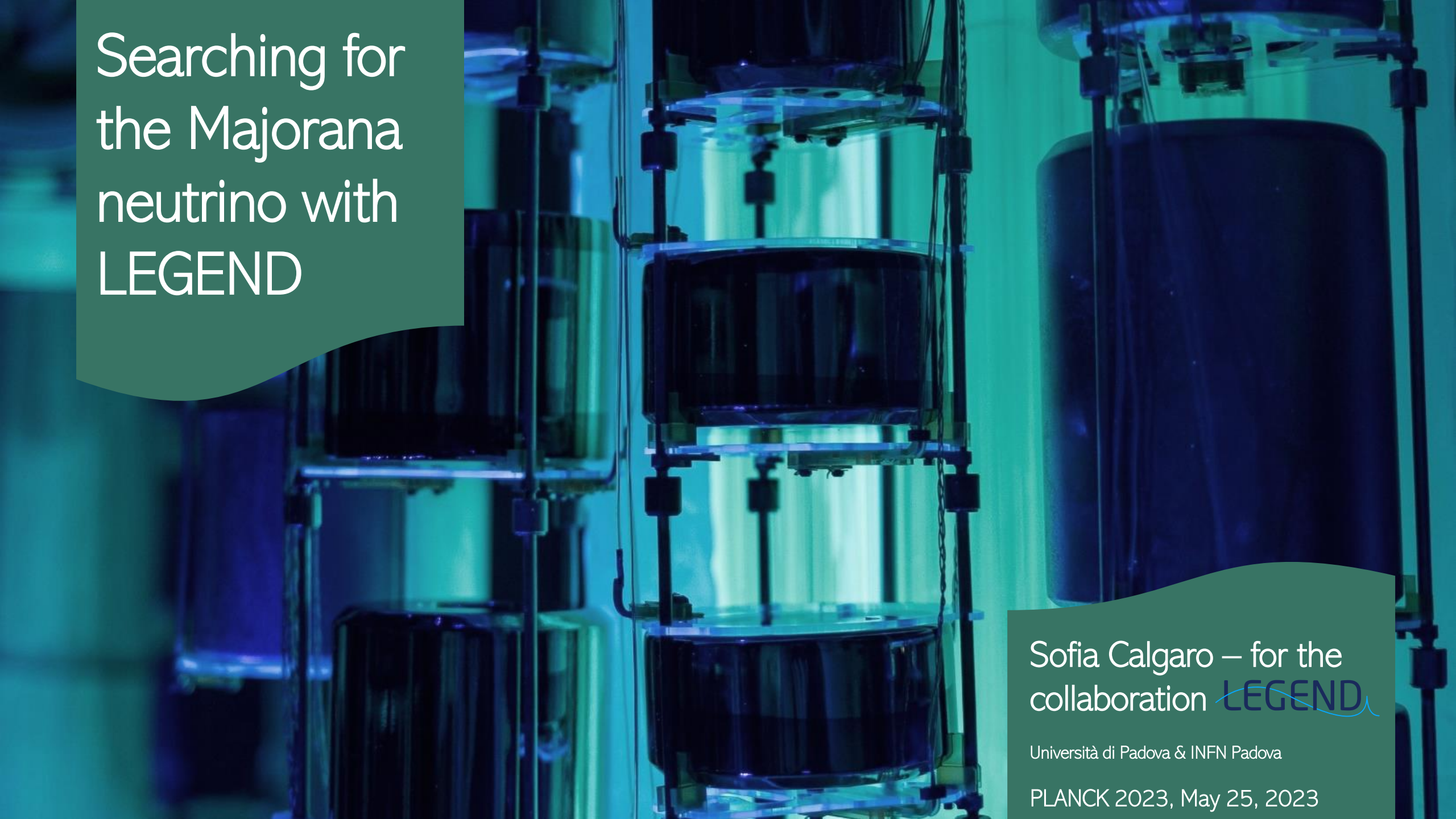


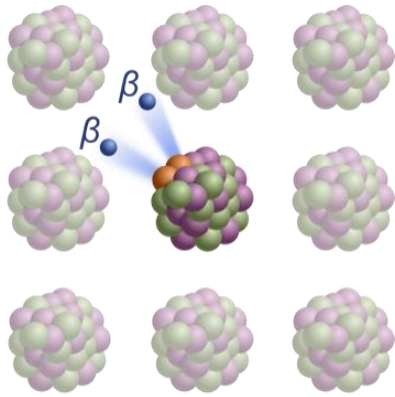


Searching for the Majorana neutrino with LEGEND

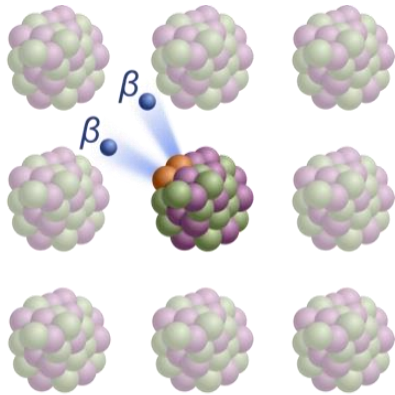
Sofia Calgari – for the
collaboration **LEGEND**

Università di Padova & INFN Padova

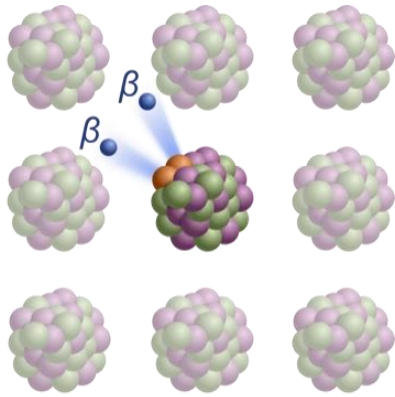
PLANCK 2023, May 25, 2023



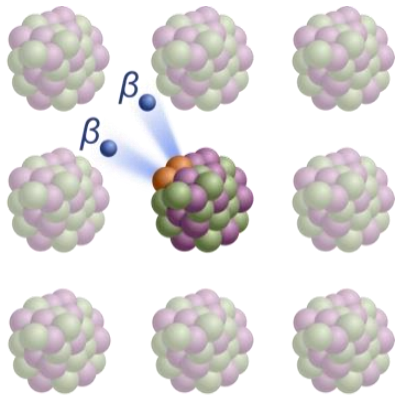
- Neutrinoless double-beta ($0\nu\beta\beta$) decay: the only known, feasible probe of the Majorana nature of neutrinos



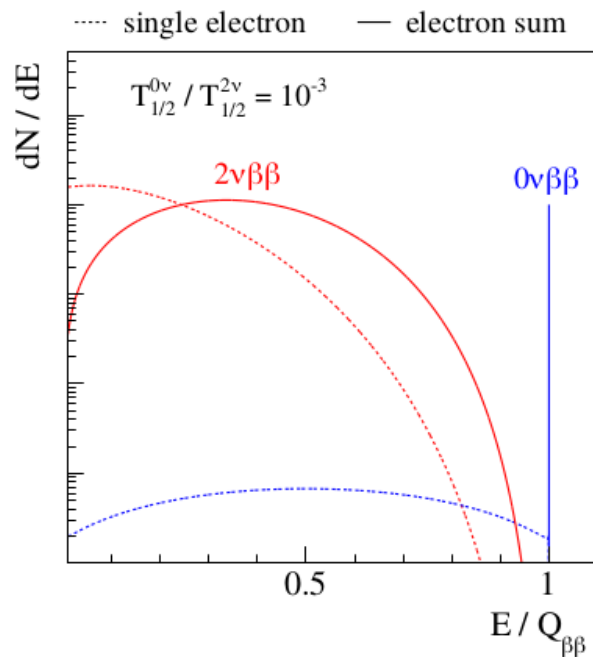
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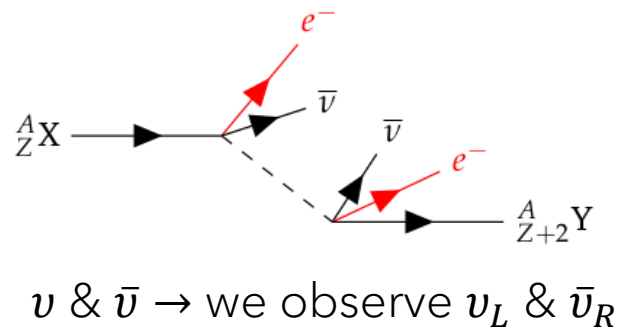
- Neutrinoless double-beta (**$0\nu\beta\beta$**) decay: the only known, feasible probe of the **Majorana nature of neutrinos**
- Its discovery would prove unambiguously 1) lepton-number-violating physics exists 2) that it is connected to the **mysterious origin of the neutrino's mass**
- Explanation for many **BSM theories** and theories that explain **dominance of matter** over antimatter in the Universe



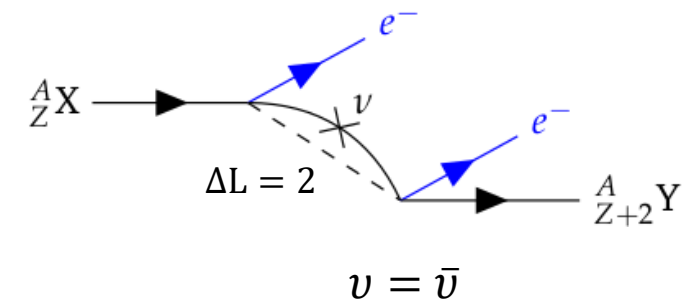
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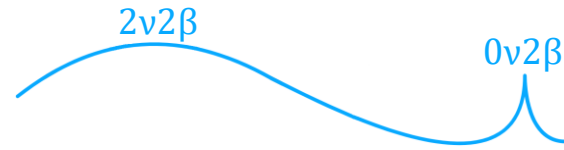


DOUBLE BETA DECAY



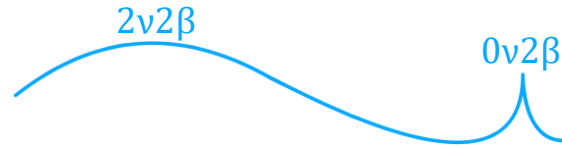
NEUTRINOLESS DOUBLE BETA DECAY





$0\nu 2\beta$ signal

$$M_{iniz} - M_{fin} - 2m_e = Q_{\beta\beta}$$



$0\nu2\beta$ signal

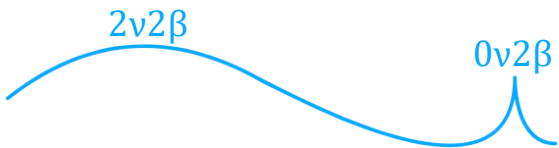
$$M_{iniz} - M_{fin} - 2m_e = Q_{\beta\beta}$$

$0\nu2\beta$ half-life

$$(T_{1/2}^{0\nu})^{-1} = \overset{\text{Phase-space integral,}}{2.36 \cdot 10^{-15} \text{ yr}^{-1}} G^{0\nu}(Q_{\beta\beta}, Z) |M^{0\nu}|^2 \left(\frac{m_{\beta\beta}}{m_e} \right)^2$$

$m_{\beta\beta} = |\sum_i U_{ei}^2 m_i| \rightarrow$ to compare results obtained with different isotopes

*Effective Majorana
neutrino Mass*



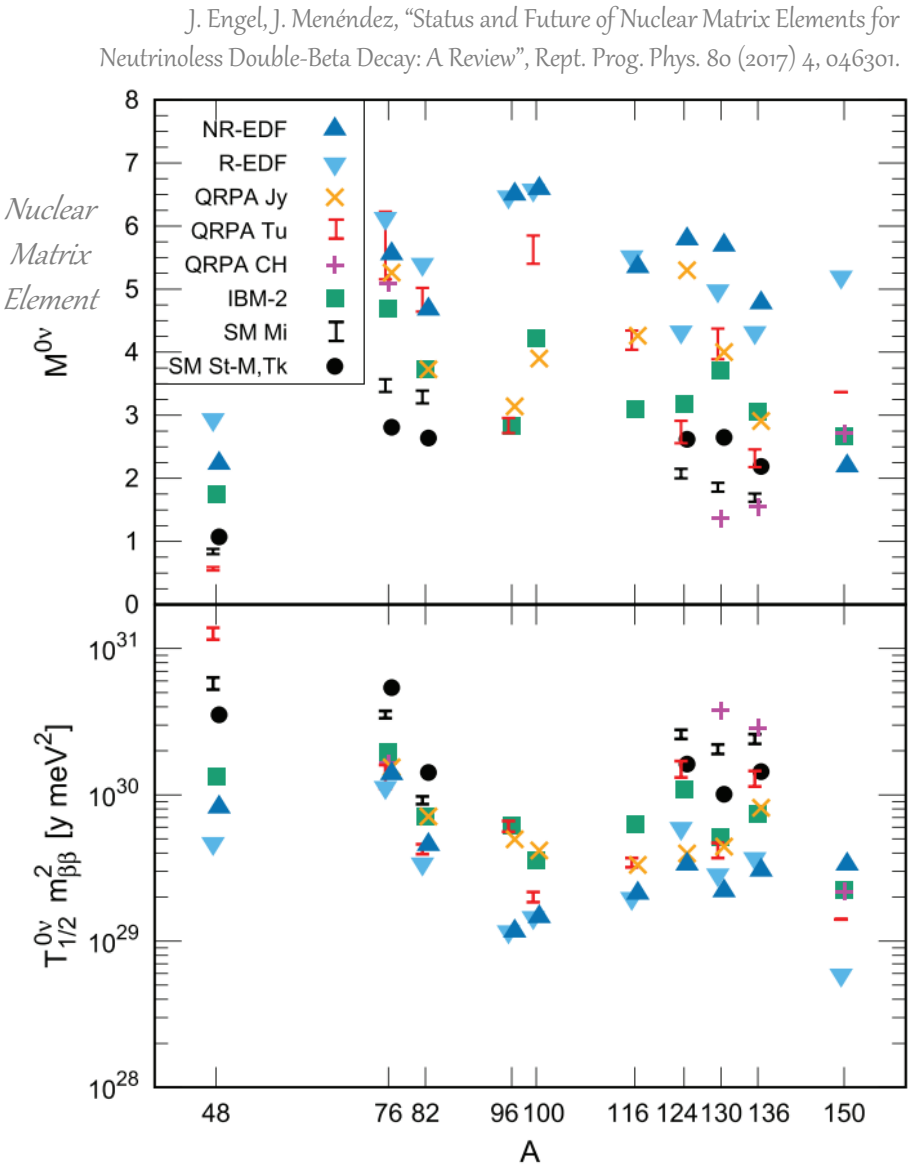
0ν2β signal

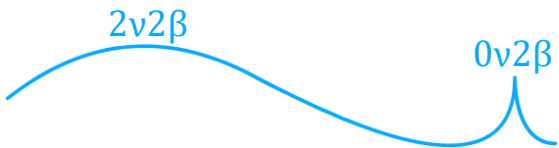
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0ν2β signal

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0ν2β half-life

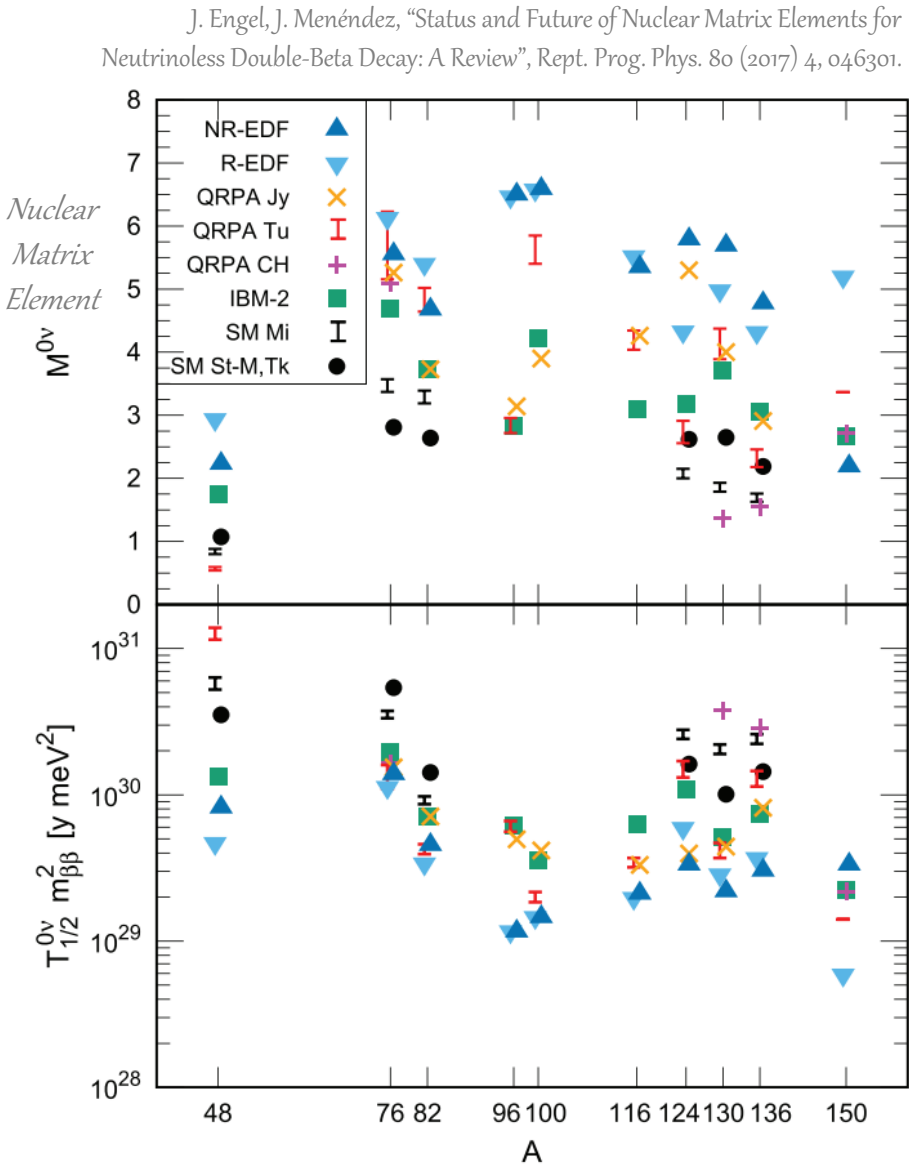
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$m_{\beta\beta} = |\sum_i U_{ei}^2 m_i| \rightarrow$ to compare results obtained with different isotopes
Effective Majorana neutrino Mass

0ν2β half-life sensitivity

$$T_{1/2}^{0\nu} \propto \begin{cases} \epsilon \cdot f \cdot \varepsilon & \text{without bkg} \\ \epsilon \cdot f \cdot \sqrt{\frac{\varepsilon}{BI \cdot \Delta E}} & \text{with bkg} \end{cases}$$

ϵ : efficiency
 f : isotopic fraction
 $\varepsilon = M \cdot t$: exposure
 ΔE : energetic resolution at $Q_{\beta\beta}$
 BI : bkg level





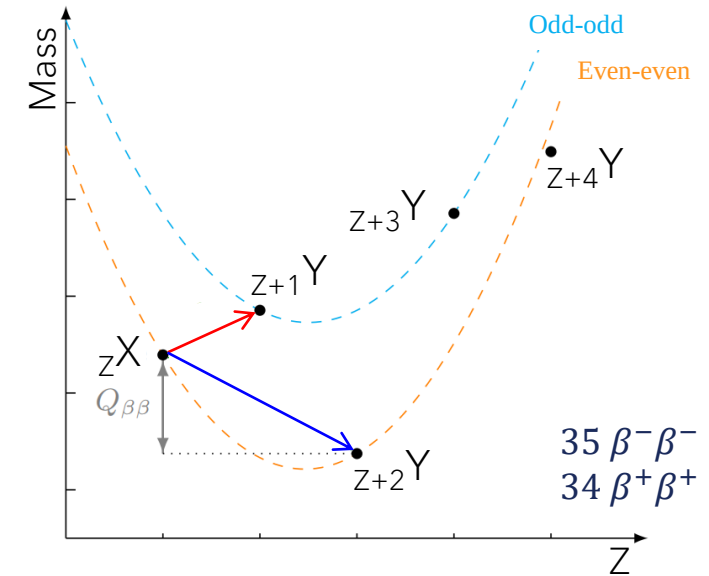
What is the best isotope to use?

No theoretical preferences

- Large theoretical uncertainties
- $G^{0\nu}$ & NME are inversely correlated: tend to compensate

Experimental preferences

- Costs
- Energetic resolution
- Background level
- Scalability (liquids, gas, cristals)

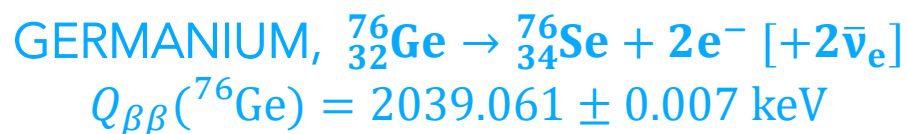
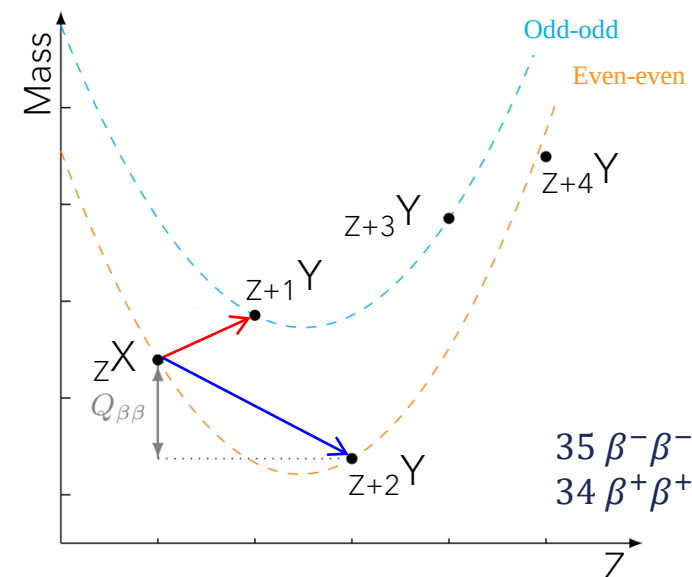


No theoretical preferences

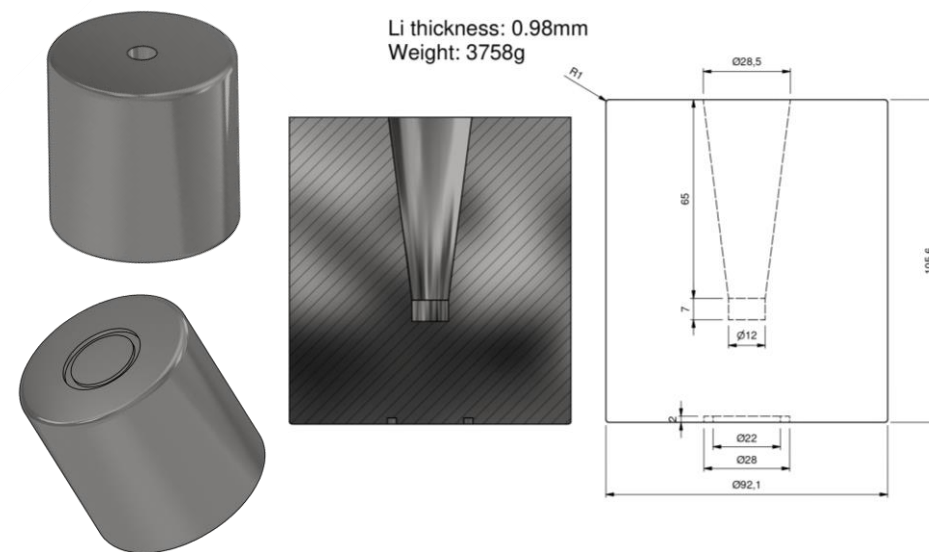
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Experimental preferences

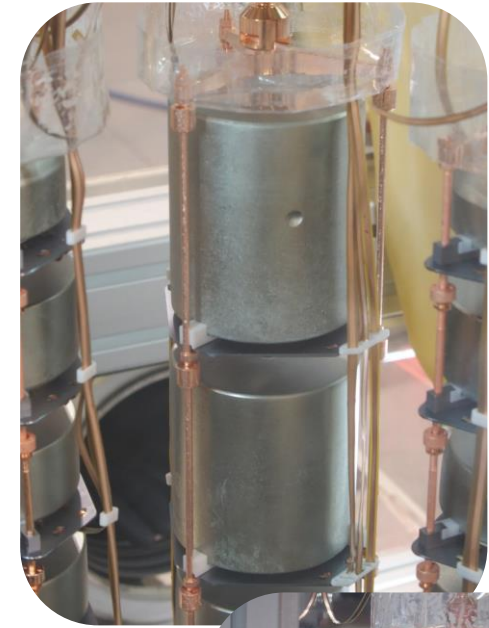
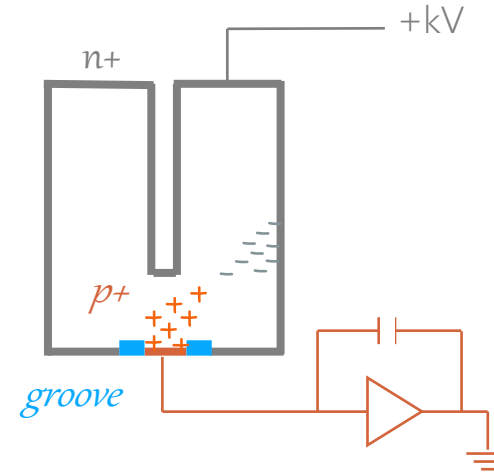
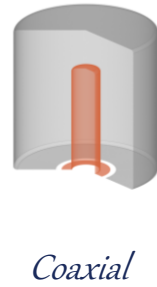
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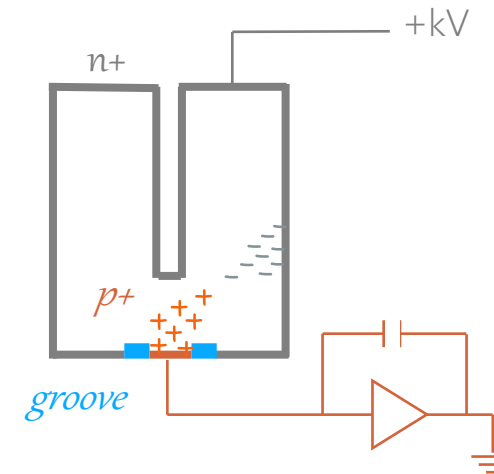
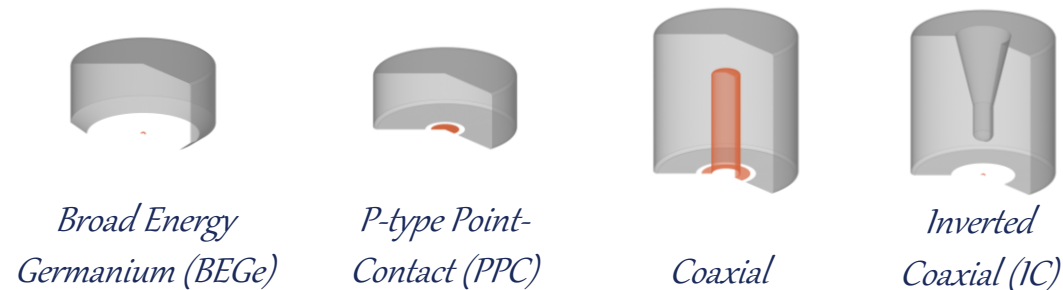
- Already established technology
- Source $\equiv$ detector $\Rightarrow$ **high detection efficiency**
- Ge: high purity $\Rightarrow$ **low intrinsic background**
- Excellent energetic resolution $\Rightarrow$ **FWHM $\sim 0.1\%$ @ $Q_{\beta\beta}$**
- **Low $Q_{\beta\beta} \Rightarrow$ difficult to reach low bkg levels**
- **$f_{76}^{\text{nat}} \sim 8\% \Rightarrow$ the enrichment in Ge-76 is necessary**



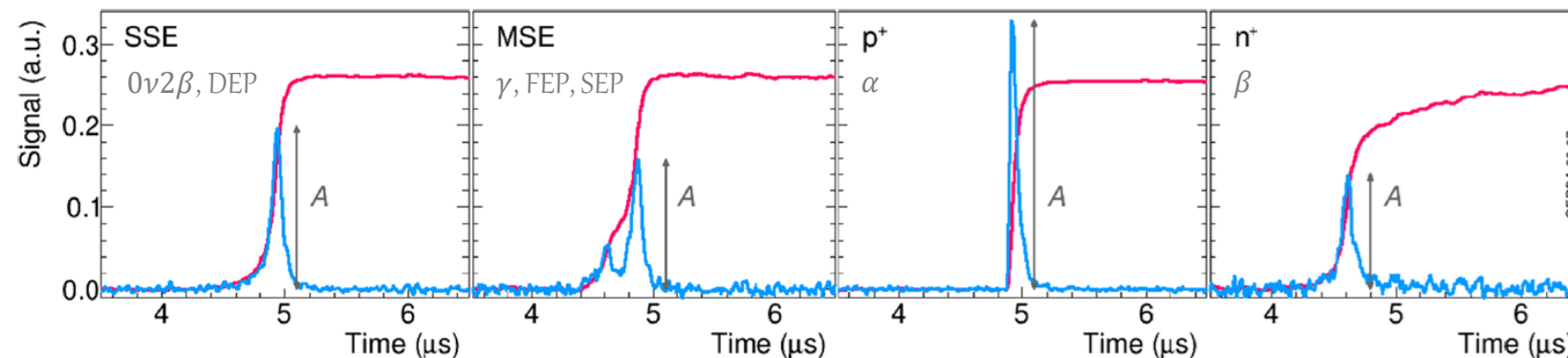
- p+ (implanted B) & n+ (diffused Li), passivated groove
- Full depleted crystals
- Different geometries: BEGe, PPC, coaxials, IC
- Mass: ~0.7-4 kg



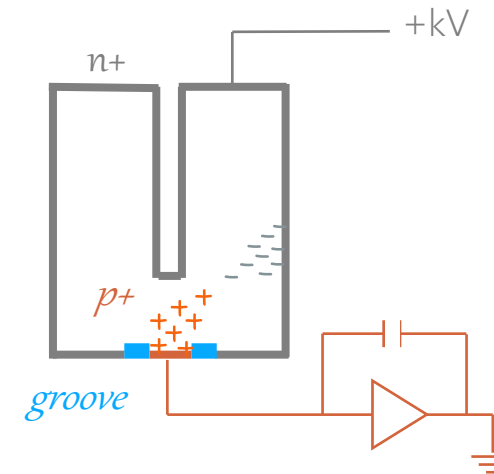
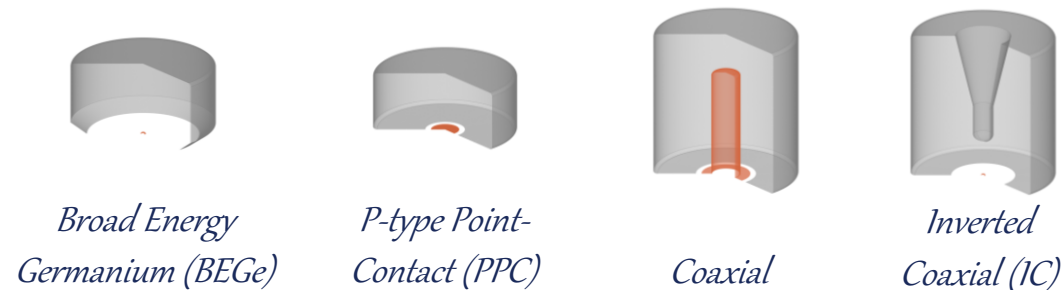
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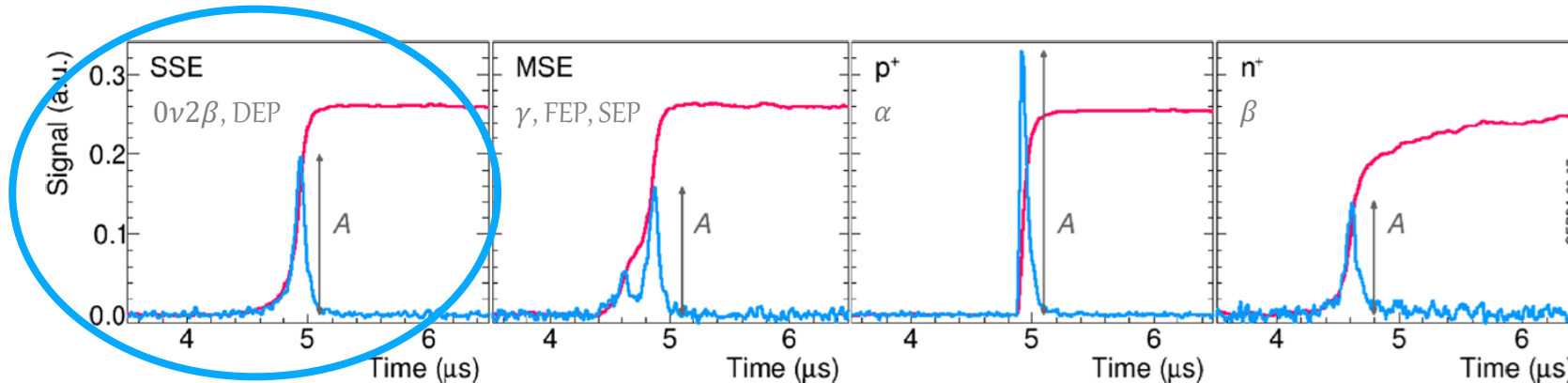
- Signal discrimination based on the signal risetime and amplitude
- **Pulse shape discrimination (PSD) analysis**

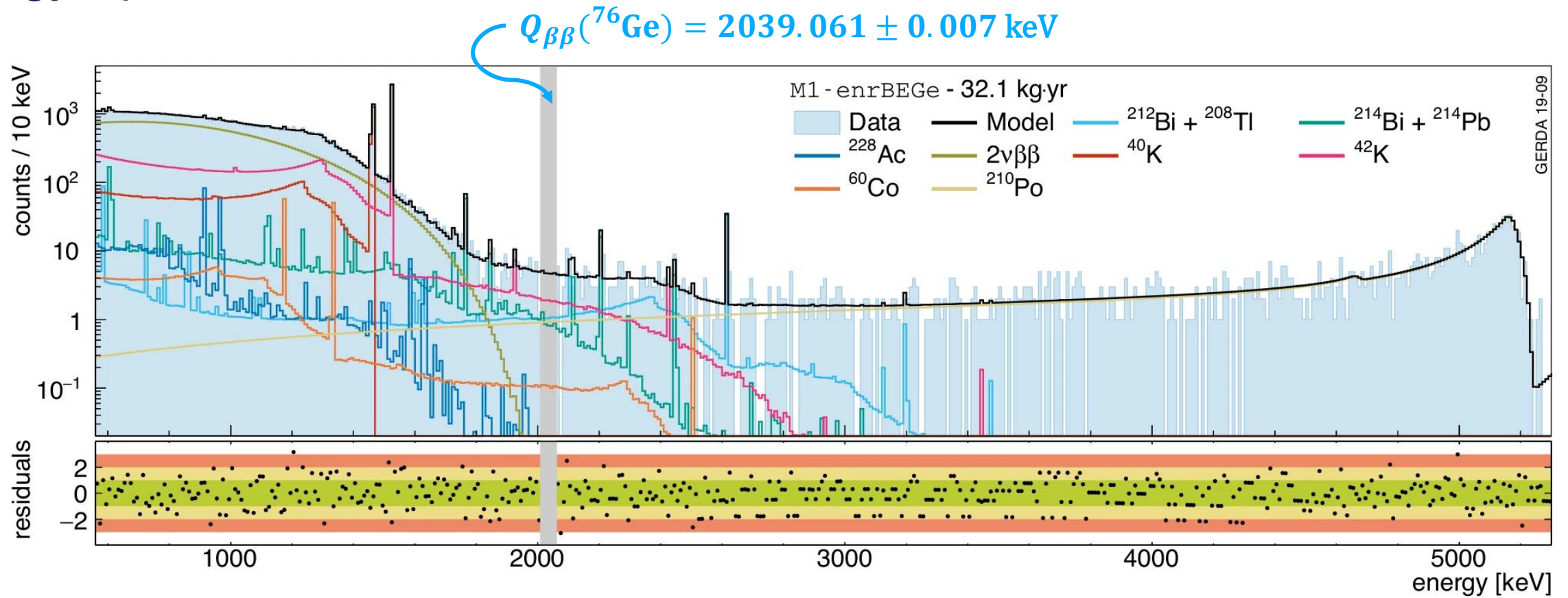


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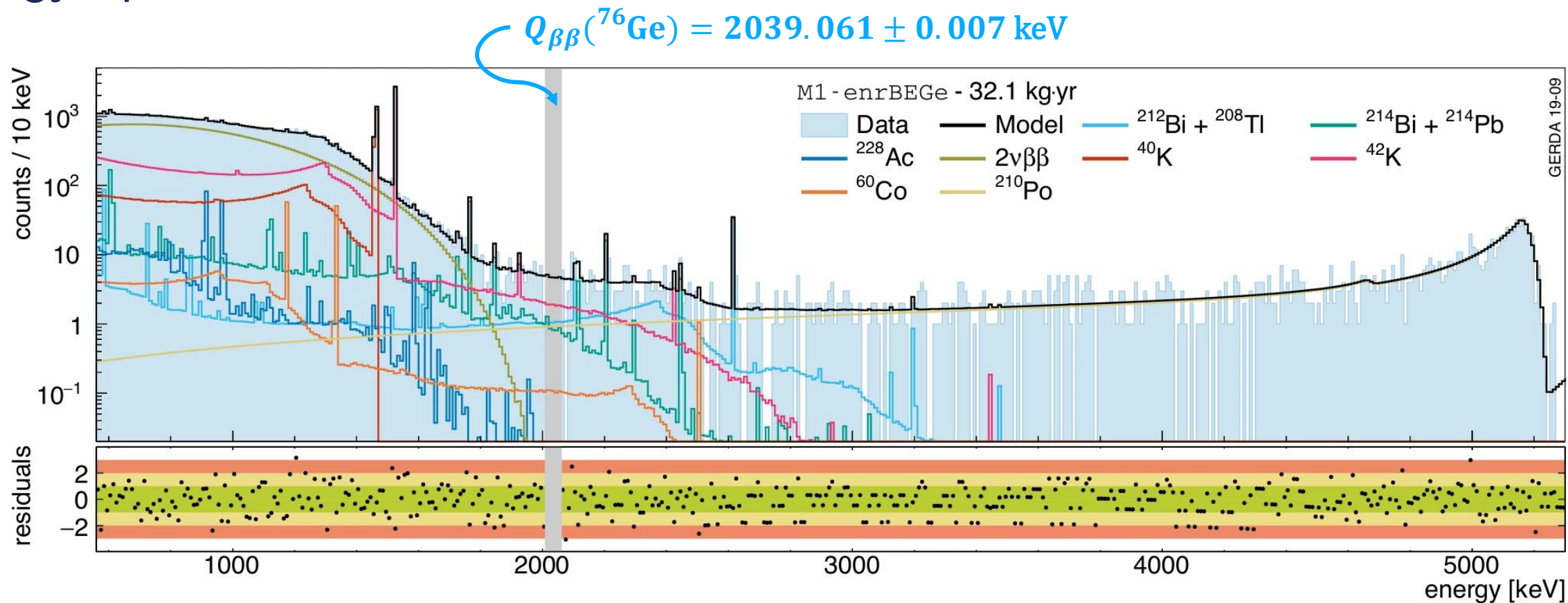


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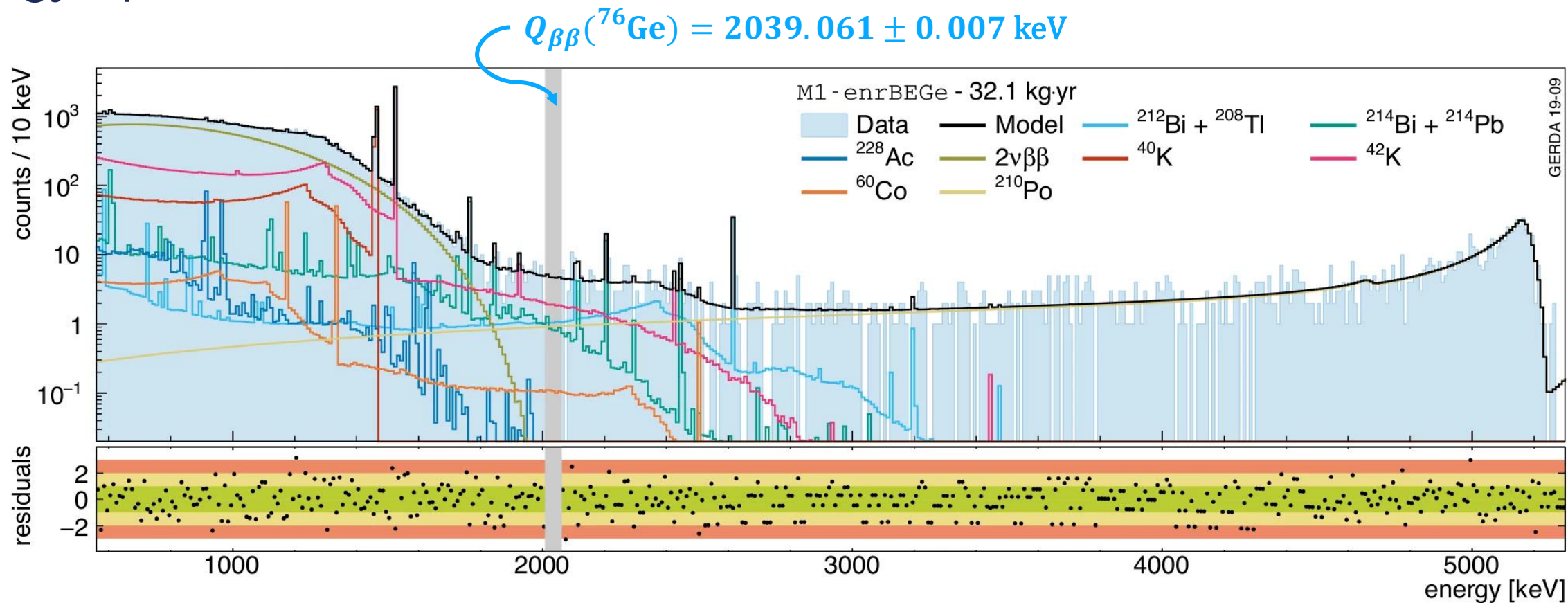




- Energy spectrum prior to the analysis cuts

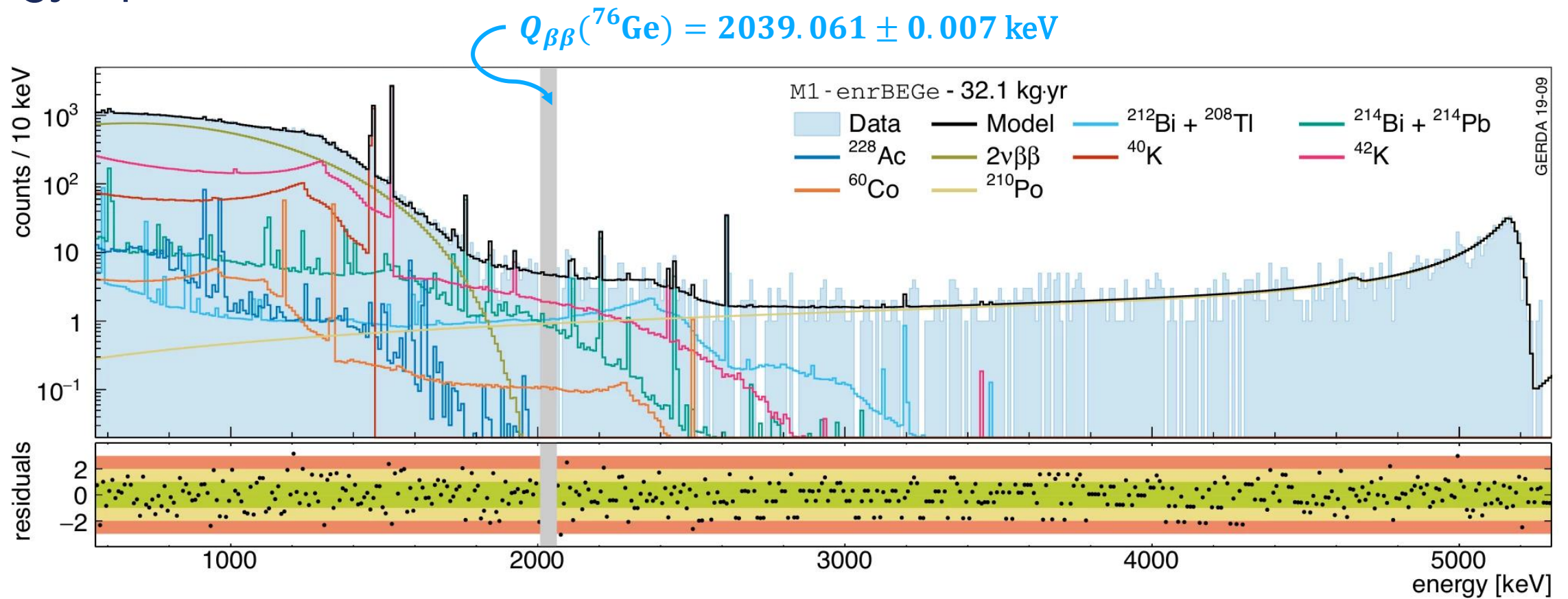


- Energy spectrum prior to the analysis cuts
- Blinded analysis in $Q_{\beta\beta} \pm 25 \text{ keV}$



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- The **background index (BI)** is evaluated starting from the study of the remnant spectrum

$$T_{1/2}^{0\nu} \propto \begin{cases} \epsilon \cdot f \cdot \epsilon & \text{with bkg} \\ \epsilon \cdot f \cdot \sqrt{\frac{\epsilon}{\text{BI} \cdot \Delta E}} & \text{without bkg} \end{cases}$$



- Energy spectrum prior to the analysis cuts
- Blinded analysis in $Q_{\beta\beta} \pm 25 \text{ keV}$
- The **background index (BI)** is evaluated starting from the study of the remnant spectrum
- **"Bkg-free regime":** <1 count in $Q_{\beta\beta} \pm 0.5 \text{ FWHM}$

$$T_{1/2}^{0\nu} \propto \begin{cases} \epsilon \cdot f \cdot \epsilon & \text{without bkg} \\ \epsilon \cdot f \cdot \sqrt{\frac{\epsilon}{BI \cdot \Delta E}} & \text{with bkg} \end{cases}$$



The LEGEND project

GERDA

GERmanium Detector Array

See Yannick Müller's talk (parallel 12)



Bkg-free!
~0.3 events for 100 kg yr

	GERDA
Mass [kg]	45
Exposure [kg·yr]	100
BI [cts/(keV·kg·yr)]	$(5.2 \pm 1.6) \cdot 10^{-4}$
Resolution [keV]	2.6 ± 0.2

GERDA

GERmanium Detector Array

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MAJORANA DEMONSTRATOR



	GERDA	MAJORANA
Mass [kg]	45	30
Exposure [kg·yr]	100	26
BI [cts/(keV·kg·yr)]	$(5.2 \pm 1.6) \cdot 10^{-4}$	$(4.7 \pm 0.8) \cdot 10^{-3}$
Resolution [keV]	2.6 ± 0.2	2.52 ± 0.08

GERDA

GERmanium Detector Array
See Yannick Müller's talk (parallel 12)



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MAJORANA DEMONSTRATOR



LEGEND

Large Enriched Germanium Experiment for
Neutrinoless $\beta\beta$ Decay



200



1000

	GERDA	MAJORANA	LEGEND-200	LEGEND-1000
Mass [kg]	45	30	200	1'000
Exposure [kg·yr]	100	26	1'000	10'000
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GERDA

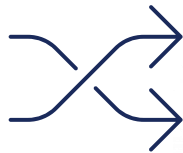
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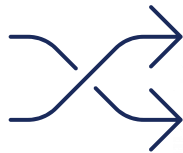
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MAJORANA
DEMONSTRATOR



LEGEND

Large Enriched Germanium Experiment for
Neutrinoless $\beta\beta$ Decay



200



1000

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Mass [kg]	45	30	200	1'000
Exposure [kg·yr]	100	26	1'000 >>> ×10	10'000 >>> ×100
BI [cts/(keV·kg·yr)]	$(5.2 \pm 1.6) \cdot 10^{-4}$	$(4.7 \pm 0.8) \cdot 10^{-3}$	$2 \cdot 10^{-4}$	10^{-5}
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GERDA

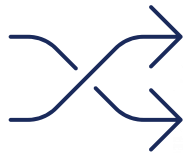
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MAJORANA
DEMONSTRATOR



LEGEND

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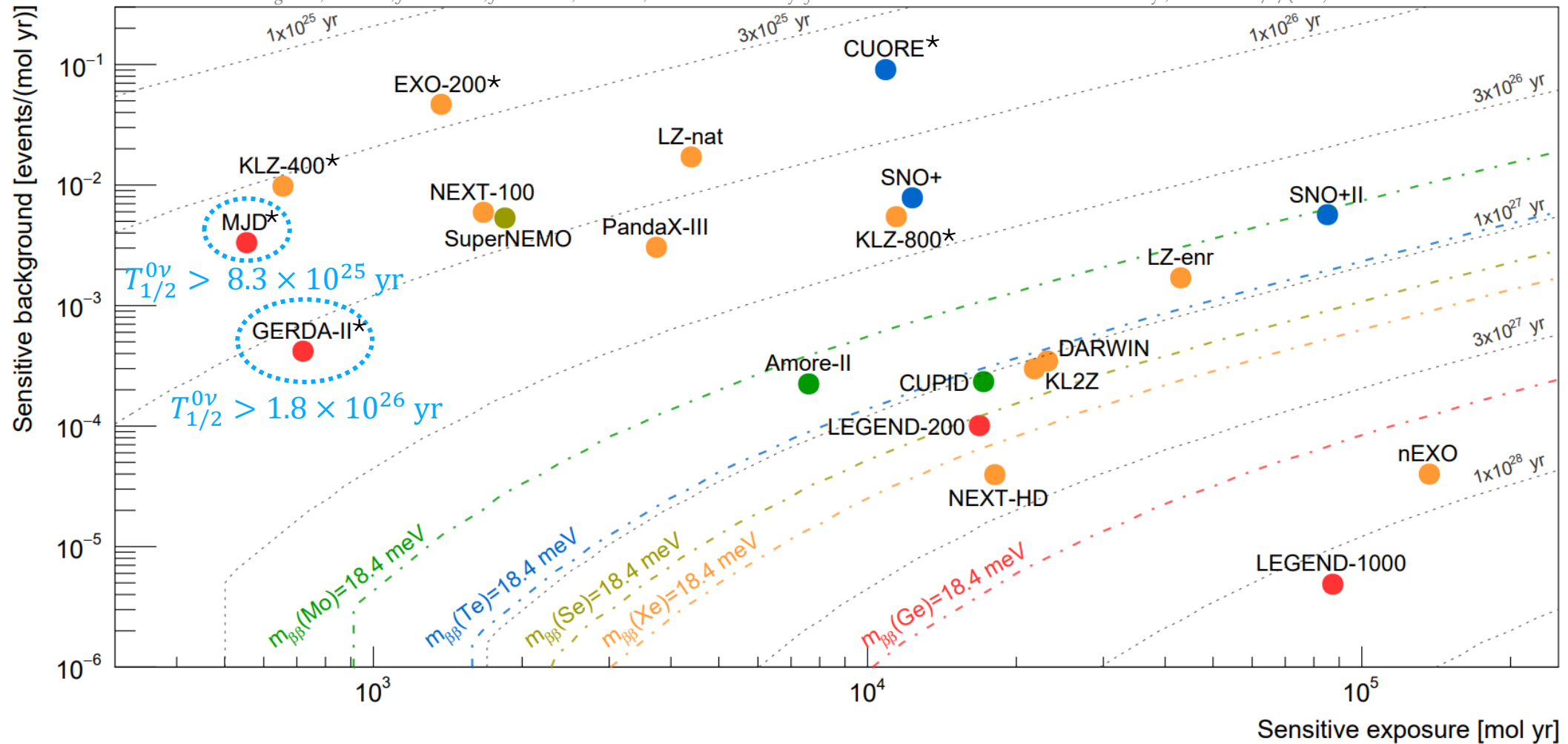
200



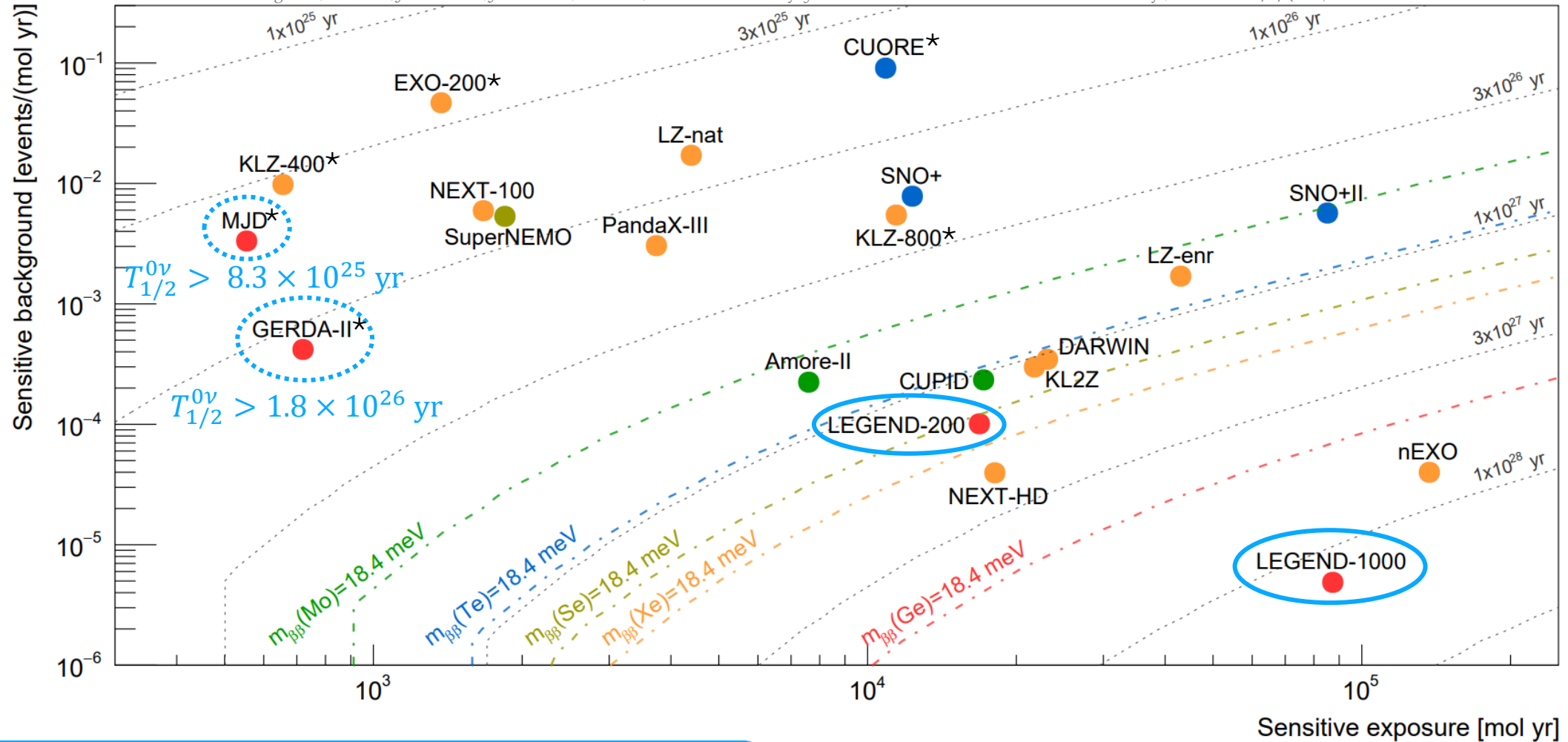
1000

	GERDA	MAJORANA	LEGEND-200	LEGEND-1000
Mass [kg]	45	30	200	1'000
Exposure [kg·yr]	100	26	1'000 >>> $\times 10$	10'000 >>> $\times 100$
BI [cts/(keV·kg·yr)]	$(5.2 \pm 1.6) \cdot 10^{-4}$	$(4.7 \pm 0.8) \cdot 10^{-3}$	$2 \cdot 10^{-4}$ <<< $\times 2.5$	10^{-5} <<< $\times 50$
Resolution [keV]	2.6 ± 0.2	2.52 ± 0.08	2.5	2.5

M. Agostini, G. Benato, J. A. Detwiler, J. Menéndez, F. Vissani, "Toward the discovery of matter creation with neutrinoless double-beta decay", arXiv:2202.01787 (2022).



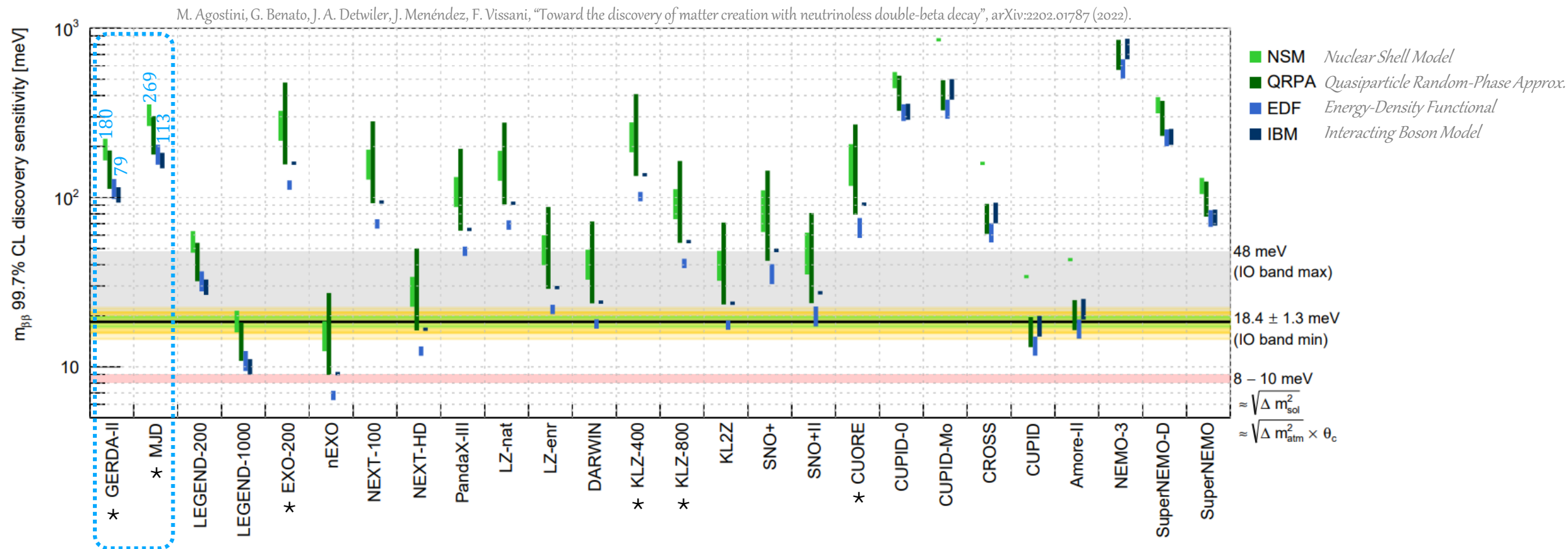
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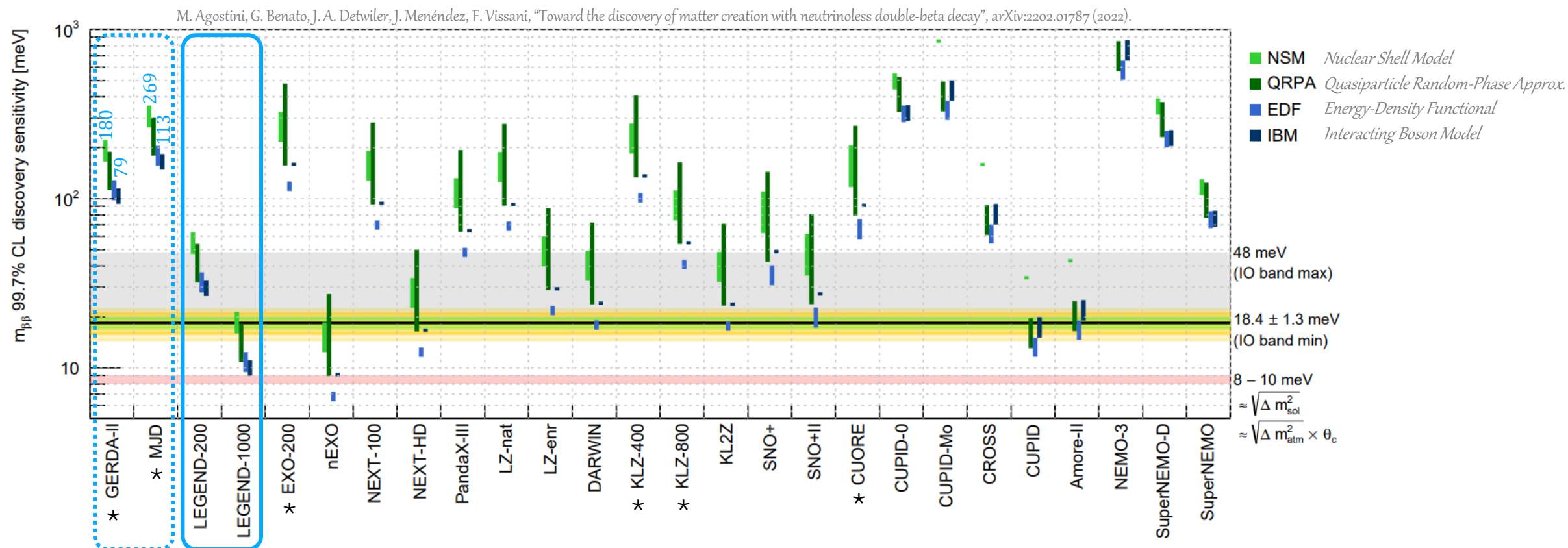
LEGEND-200: $T_{1/2}^{0\nu} > 10^{27}$ yr @ 90% CL

LEGEND-1000: $T_{1/2}^{0\nu} > 1.6 \cdot 10^{28}$ yr @ 90% CL

- Assumption: 10 yr of data acquisition for all experiments, but for those that already finished to operate (*)
- The colorful-dashed lines correspond to the half-life sensibility requested to test the inverted hierarchy with isotopes Mo-100, Te-130, Se-82, Xe-136, Ge-76



- Gray band: range of values $m_{\beta\beta}$ for the inverse hierarchy and $m_{\text{light}} \rightarrow 0$
- $m_{\beta\beta} = 18.4$ meV: minimum allowed value for the IO
- 1σ , 2σ , 3σ uncertainty bands for $m_{\beta\beta} = 18.4$ meV are shown in green, orange and yellow, respectively
- Red band at 8-10 meV: future aim for $0\nu\beta\beta$ next-generation experiments

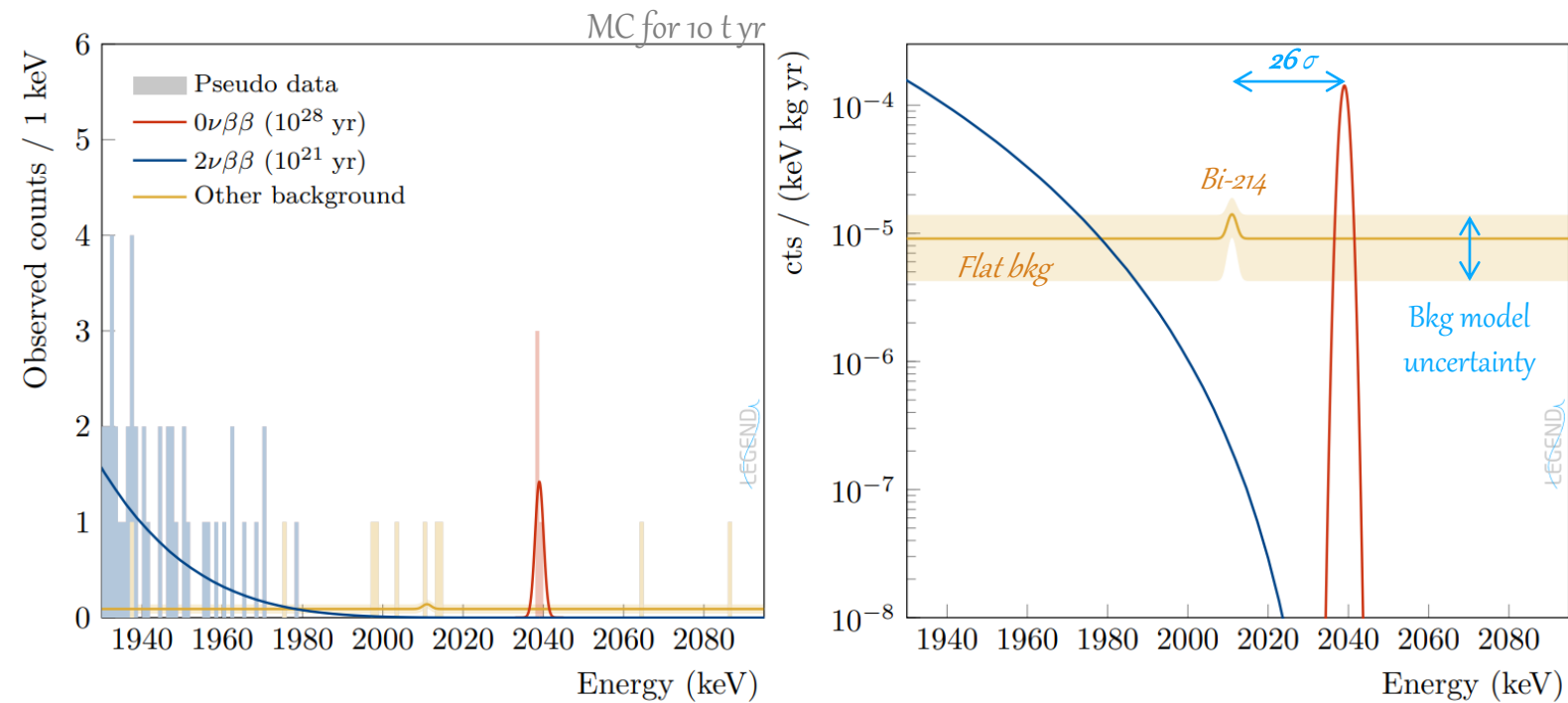


LEGEND-200: $m_{\beta\beta} < 34 - 78$ meV
 LEGEND-1000: $m_{\beta\beta} < 8.5 - 19.4$ meV

- Gray band: range of values $m_{\beta\beta}$ for the inverse hierarchy and $m_{\text{light}} \rightarrow 0$
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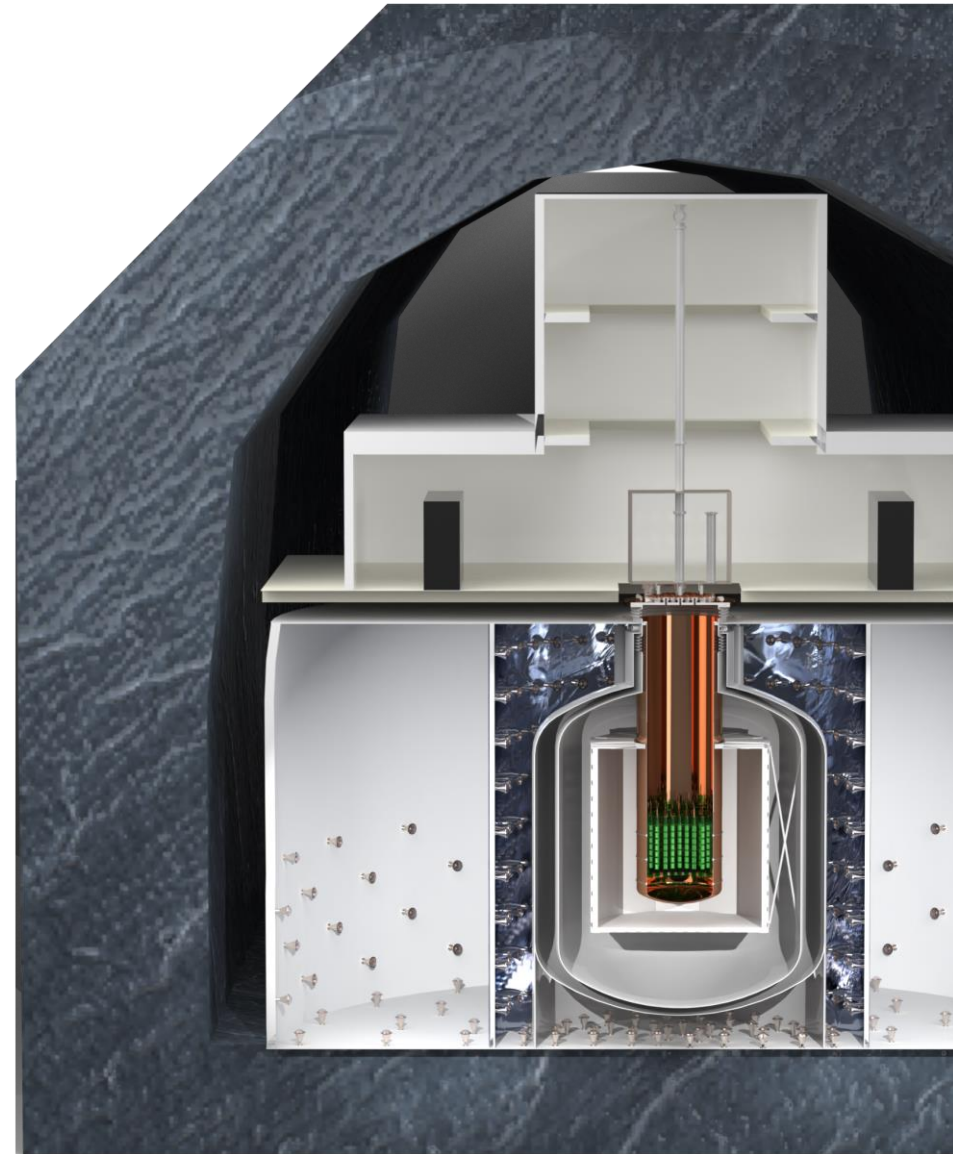


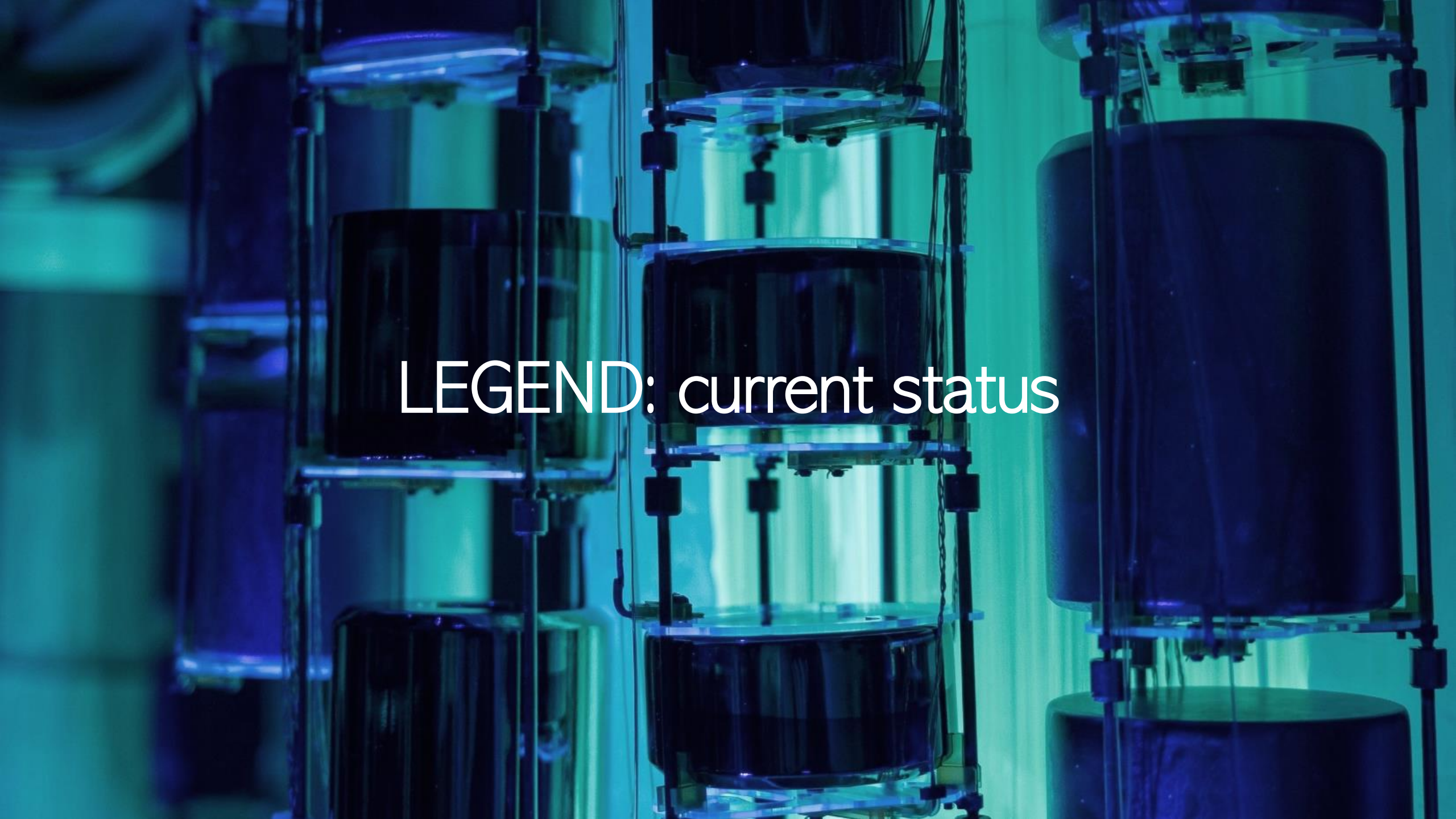
The ultimate goal



“By combining the lowest background levels and the best energy resolution in the field, **LEGEND-1000** will perform a quasi-background-free search and **can make an unambiguous discovery of $0\nu\beta\beta$ decay** with just a handful of counts at the $Q_{\beta\beta}$ ”

Pre-conceptual Design Report: <https://inspirehep.net/literature/1892243>

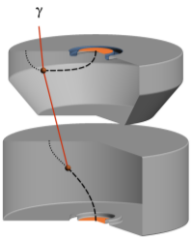




LEGEND: current status



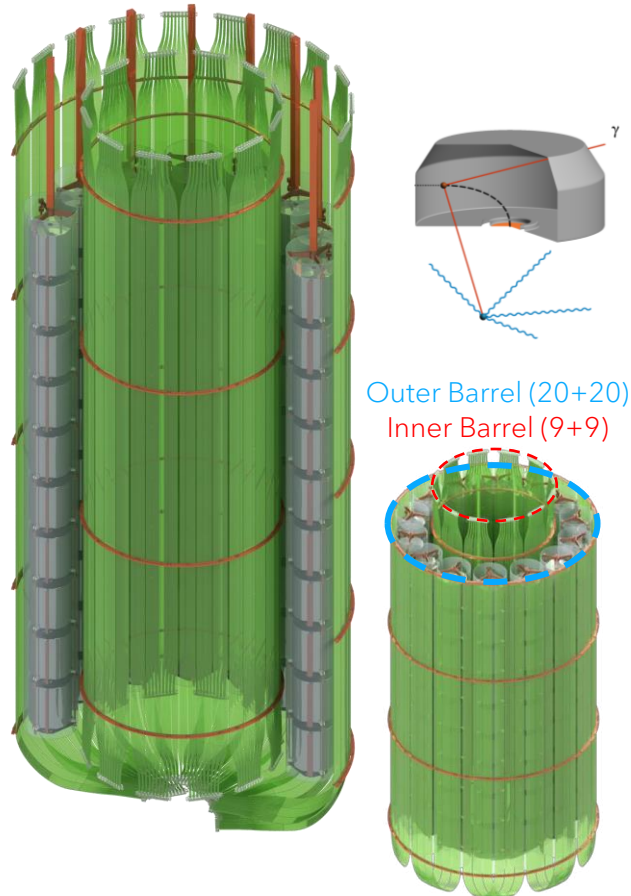
Array
of germanium diodes







- LAr veto
- 58 read-out modules of SiPMs coupled to WLS fibers



- Criostat
- Ø 4m, H 5.88m
 - LAr: 64 m³

- Ultrpure water tank
- Shields n, γ
 - 66 PMTs (Cherenkov) + plastic scintillators for μ
 - Ø 10m, H 8.5 m, V 590 m³
 - Clean room at 9.7 m

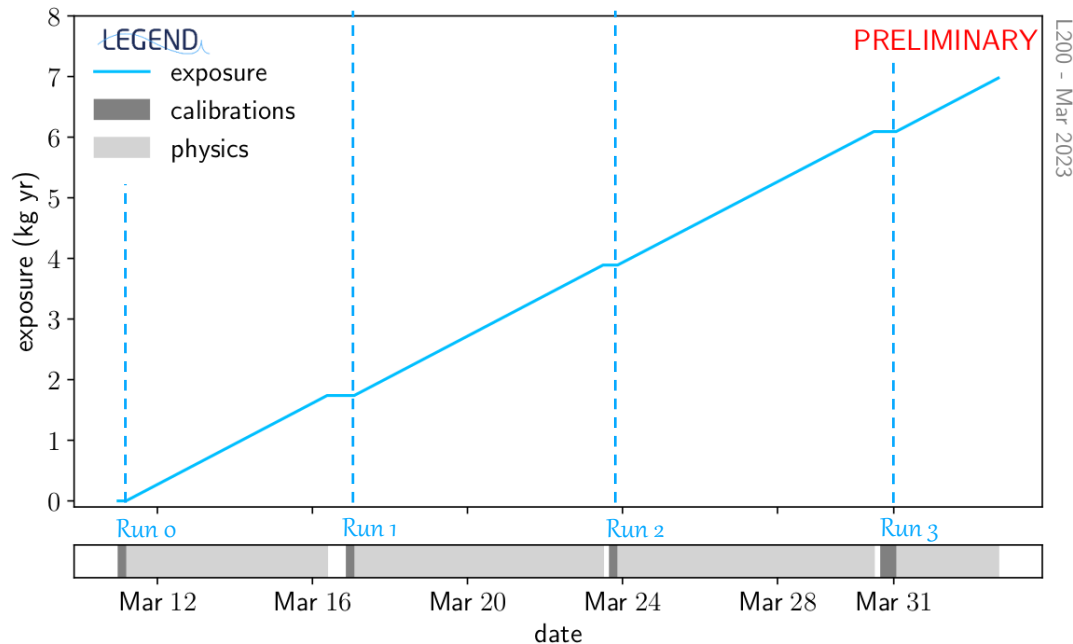




- LNGS-Hall A @ 3600 m w.e.
- Same infrastructure of GERDA
- Muon flux reduced to $1.25 / (\text{m}^2\text{h})$ - factor $\mathcal{O}(10^6)$

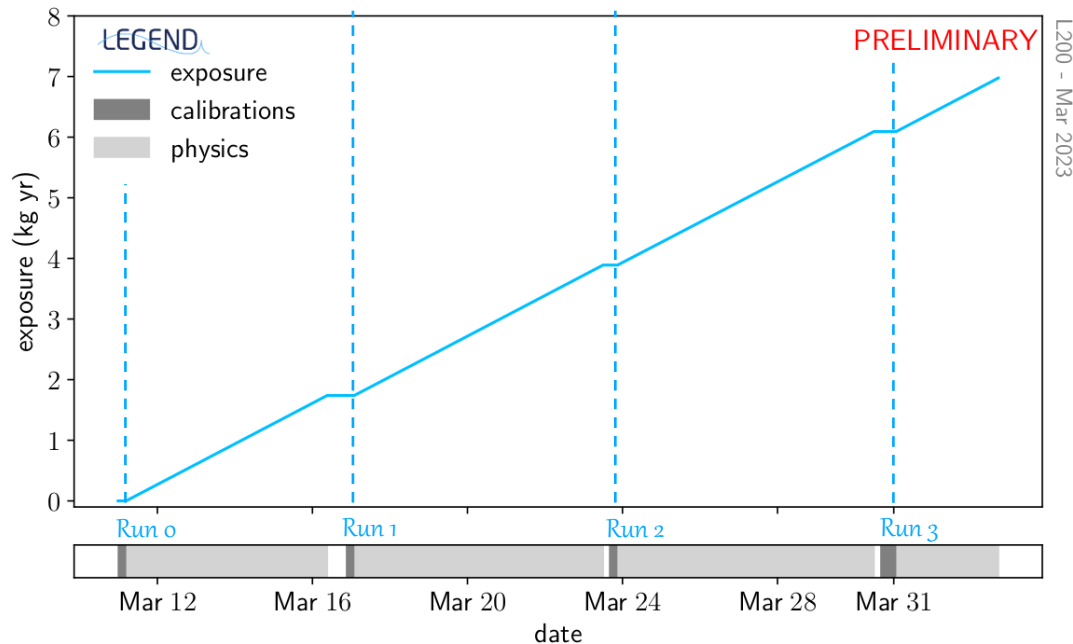
- 101 detectors (10 strings): 6 coaxials, 26 PPC, 28 BEGe, 41 ICPCs
- $m_{\text{tot}}=142.4 \text{ kg}$ (122 kg for the analysis)

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- $m_{\text{tot}} = 142.4 \text{ kg}$ (122 kg for the analysis)
- Stable data taking started **12 March 2023**
- Excluded unstable detectors (pulser, $R=0.05 \text{ Hz}$)

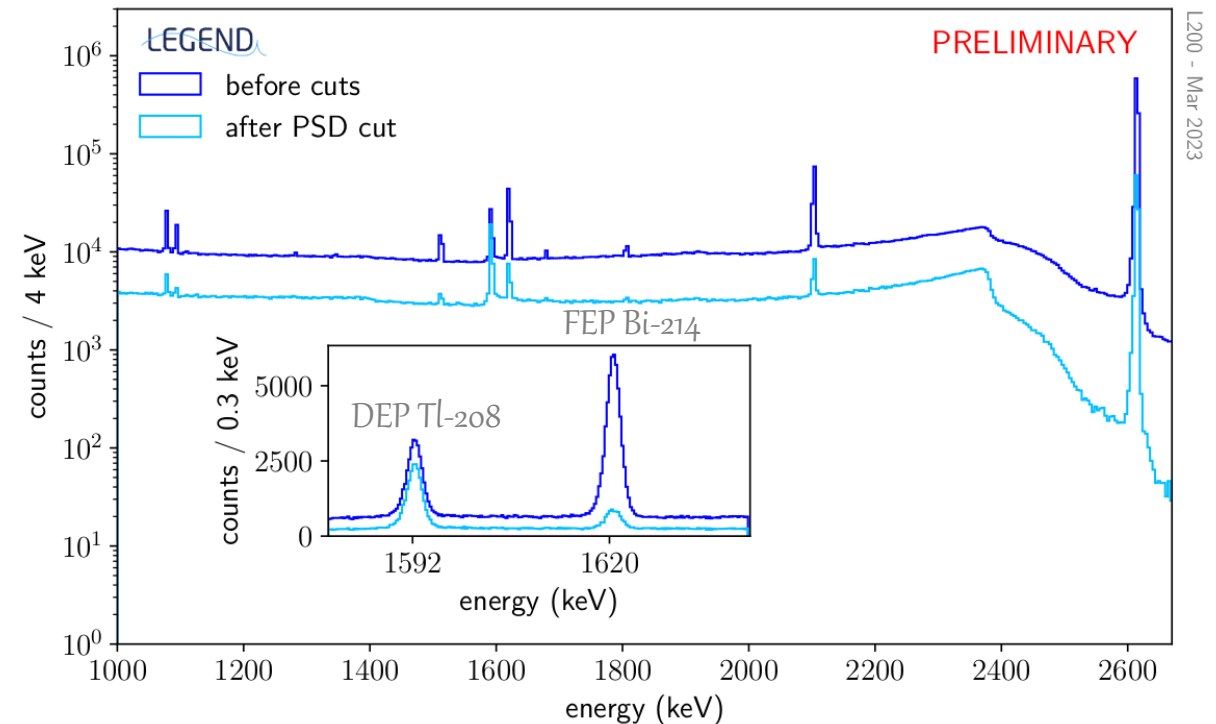


Exposure: $\sim 7 \text{ kg yr}$ ($> 2 \text{ kg yr / week}$)

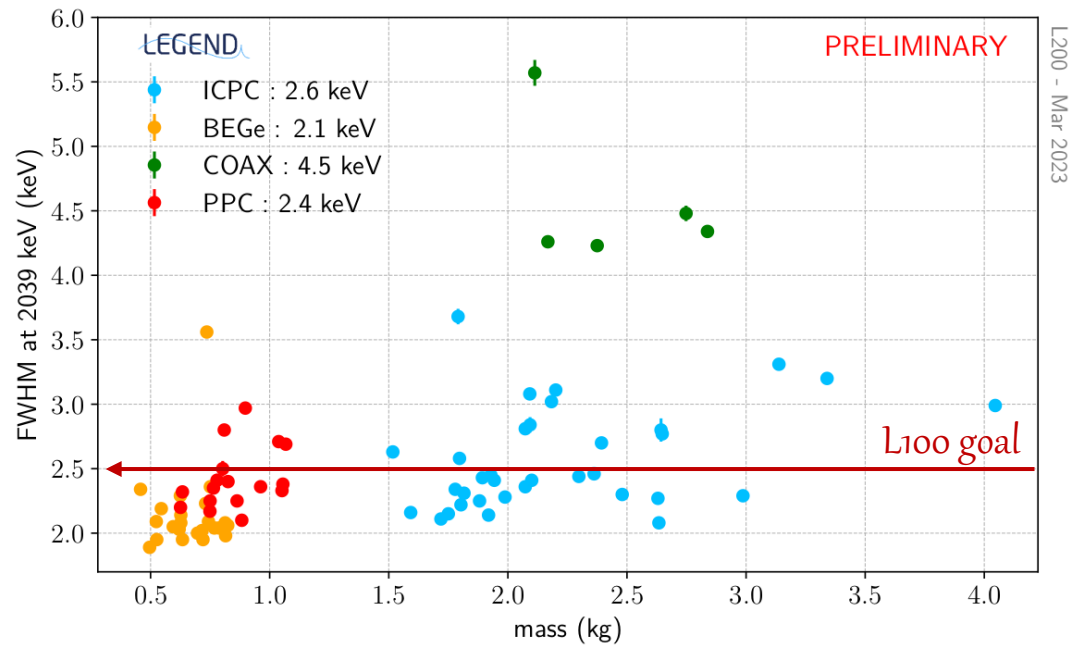
- 101 detectors (10 strings): 6 coaxials, 26 PPC, 28 BEGe, 41 ICPCs
- $m_{\text{tot}} = 142.4 \text{ kg}$ (122 kg for the analysis)
- Stable data taking started **12 March 2023**
- Excluded unstable detectors (pulser, $R=0.05 \text{ Hz}$)



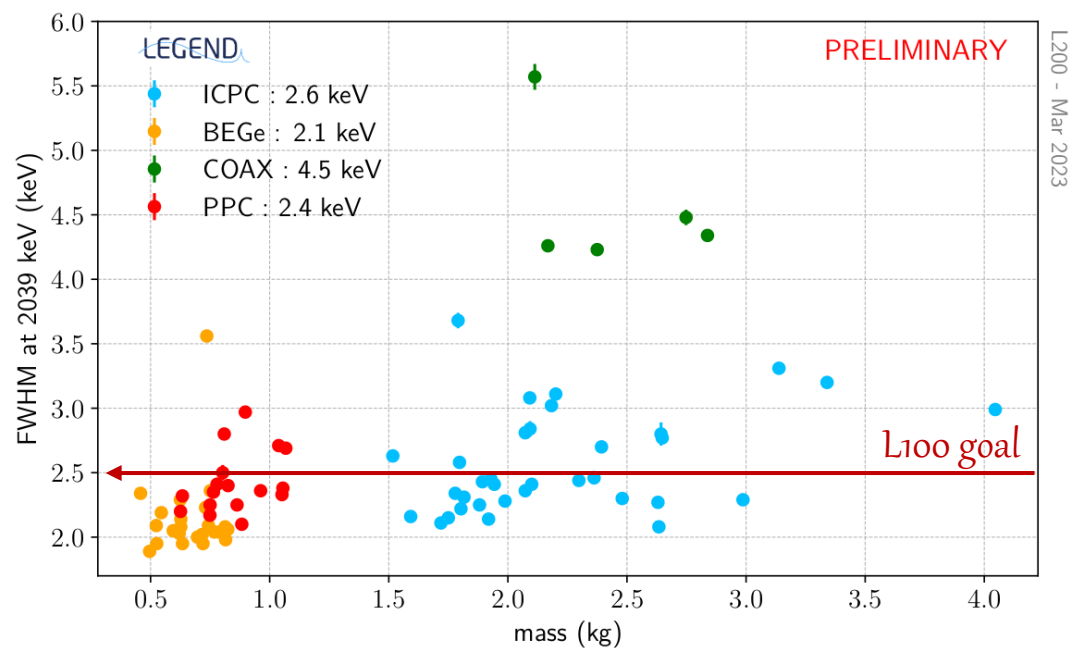
Exposure: $\sim 7 \text{ kg yr}$ ($> 2 \text{ kg yr / week}$)



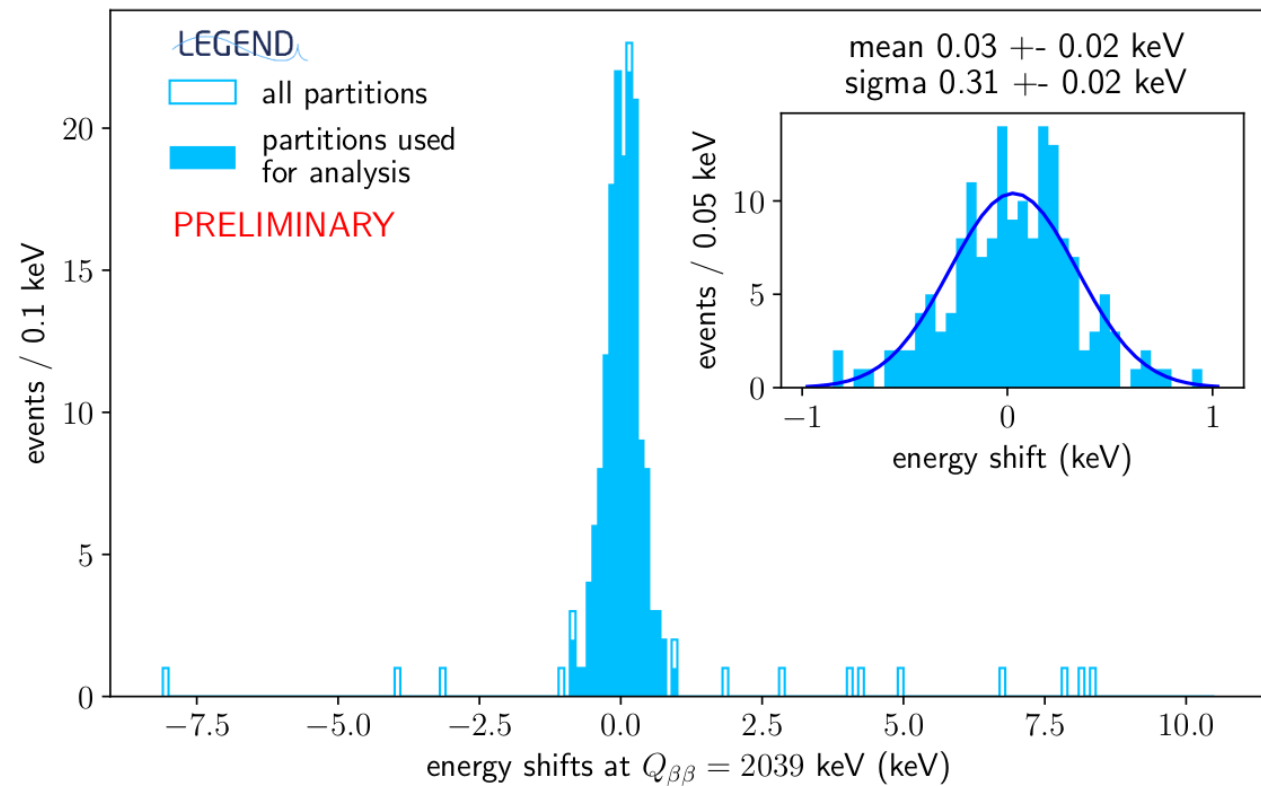
PSD tested with 13 sources of Th-228:
multi-site events (FEP Bi-214): $\sim 10\%$ s.f.
single-site events (DEP Tl-208): $\sim 90\%$ s.f.



Energy resolution:
FWHM ~ 2.4 keV @ $Q_{\beta\beta}$
(4.5 keV for coaxials)

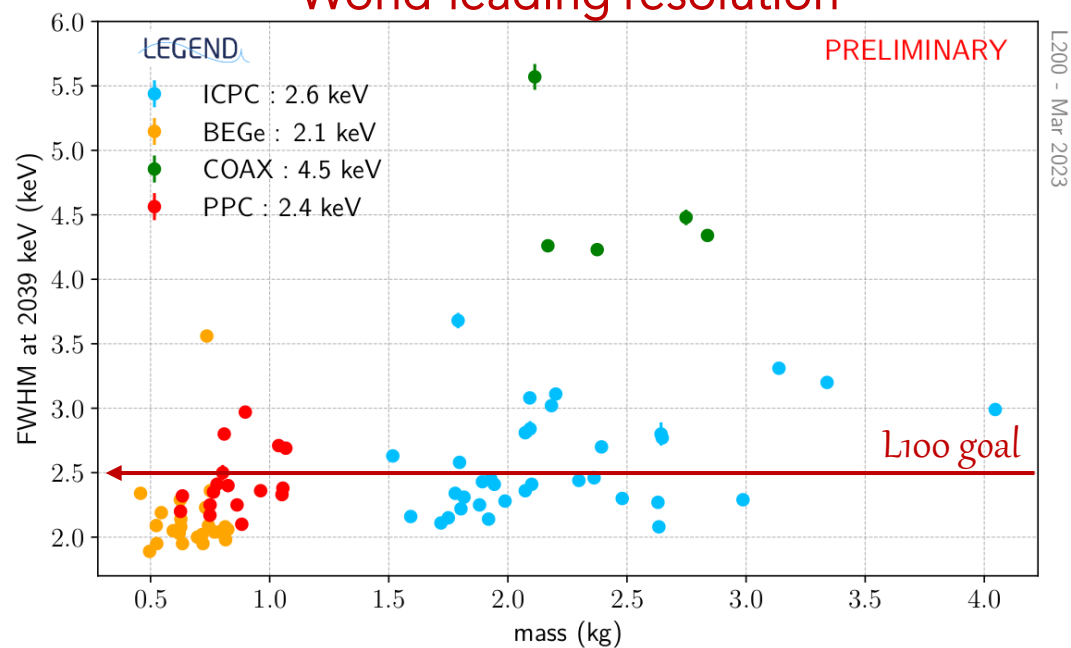


Energy resolution:
FWHM ~ 2.4 keV @ $Q_{\beta\beta}$
(4.5 keV for coaxials)



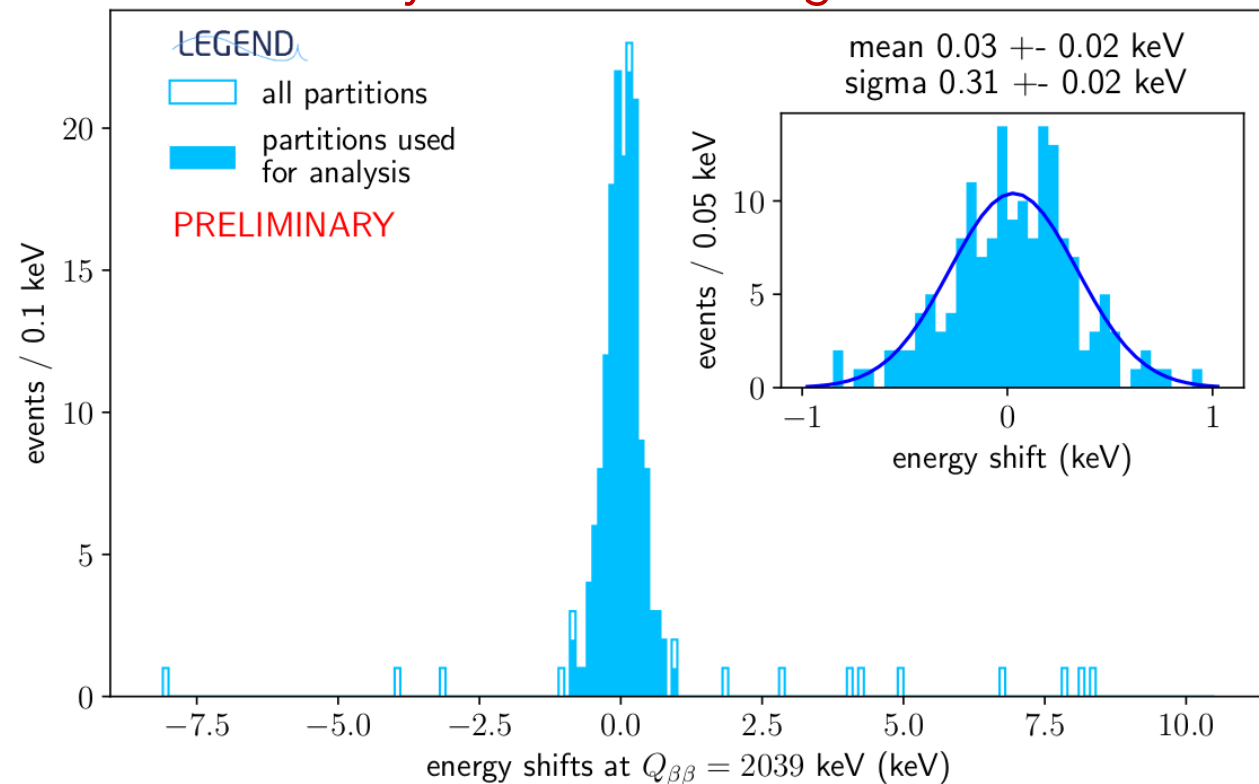
Energy shifts
between subsequent
calibrations

World-leading resolution

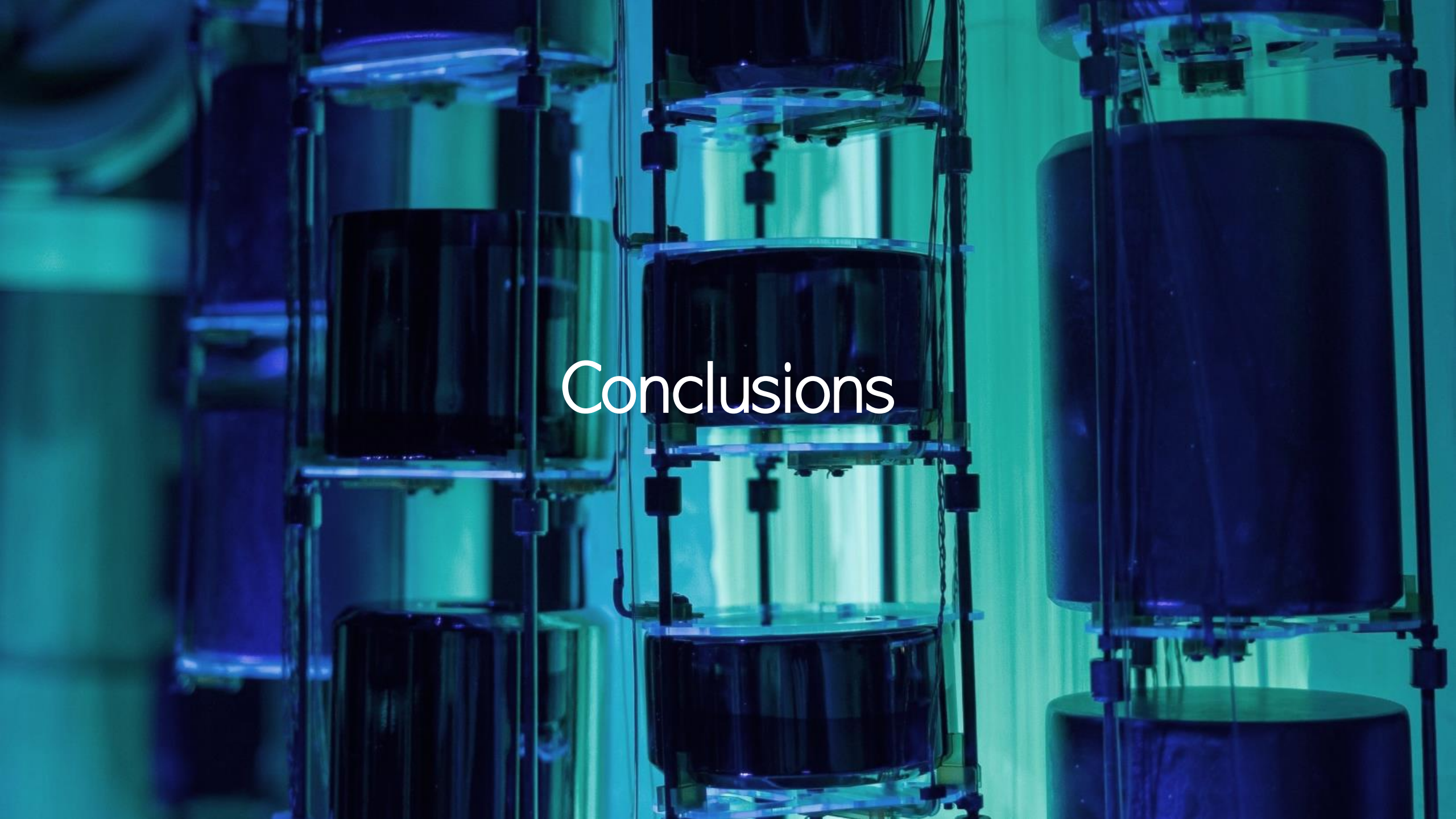


Energy resolution:
FWHM ~ 2.4 keV @ $Q_{\beta\beta}$
(4.5 keV for coaxials)

Extremely stable for >120 kg of detectors



Energy shifts
between subsequent
calibrations



Conclusions

LEGEND = Large Enriched Germanium Experiment for Neutrinoless $\beta\beta$ Decay

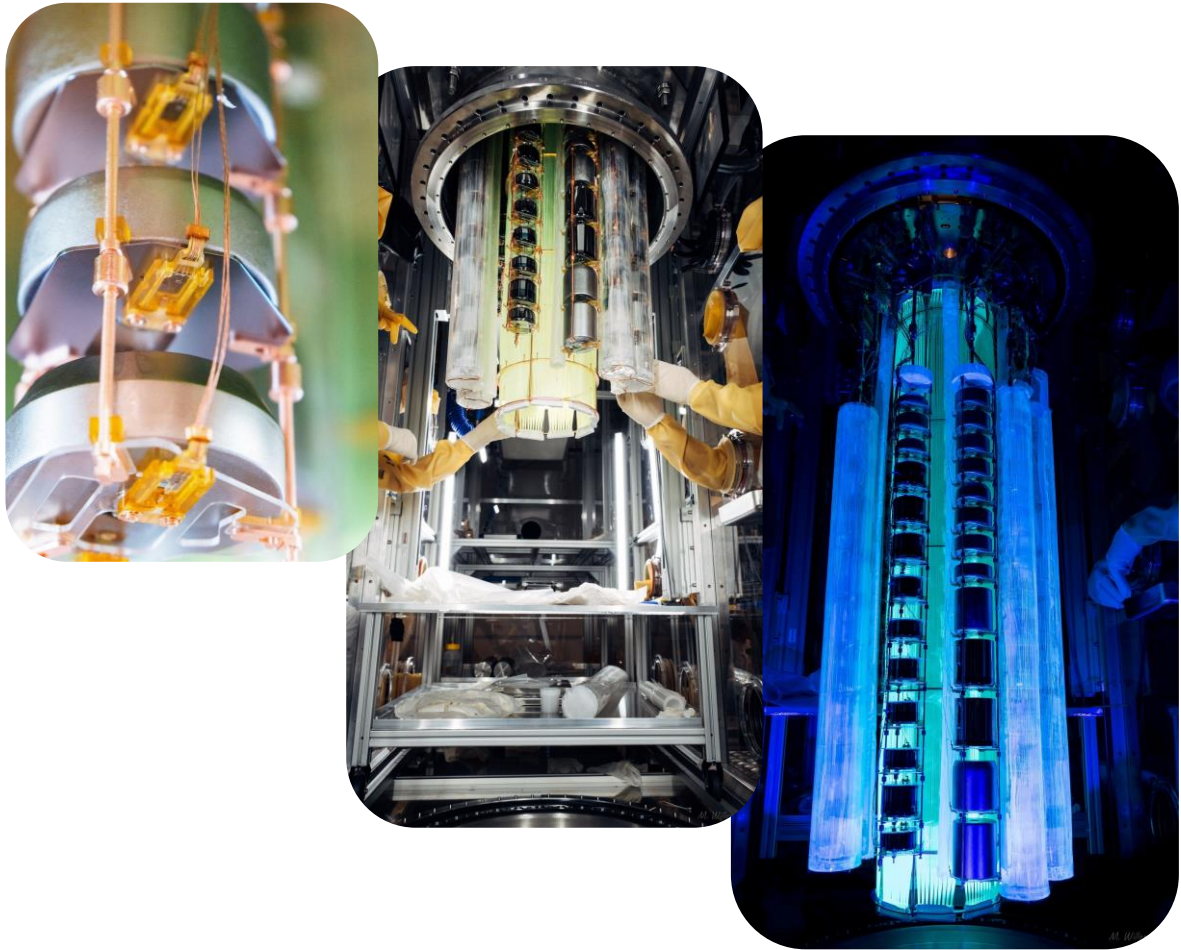
- Neutrinoless Double Beta ($0\nu2\beta$) Decay
 - Requires the construction of large experiments
- Germanium Experiments
 - GERDA and MJD demonstrated the advantages of using the germanium detector technology
- LEGEND
 - Optimized for the search for the $0\nu2\beta$ decay
 - ICPC detectors already satisfy L1000 requirements
 - PSD cuts consisten with expectations
 - ~50 kg of detectors in hand/production
 - BI in late summer, first $0\nu2\beta$ decay results in 2024



LEGEND = Large Enriched Germanium Experiment for Neutrinoless $\beta\beta$ Decay

Physics	Signature	Energy Range
Bosonic dark matter	Peak at DM mass	$< 1 \text{ MeV}$
Electron decay	Peak at 11.8 keV	$\sim 10 \text{ keV}$
Pauli exclusion principle violation	Peak at 10.6 keV	$\sim 10 \text{ keV}$
Solar axions	Peaked spectra, daily modulation	$2 - 10 \text{ keV}$
Majoron emission	$2\nu\beta\beta$ spectral distortion	$< Q_{\beta\beta}$
Exotic fermions	$2\nu\beta\beta$ spectral distortion	$< Q_{\beta\beta}$
Lorentz violation	$2\nu\beta\beta$ spectral distortion	$< Q_{\beta\beta}$
Exotic currents in $2\nu\beta\beta$ decay	$2\nu\beta\beta$ spectral distortion	$< Q_{\beta\beta}$
Time-dependent $2\nu\beta\beta$ decay rate	Modulation of $2\nu\beta\beta$ spectrum	$< Q_{\beta\beta}$
WIMP and related searches	Exponential excess, annual modulation	$< 10 \text{ keV}$
Baryon decay	Timing coincidence	$> 10 \text{ MeV}$
Fractionally charged cosmic-rays	Straight tracks	few keV
Fermionic dark matter	Nuclear recoil/deexcitation	$< \text{few MeV}$
Inelastic boosted dark matter	Positron production	$< \text{few MeV}$
BSM physics in Ar	Features in Ar veto spectrum	ECEC in ^{36}Ar

...and many more!
L1000 BSM paper on the way

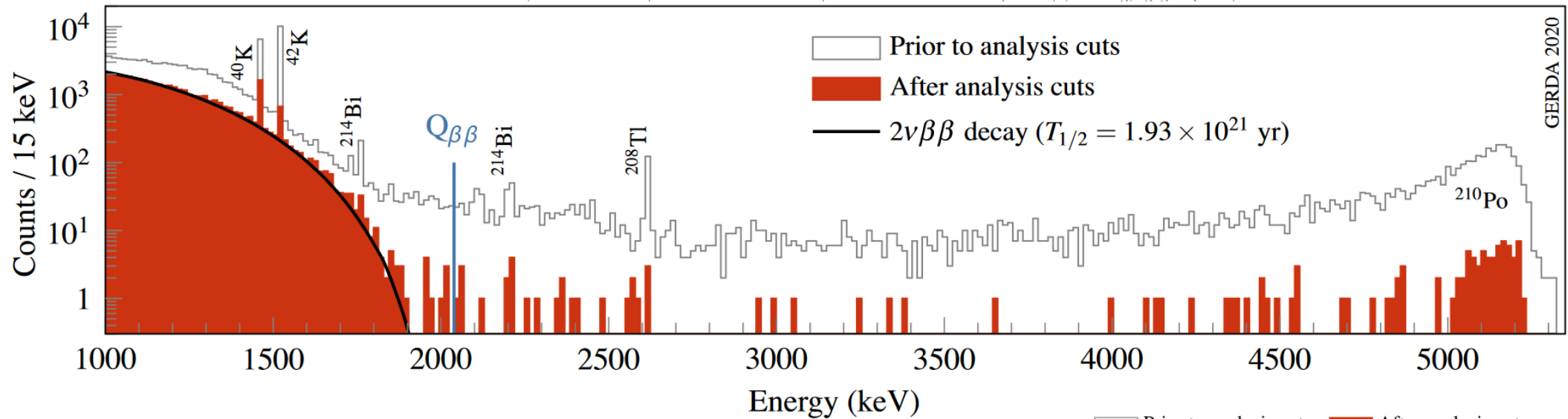




THANK YOU FOR YOUR ATTENTION



Backup slides



- November 2011 – September 2013 (Phase I)
- December 2015 – November 2019 (Phase II)
- Analysis performed over k partitions:

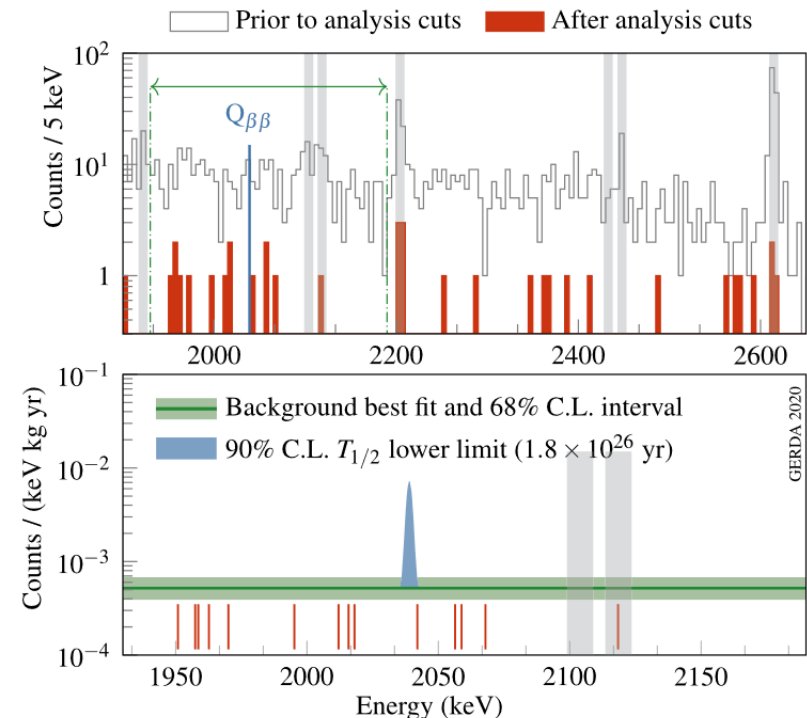
$$\mathcal{L} = \prod_k \left[\frac{(\mu_{s,k} + \mu_{b,k})^{N_k} e^{-(\mu_{s,k} + \mu_{b,k})}}{N_k!} \times \prod_{i=1}^{N_k} \frac{1}{\mu_{s,k} + \mu_{b,k}} \times \left(\frac{\mu_{b,k}}{\Delta E} + \frac{\mu_{s,k}}{\sqrt{2\pi}\sigma_k} e^{-\frac{(E_i - Q_{\beta\beta})^2}{2\sigma_k^2}} \right) \right]$$

- Best limit (and sensitivity assuming no signal) in the world:

$$T_{1/2}^{0\nu} > 1.8 \cdot 10^{26} \text{ yr @ 90\% CL}$$

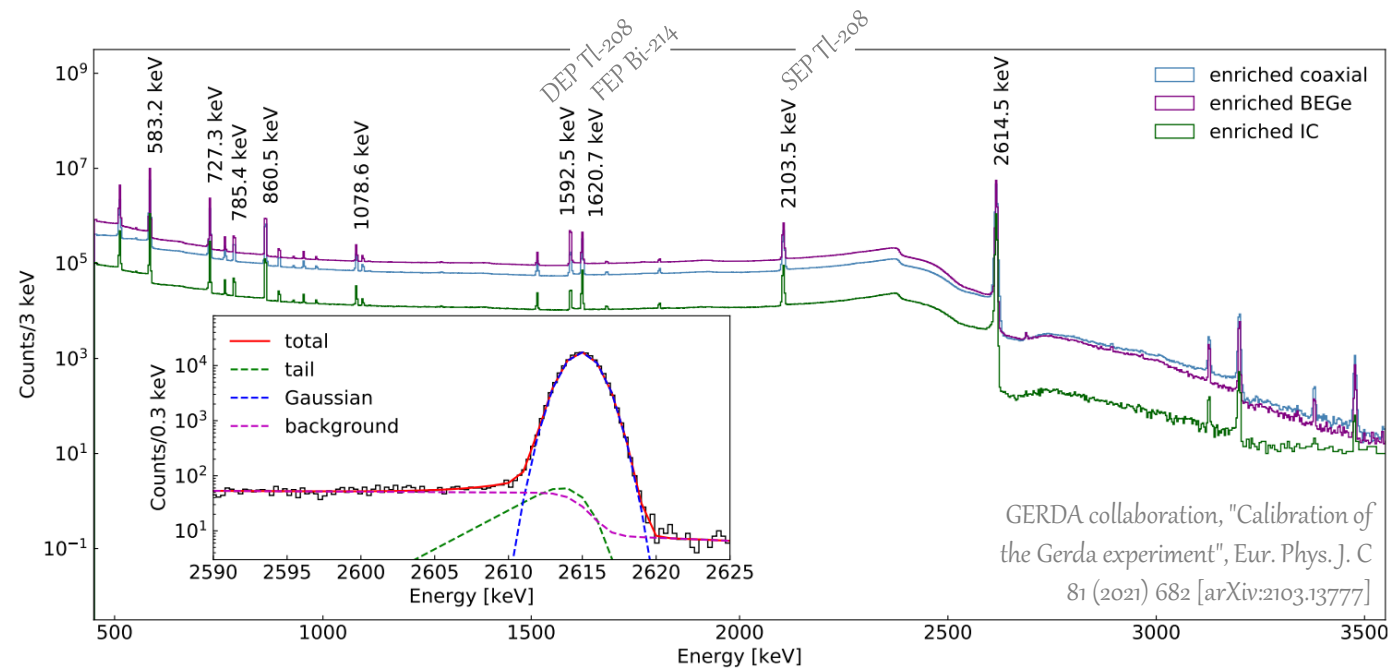
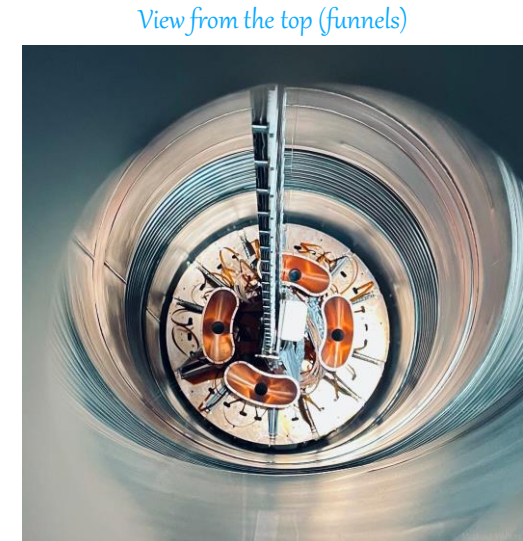
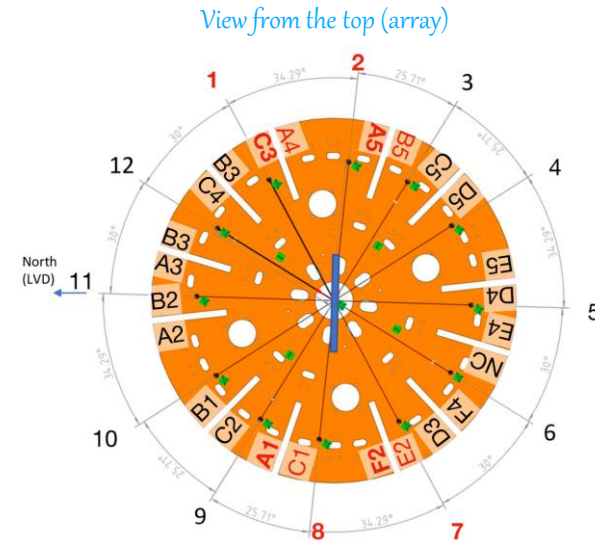
- Best BI in the *Region Of Interest* (ROI):

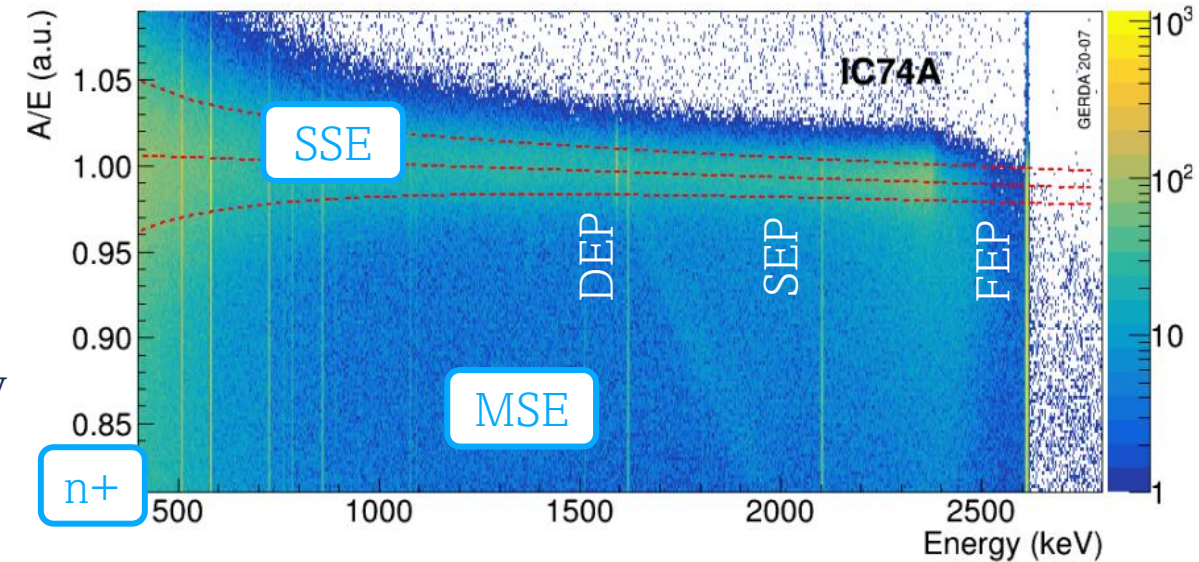
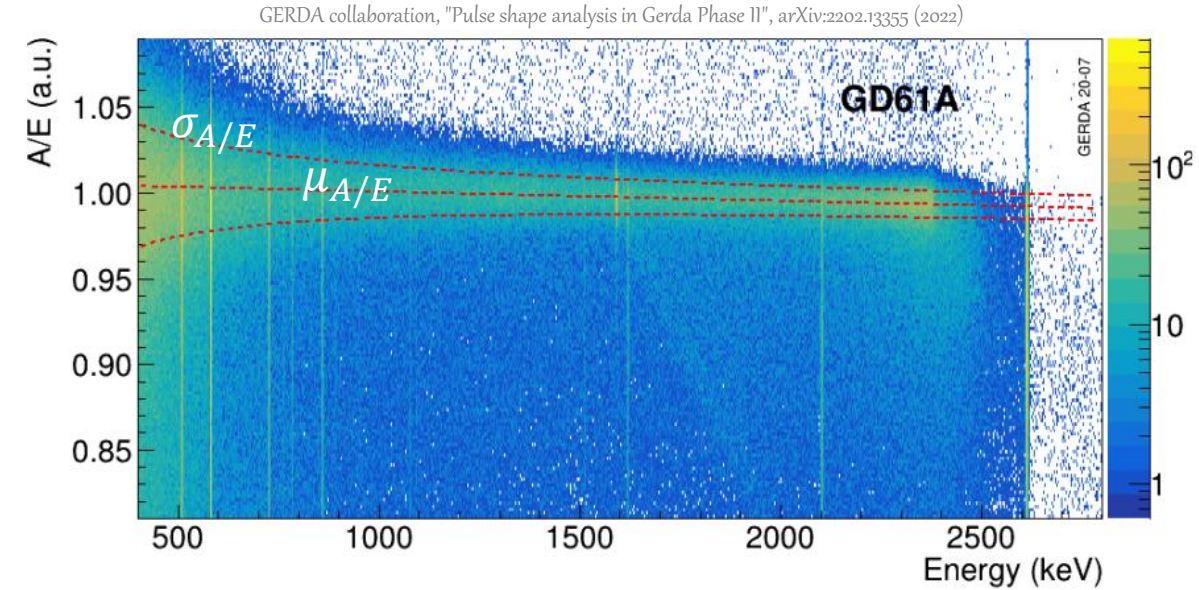
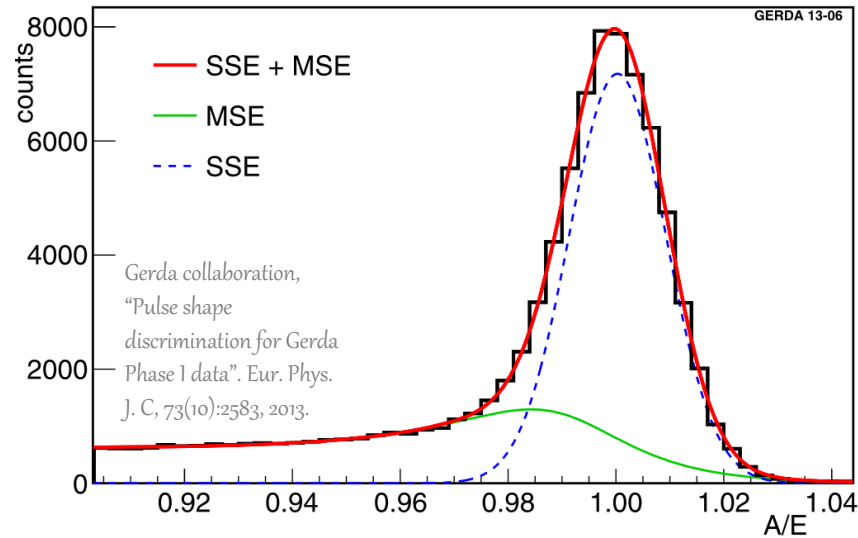
$$BI = 5.2_{-1.3}^{+1.6} \cdot 10^{-4} \text{ cts}/(\text{keV} \cdot \text{kg} \cdot \text{yr})$$



Calibration Procedure

- 4 (+1) *Source Insertion Systems* (SIS), 16 Th-228 sources
- Use γ lines of known energies to pass from ADC to keV (physical unit)
- Peak fitting algorithm to extract each detector's resolution
- 3 terms:
 - Signal = gaussian
 - Bkg #1 = linear + step (flat bkg below/above the peak derived from multiple Compton scatterings)
 - Bkg #2 = tail at low energies (due to incomplete charge collection effects + pile-up events residuals)





- A/E: gaussian (SSE) + tail at low energies (MSE):

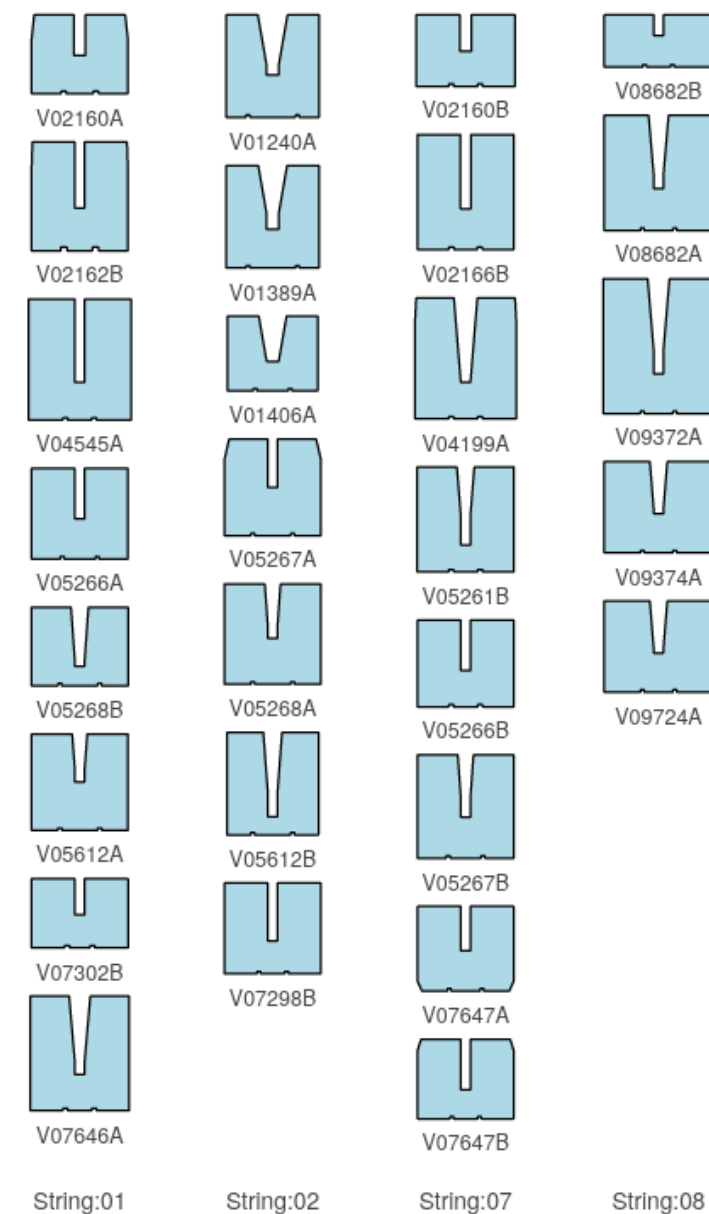
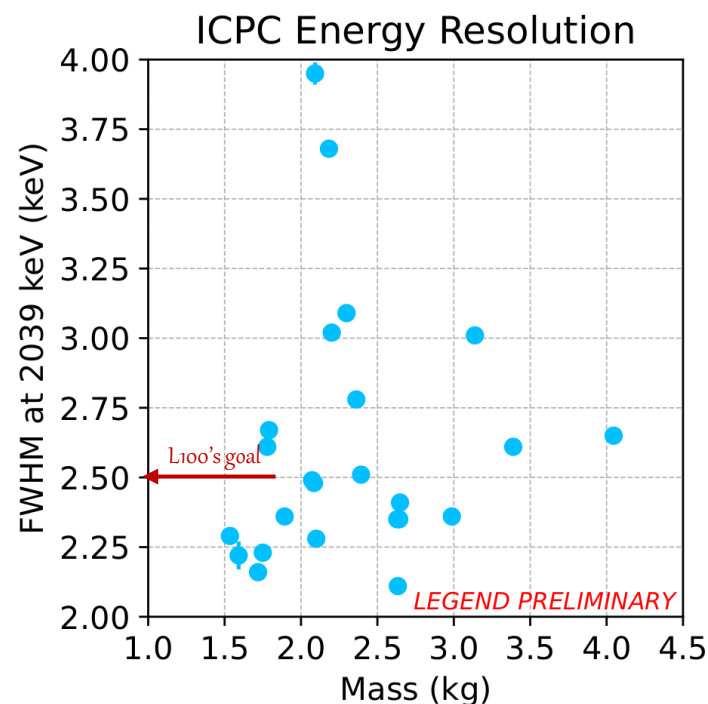
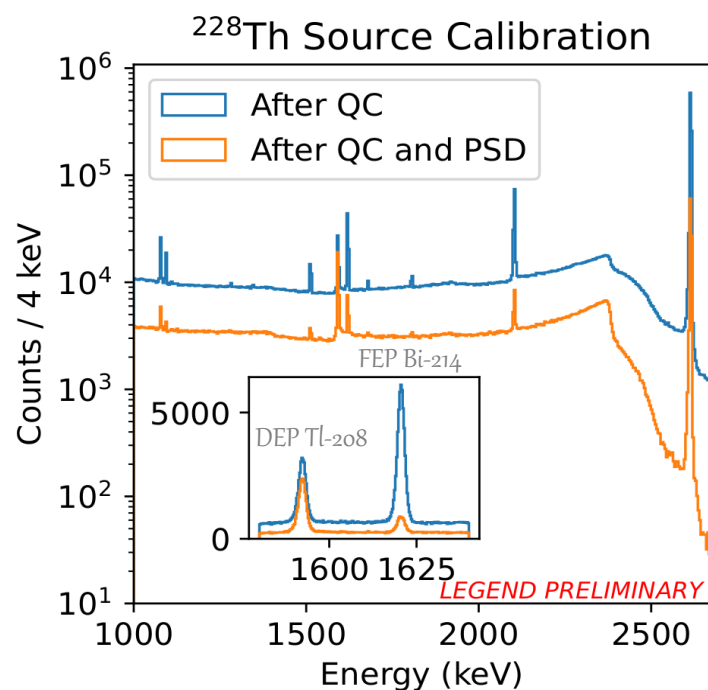
$$f(x = A/E) = \frac{n}{\sigma_{A/E} \cdot \sqrt{2\pi}} \cdot e^{-\frac{(x - \mu_{A/E})^2}{2\sigma_{A/E}^2}} + m \cdot \frac{e^{f \cdot (x-l)} + d}{e^{(x-l)/t} + l}$$

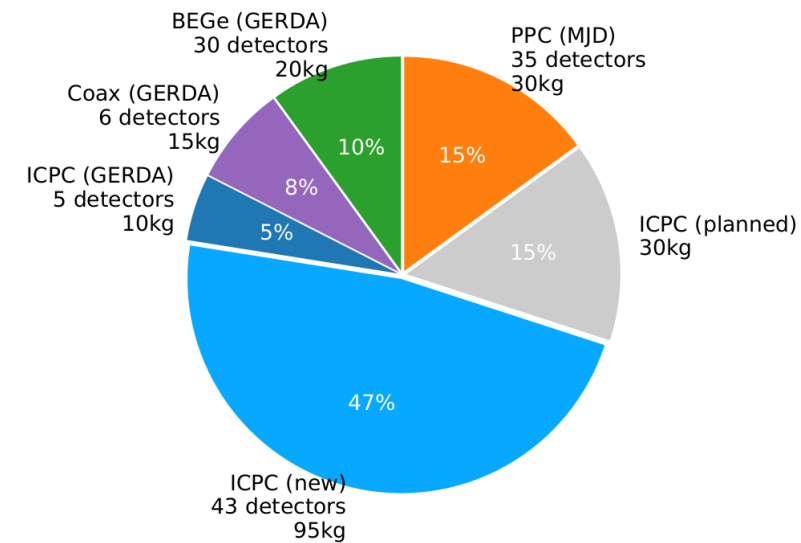
- Cuts over A/E are performed for each detector separately
- “A/E classifier”: energy independent,

$$\zeta = ([A/E] / \mu_{A/E}(E) - 1) / \sigma_{A/E}$$
- ζ distributed around 0, with std=1 for SSEs

28 ICPC detectors, 4 strings

- $m_{\text{tot}} = 63.905 \text{ kg}$
- Data taking period: June - December 2022
- Excluded unstable detectors (pulser, $R = 0.05 \text{ Hz}$)
- Exposure: 2.248 kg yr



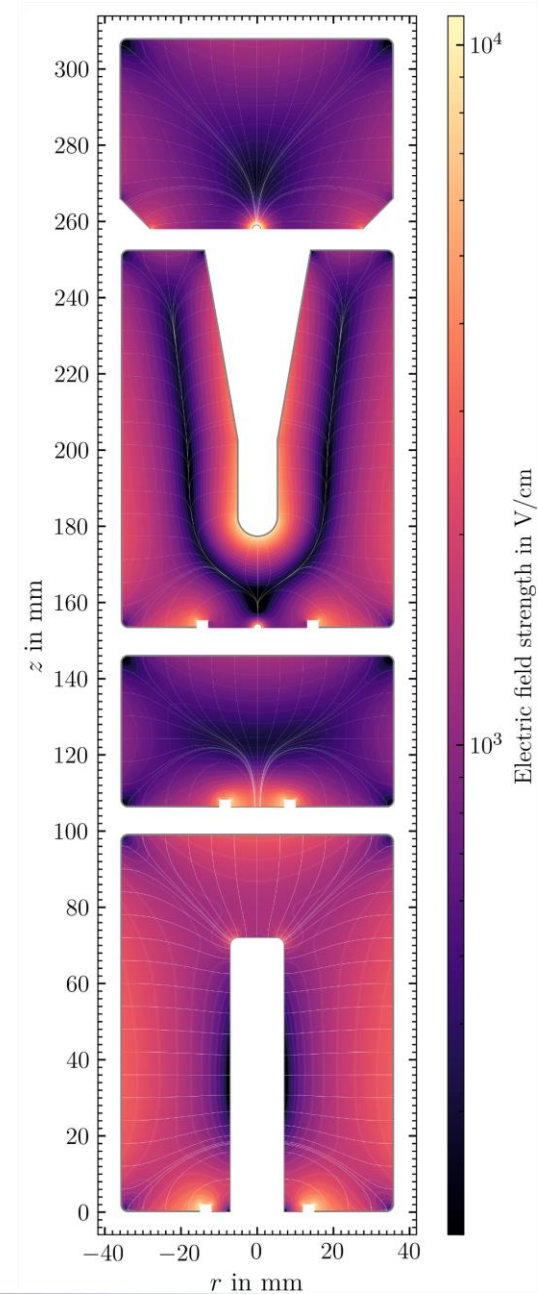


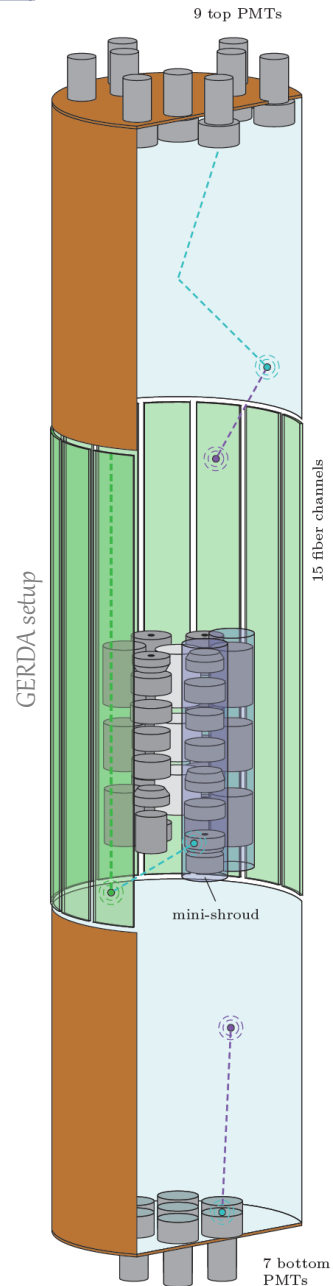
Broad Energy
Germanium (BEGe)

P-type Point-
Contact (PPC)

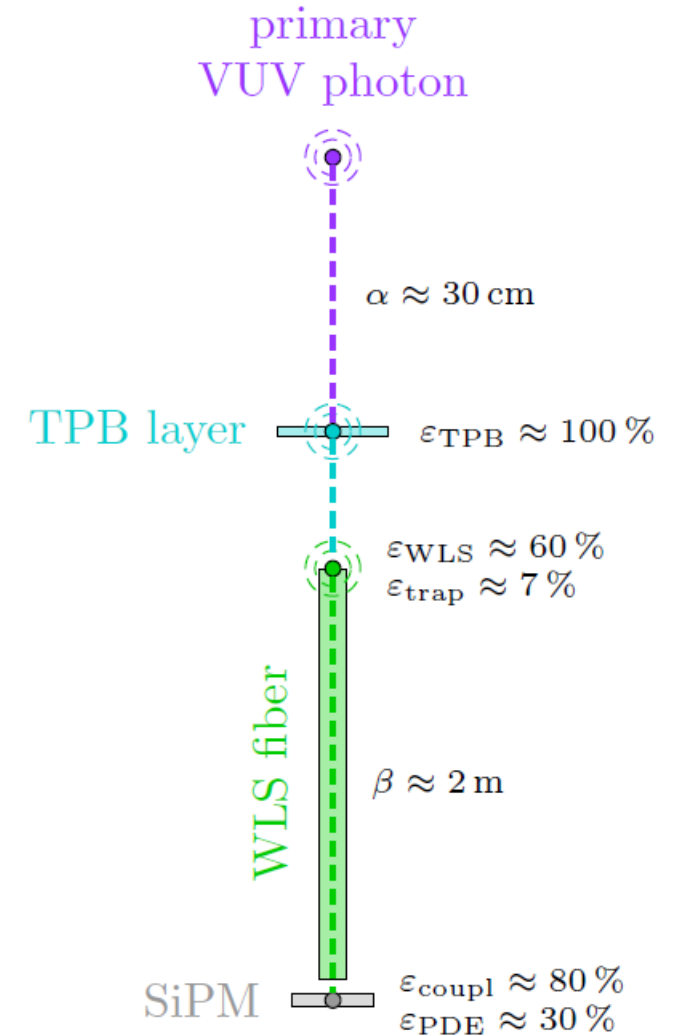
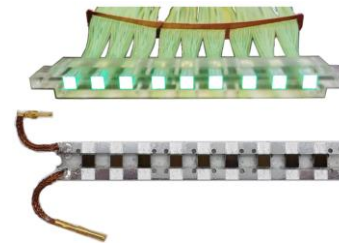
Coaxial

Inverted
Coaxial (IC)

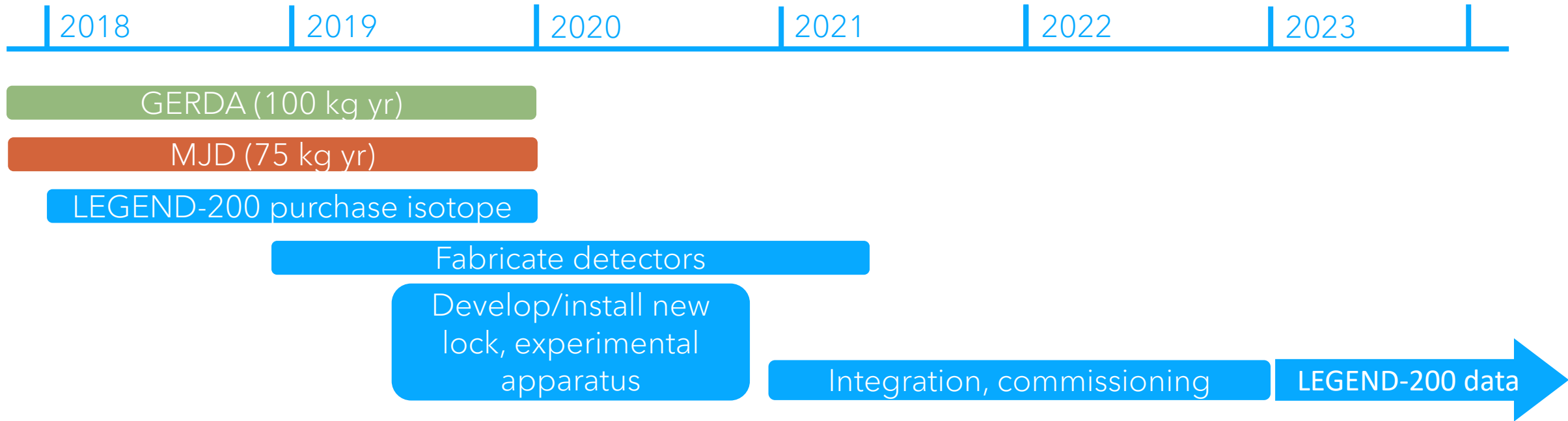




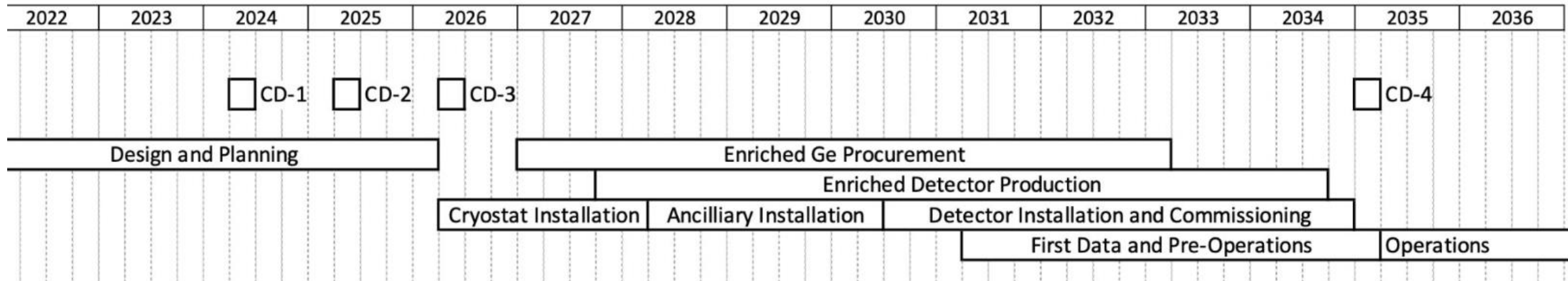
- Two barrels in LAr @ $\sim 84\text{K}$
 - $\varnothing 590\text{ mm}$, H 1500 mm
 - *Inner Barrel*, IB: 9 fibres modules = 18 SiPM channels
 - *Outer Barrel*, OB: 20 fibres modules = 40 SiPM channels
- Array of 9 SiPMs read in parallel
 - Ketek PM33100T 3x3 mm
 - 81 WaveLength Shifting (WLS) fibers
- Nylon Mini-shrouds (MS)
 - Suppression of K-42 ions background
 - Transparent to LAr scintillation light (covered with TetraPhenyl Butadiene, TPB)
 - Nylon: low-background material
- Fiber shroud cylinder
 - $\varnothing 1375\text{ mm}$, H 3000 mm
 - Reflected photons reach fibers without any LAr attenuation
 - Help to detect photons produced outside the barrels

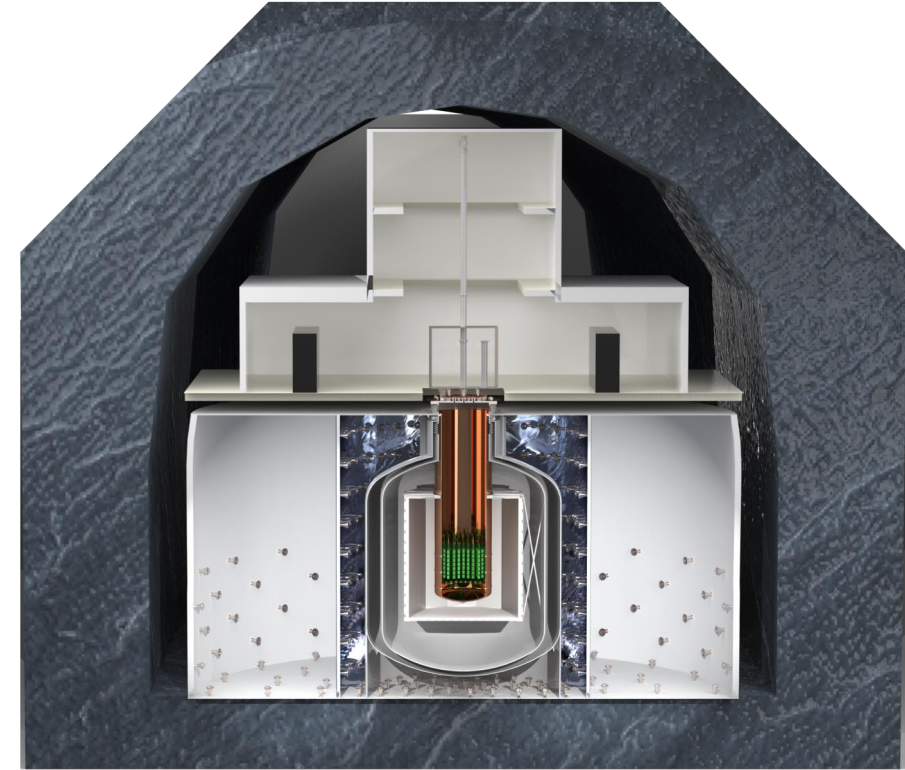
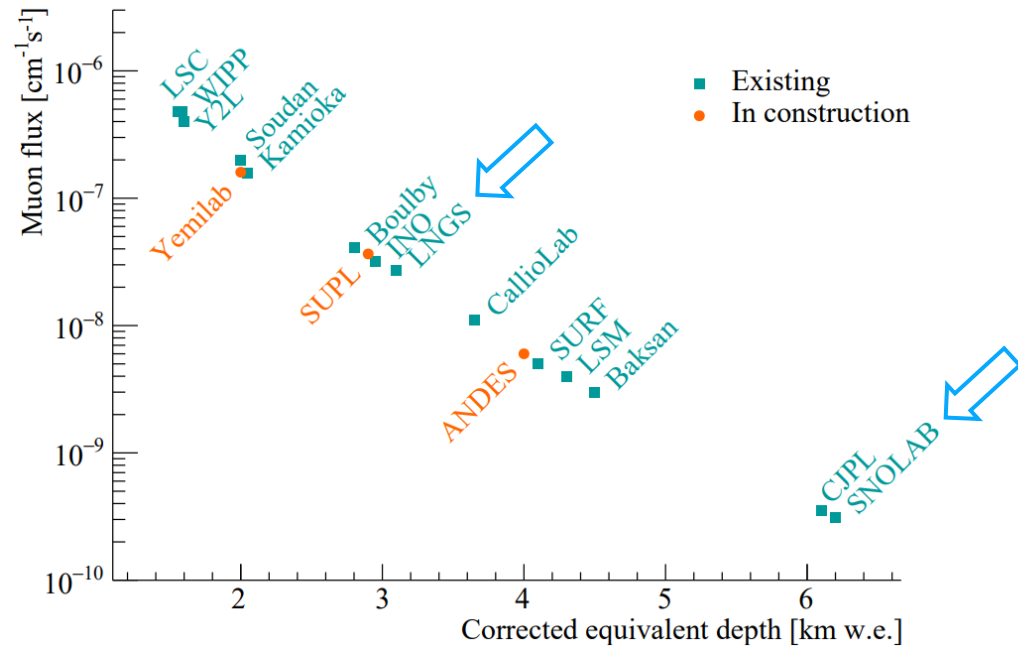


GERDA collaboration, "Liquid argon light collection and veto modeling in GERDA Phase II", [arXiv:2212.02856](https://arxiv.org/abs/2212.02856) (2022)



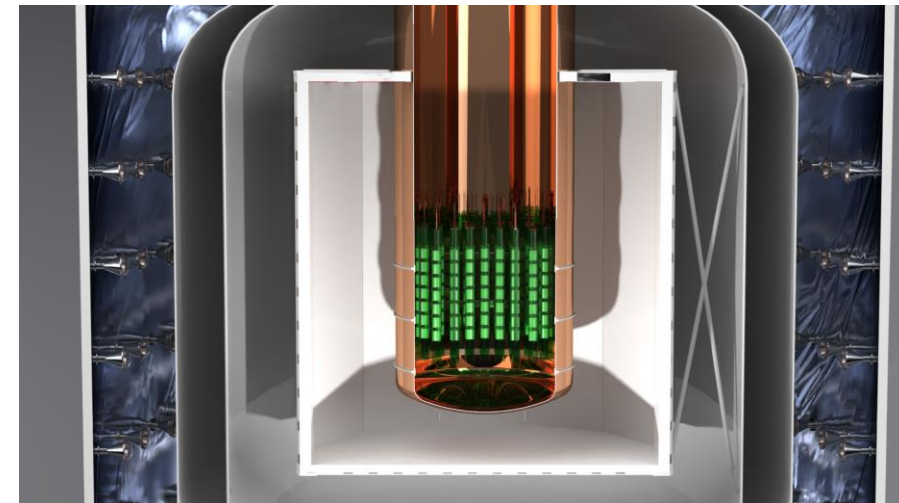
LEGEND-1000

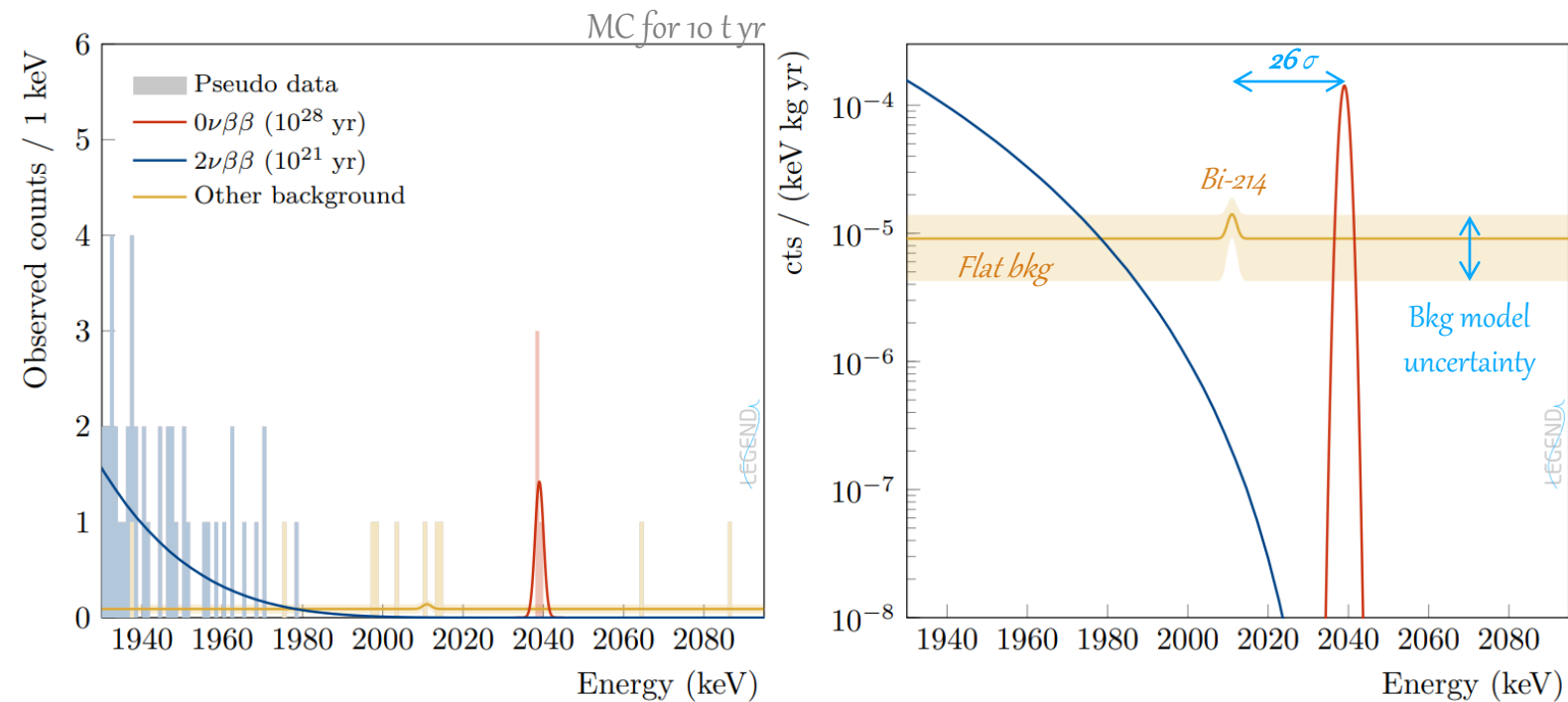




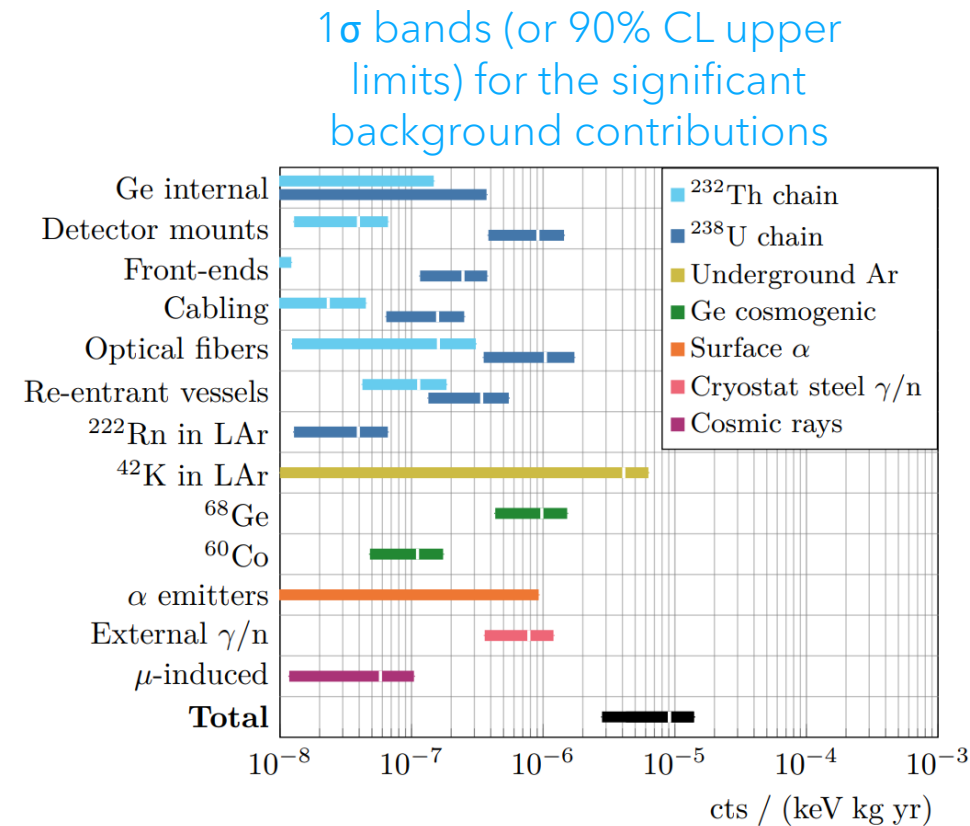
- Baseline design: **1000 kg** of germanium in **4 re-entrant tubes** containing **LAr underground** (reduces ^{42}K produced from ^{42}Ar)
- Background level reduced by a factor 20 wrt LEGEND-200
- **SNOLAB** (6010 m w.e.) or **LNGS** (3600 m w.e.) – DOE, ~July '23
- CD-0 approved by the United States Department of Energy (DOE)

Pre-conceptual Design Report: <https://inspirehep.net/literature/1892243>

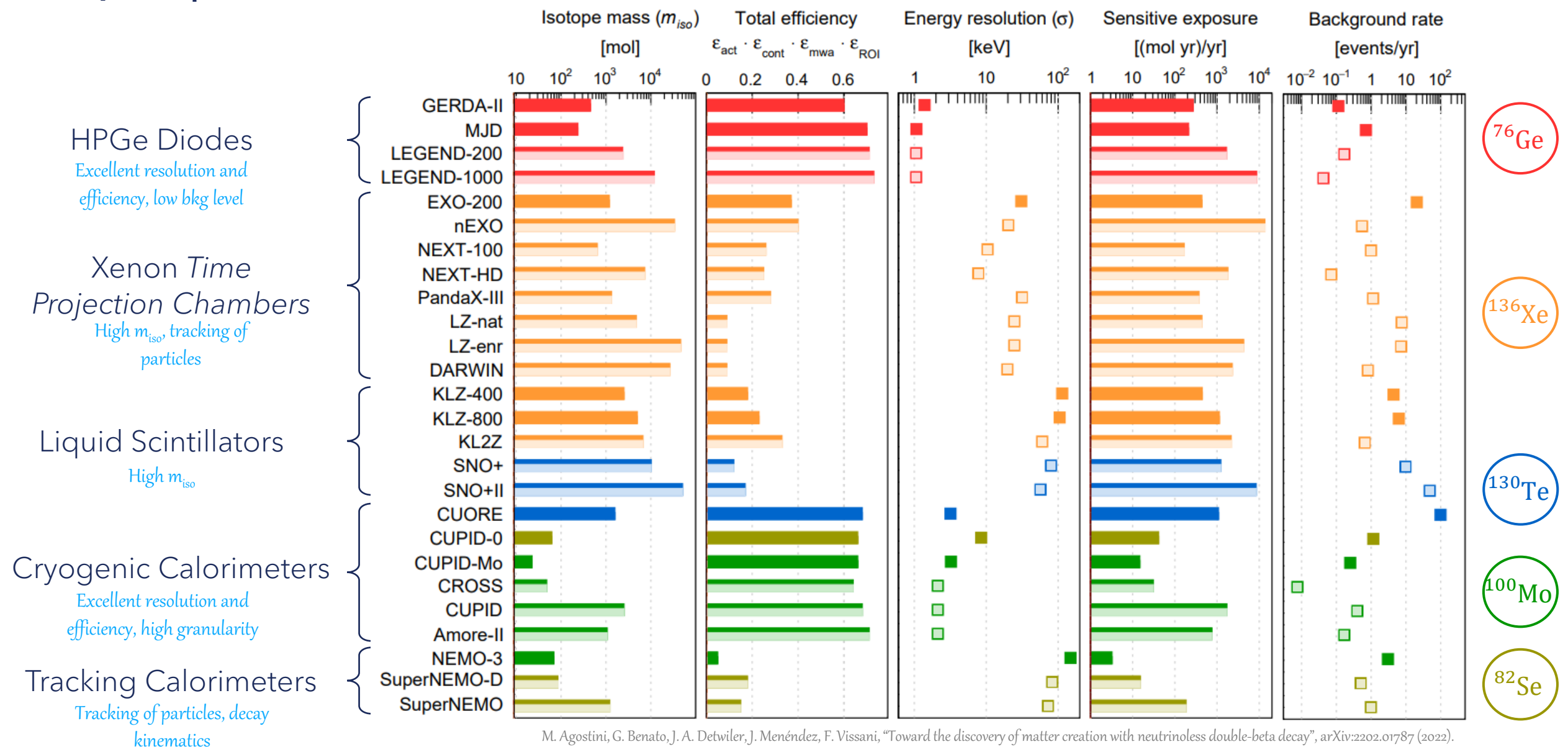




- Background contributions derived from the **background model of LEGEND-1000** (calibrated over data collected by GERDA & MJD)
- Less than 1 background event expected around the $Q_{\beta\beta}$ peak
- No $2\nu 2\beta$ in proximity of $Q_{\beta\beta}$
- **3-4 signal events**, visible by eye



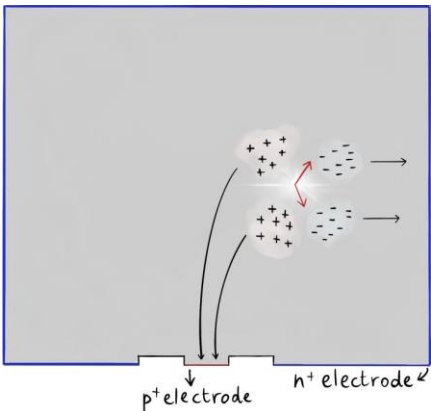
Pre-conceptual Design Report: <https://inspirehep.net/literature/1892243>



M. Agostini, G. Benato, J. A. Detwiler, J. Menéndez, F. Vissani, "Toward the discovery of matter creation with neutrinoless double-beta decay", arXiv:2202.01787 (2022).

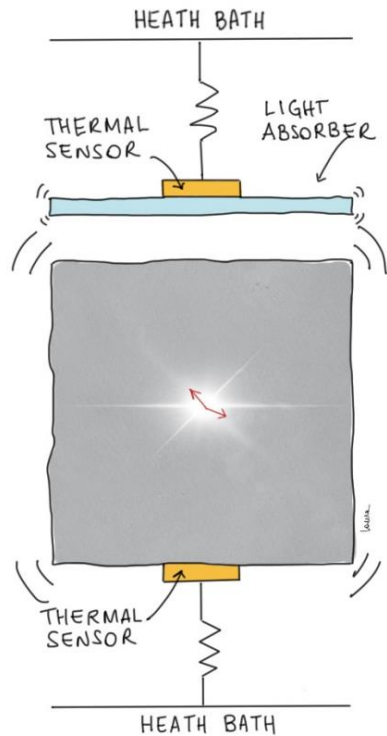
M. Agostini, G. Benato, J. A. Detwiler, J. Menéndez, F. Vissani, “Toward the discovery of matter creation with neutrinoless double-beta decay”, arXiv:2202.01787 (2022).

Experiment	Isotope	Status	Lab	m_{iso} [mol]	ε_{act} [%]	$\varepsilon_{\text{cont}}$ [%]	ε_{mva} [%]	σ [keV]	ROI [σ]	ε_{ROI} [%]	$\mathcal{E}$ [$\frac{\text{mol}\cdot\text{yr}}{\text{yr}}$]	$\mathcal{B}$ [$\frac{\text{events}}{\text{mol}\cdot\text{yr}}$]	λ_b [$\frac{\text{events}}{\text{yr}}$]	$T_{1/2}$ [yr]	$m_{\beta\beta}$ [meV]
High-purity Ge detectors (Sec. VI.B)															
GERDA-II	^{76}Ge	completed	LNGS	$4.5 \cdot 10^2$	88	91	79	1.4	-2,2	95	273	$4.2 \cdot 10^{-4}$	$1.1 \cdot 10^{-1}$	$1.2 \cdot 10^{26}$	93-222
MJD	^{76}Ge	completed	SURF	$3.1 \cdot 10^2$	91	91	86	1.1	-2,2	95	212	$3.3 \cdot 10^{-3}$	$7.1 \cdot 10^{-1}$	$4.7 \cdot 10^{25}$	149-355
LEGEND-200	^{76}Ge	construction	LNGS	$2.4 \cdot 10^3$	91	91	90	1.1	-2,2	95	1 684	$1.0 \cdot 10^{-4}$	$1.7 \cdot 10^{-1}$	$1.5 \cdot 10^{27}$	27-63
LEGEND-1000	^{76}Ge	proposed		$1.2 \cdot 10^4$	92	92	90	1.1	-2,2	95	8 736	$4.9 \cdot 10^{-6}$	$4.3 \cdot 10^{-2}$	$1.3 \cdot 10^{28}$	9-21
Xenon time projection chambers (Sec. VI.C)															
EXO-200	^{136}Xe	completed	WIPP	$1.2 \cdot 10^3$	46	100	84	31	-2,2	95	438	$4.7 \cdot 10^{-2}$	$2.1 \cdot 10^{+1}$	$2.4 \cdot 10^{25}$	111-477
nEXO	^{136}Xe	proposed	SNOLAB	$3.4 \cdot 10^4$	64	100	66	20	-2,2	95	13 700	$4.0 \cdot 10^{-5}$	$5.5 \cdot 10^{-1}$	$7.4 \cdot 10^{27}$	6-27
NEXT-100	^{136}Xe	construction	LSC	$6.4 \cdot 10^2$	88	76	49	10	-1.0,1.8	80	167	$5.9 \cdot 10^{-3}$	$9.9 \cdot 10^{-1}$	$7.0 \cdot 10^{25}$	66-281
NEXT-HD	^{136}Xe	proposed		$7.4 \cdot 10^3$	95	89	44	7.7	-0.5,1.7	65	1 809	$4.0 \cdot 10^{-5}$	$7.2 \cdot 10^{-2}$	$2.2 \cdot 10^{27}$	12-50
PandaX-III-200	^{136}Xe	construction	CJPL	$1.3 \cdot 10^3$	77	74	65	31	-1.2,1.2	76	374	$3.0 \cdot 10^{-3}$	$1.1 \cdot 10^{+0}$	$1.5 \cdot 10^{26}$	45-194
LZ-nat	^{136}Xe	construction	SURF	$4.7 \cdot 10^3$	14	100	80	25	-1.4,1.4	84	440	$1.7 \cdot 10^{-2}$	$7.5 \cdot 10^{+0}$	$7.2 \cdot 10^{25}$	64-277
LZ-enr	^{136}Xe	proposed	SURF	$4.6 \cdot 10^4$	14	100	80	25	-1.4,1.4	84	4302	$1.7 \cdot 10^{-3}$	$7.3 \cdot 10^{+0}$	$7.1 \cdot 10^{26}$	20-87
Darwin	^{136}Xe	proposed		$2.7 \cdot 10^4$	13	100	90	20	-1.2,1.2	76	2 312	$3.5 \cdot 10^{-4}$	$8.0 \cdot 10^{-1}$	$1.1 \cdot 10^{27}$	17-72
Large liquid scintillators (Sec. VI.D)															
KLZ-400	^{136}Xe	completed	Kamioka	$2.5 \cdot 10^3$	44	100	97	114	0,1.4	42	450	$9.8 \cdot 10^{-3}$	$4.4 \cdot 10^{+0}$	$3.3 \cdot 10^{25}$	95-408
KLZ-800	^{136}Xe	taking data	Kamioka	$5.0 \cdot 10^3$	55	100	100	105	0,1.4	42	1 143	$5.5 \cdot 10^{-3}$	$6.2 \cdot 10^{+0}$	$2.0 \cdot 10^{26}$	38-164
KL2Z	^{136}Xe	proposed	Kamioka	$6.7 \cdot 10^3$	80	100	97	60	0,1.4	42	2 176	$3.0 \cdot 10^{-4}$	$6.5 \cdot 10^{-1}$	$1.1 \cdot 10^{27}$	17-71
SNO+I	^{130}Te	construction	SNOLAB	$1.0 \cdot 10^4$	20	100	97	80	-0.5,1.5	62	1 232	$7.8 \cdot 10^{-3}$	$9.7 \cdot 10^{+0}$	$1.8 \cdot 10^{26}$	31-144
SNO+II	^{130}Te	proposed	SNOLAB	$5.1 \cdot 10^4$	27	100	97	57	-0.5,1.5	62	8 521	$5.7 \cdot 10^{-3}$	$4.8 \cdot 10^{+1}$	$5.7 \cdot 10^{26}$	17-81
Cryogenic calorimeters (Sec. VI.E)															
CUORE	^{130}Te	taking data	LNGS	$1.6 \cdot 10^3$	100	88	92	3.2	-1.4,1.4	84	1 088	$9.1 \cdot 10^{-2}$	$9.9 \cdot 10^{+1}$	$5.1 \cdot 10^{25}$	58-270
CUPID-0	^{82}Se	completed	LNGS	$6.2 \cdot 10^1$	100	81	86	8.5	-2,2	95	41	$2.8 \cdot 10^{-2}$	$1.2 \cdot 10^{+0}$	$4.4 \cdot 10^{24}$	283-551
CUPID-Mo	^{100}Mo	completed	LSM	$2.3 \cdot 10^1$	100	76	91	3.2	-2,2	95	15	$1.7 \cdot 10^{-2}$	$2.5 \cdot 10^{-1}$	$1.7 \cdot 10^{24}$	293-858
CROSS	^{100}Mo	construction	LSC	$4.8 \cdot 10^1$	100	75	90	2.1	-2,2	95	31	$2.5 \cdot 10^{-4}$	$7.6 \cdot 10^{-3}$	$4.9 \cdot 10^{25}$	54-160
CUPID	^{100}Mo	proposed	LNGS	$2.5 \cdot 10^3$	100	79	90	2.1	-2,2	95	1 717	$2.3 \cdot 10^{-4}$	$4.0 \cdot 10^{-1}$	$1.1 \cdot 10^{27}$	12-34
AMoRE-II	^{100}Mo	proposed	Yemilab	$1.1 \cdot 10^3$	100	82	91	2.1	-2,2	95	760	$2.2 \cdot 10^{-4}$	$1.7 \cdot 10^{-1}$	$6.7 \cdot 10^{26}$	15-43
Tracking calorimeters (Sec. VI.F)															
NEMO-3	^{100}Mo	completed	LSM	$6.9 \cdot 10^1$	100	100	11	148	-1.6,1.1	42	3	$9.4 \cdot 10^{-1}$	$3.0 \cdot 10^{+0}$	$5.6 \cdot 10^{23}$	505-1485
SuperNEMO-D	^{82}Se	construction	LSM	$8.5 \cdot 10^1$	100	100	28	83	-4.2,2.4	64	15	$3.3 \cdot 10^{-2}$	$5.0 \cdot 10^{-1}$	$8.6 \cdot 10^{24}$	201-391
SuperNEMO	^{82}Se	proposed	LSM	$1.2 \cdot 10^3$	100	100	28	72	-4.1,2.8	54	185	$5.3 \cdot 10^{-3}$	$9.8 \cdot 10^{-1}$	$7.8 \cdot 10^{25}$	67-131



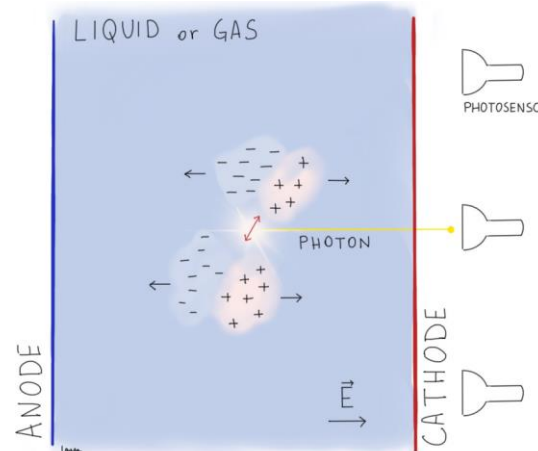
High-purity Ge (HPGe) detector

$e^- - h$ clusters created by ionization are collected to the electrodes by an electric field



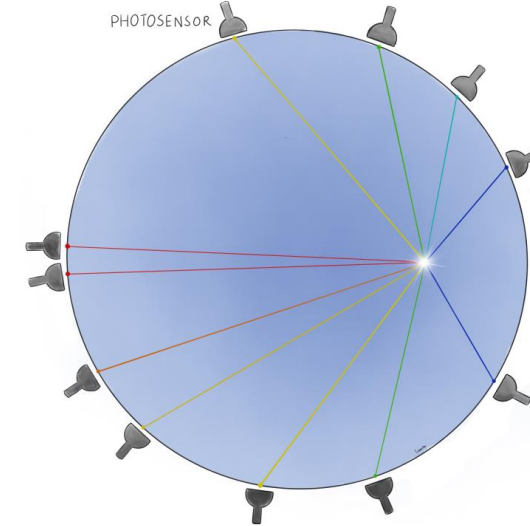
Cryogenic Calorimeters

Phonon and scintillation light signals are read-out through superconductive thermal sensors



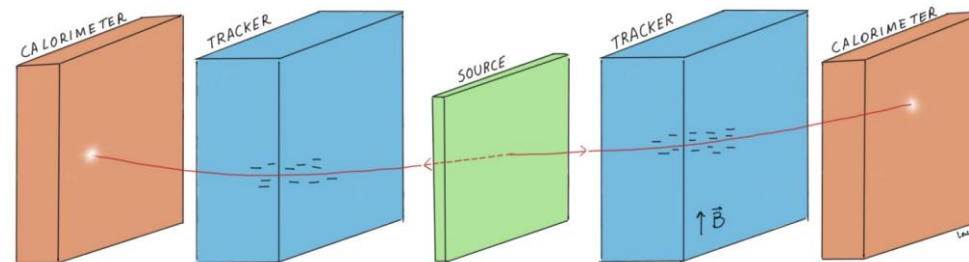
Xe Time Projection Chambers

$e^- - h$ clusters created by ionization are collected to the electrodes by an electric field. In addition, scintillation light is detected by light sensors, providing the timing of the event



Liquid Scintillators

The position of an event can be reconstructed through the time of flight of the scintillation photons



Tracking Calorimeters

The charge, momentum, and energy of the particles ejected by the source is measured through a combination of magnetic-field tracks and calorimeters



May 2023 - Roma Tre



CIEMAT
Comenius Univ.
Czech Tech. Univ. Prague and IEAP
Daresbury Lab.
Duke Univ. and TUNL
Gran Sasso Science Inst.
Indiana Univ. Bloomington
Inst. Nucl. Res. Rus. Acad. Sci.
Jagiellonian Univ.
Joint Inst. for Nucl. Res.
Joint Res. Centre Geel
Lab. Naz. Gran Sasso
Lancaster Univ.
Leibniz Inst. for Crystal Growth

Leibniz Inst. for Polymer Research
Los Alamos Natl. Lab.
Max Planck Inst. for Nucl. Phys.
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Natl. Res. Center Kurchatov Inst.
Natl. Res. Nucl. Univ. MEPhI
North Carolina State Univ.
Oak Ridge Natl. Lab.
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Roma Tre Univ. and INFN
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Univ. of California and LBNL
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Univ. of Cagliari and INFN
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Univ. of Regina
Univ. of South Carolina
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