

FIMPs: from freeze-in to conversion freeze-out

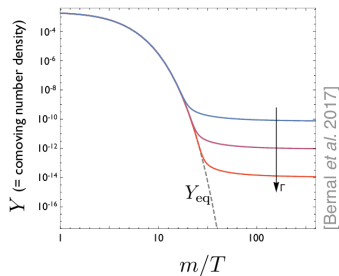
Laura Lopez Honorez



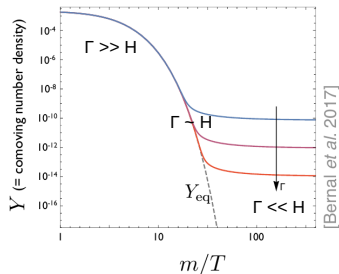
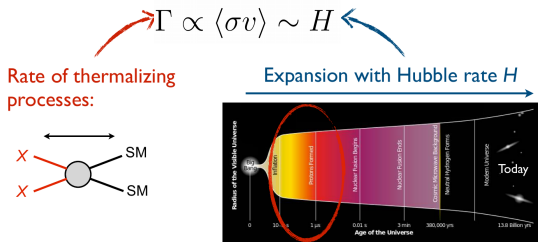
mainly based on **JHEP 1907 (2019) 136** & **JHEP 1809 (2018) 037**
in collaboration with L. Calibbi, S. Junius, S. Lowette & A. Mariotti

Scalars 2019 - Warsaw - 11-14/9/19

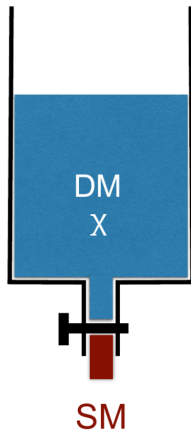
WIMP Freeze-out: The simple picture



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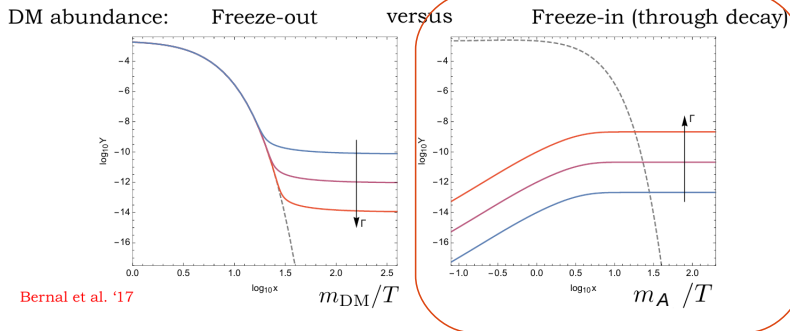
WIMP Freeze-out: The simple picture



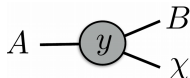
Freeze-out
DM annihilation driven

$$\Omega h^2 \propto 1/\langle\sigma v\rangle_{\chi\chi}$$

Freeze-out versus Freeze-in: going to $y \ll 1$

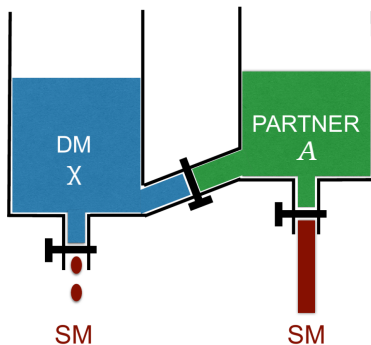


$$\Omega h^2 \sim 0.12 \left(\frac{\Gamma_A}{4 \times 10^{-15} \text{ GeV}} \right) \left(\frac{600 \text{ GeV}}{m_A} \right)^2 \left(\frac{m_{DM}}{10 \text{ keV}} \right)$$



$$y \sim 10^{-8} \quad \text{for} \quad \Gamma_A \sim \frac{y^2}{8\pi} m_A$$

Annihilations vs coannihilations: going to $\Delta m \ll m$



Freeze-out
Partner annihilation driven

$$\Omega h^2 \propto 1/\langle\sigma v\rangle_{eff}$$

$$\langle\sigma v\rangle_{eff} \propto e^{-2x_f \frac{\Delta m}{m_\chi}} \langle\sigma v\rangle_{AA}$$

$$\Delta m = m_A - m_\chi$$

Simplified Model



Simplified Model



Simplified Model: 3 extra parameters m_χ, m_A, y

Dark matter χ ($\sim$ Majorana) fermion coupled to dark A and SM B through Yukawa-like interactions [see E. Ma, F. Kahlhoefer & O. Lebedev talks for FI with A from visible sector]

$$\mathcal{L} \subset y \chi AB$$

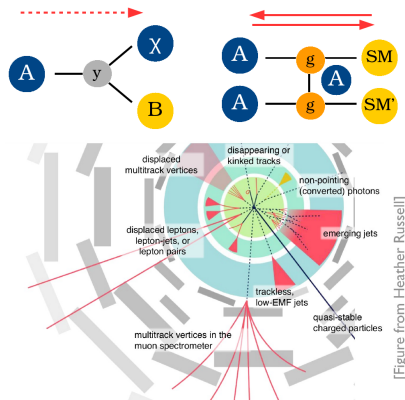
- Dark sector (Z_2 odd): $m_A > m_\chi$

Simplified Model: 3 extra parameters m_χ, m_A, y

Dark matter χ ($\sim$ Majorana) fermion coupled to dark A and SM B through Yukawa-like interactions [see E. Ma, F. Kahlhoefer & O. Lebedev talks for FI with A from visible sector]

$$\mathcal{L} \subset y \chi A B$$

- Dark sector (Z_2 odd): $m_A > m_\chi$
- A is $SU(2) \times U(1)$ charged
 - fast $A^\dagger A \leftrightarrow \text{SM SM}$ through gauge interactions at early time
 - A is produced at colliders today
- χ - A -SM interactions:
 - $\chi \equiv \text{FIMP} \leftrightarrow y \ll 10^{-4}$
 - long lived A at colliders through $A \rightarrow B\chi$



[Figure from Heather Russell]

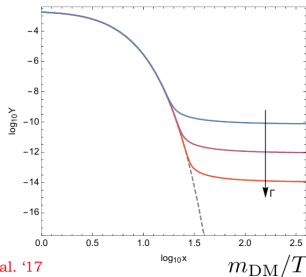
[See also 1903.04497]

Frozen in FIMP:

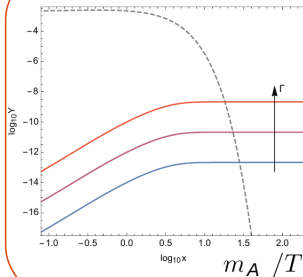
Large mass splittings within the singlet doublet case

Freeze-in

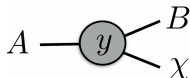
DM abundance: Freeze-out versus Freeze-in (through decay)



Bernal et al. '17



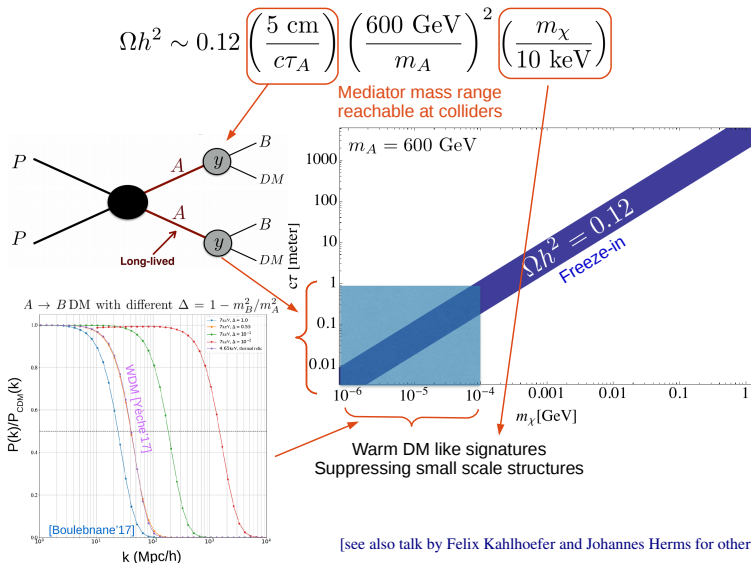
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$$y \sim 10^{-8} \quad \text{for} \quad \Gamma_A \sim \frac{y^2}{8\pi} m_A$$

Freeze-in : displaced vertices and cosmology interplay

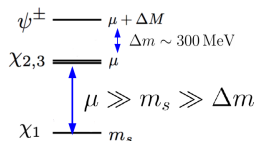
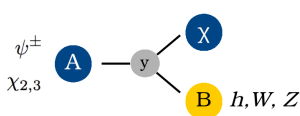
see also e.g. [Hall'09, Co'15, Hessler'16, d'Eramo'17, Buchmueller'17, Heeck'17, Boulebnane'17, Brooijmans'18, Garny'18]



[see also talk by Felix Kahlhoefer and Johannes Herms for other probes of 

Framework: H-coupled fermionic Dark Matter

see also [Mahubani'05, D'Eramo'07, Enberg'07, Cohen'11, Clifford'14, Calibi'15, Freitas'15; etc]



$$-\mathcal{L} \supset \mu \psi_d \cdot \psi_u + y (\psi_d \cdot H \psi_s + H^\dagger \psi_u \psi_s) + m_s/2 \psi_s \psi_s + h.c.,$$

- SM + 3 dark fermions $\rightsquigarrow$ Z_2 symmetry for DM stability

$$(\psi_u)_{2, \frac{1}{2}} = \begin{pmatrix} \psi^+ \\ \psi_u^0 \end{pmatrix}, \quad (\psi_d)_{2, -\frac{1}{2}} = \begin{pmatrix} \psi_d^0 \\ \psi^- \end{pmatrix}, \quad (\psi_s)_{1, 0}$$

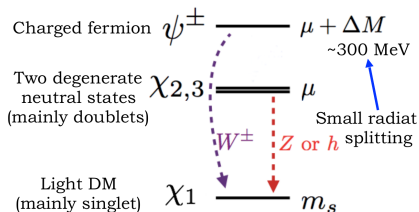
$$(\chi_1, \chi_2, \chi_3)^T = U (\psi_s, \psi_d^0, \chi_u^0)^T$$

- We work in the limit: $y \ll 1$ and $\mu \gg m_s$

Freeze-in from doublet decay

Freeze-in in the limit of $y \ll 1$ & $\mu \ll m_s$:

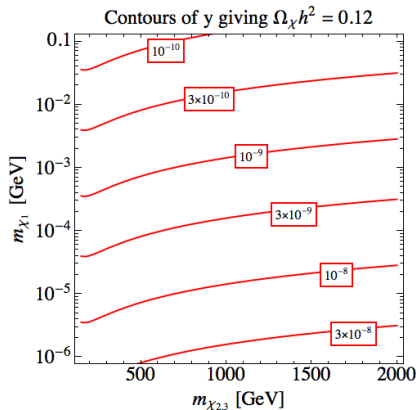
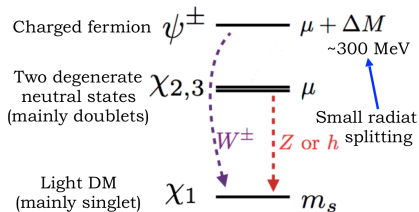
$$Y_{\chi_1} = \frac{270 M_{Pl}}{(1.66) 8\pi^3 g_*^{3/2}} \left(\sum_{B=Z,h} \frac{\Gamma[\chi_3 \rightarrow B\chi_1]}{m_{\chi_3}^2} + \sum_{B=Z,h} \frac{\Gamma[\chi_2 \rightarrow B\chi_1]}{m_{\chi_2}^2} + g_\psi \frac{\Gamma[\psi^+ \rightarrow W^+\chi_1]}{m_\psi^2} \right)$$



Freeze-in from doublet decay

Freeze-in in the limit of $y \ll 1$ & $\mu \ll m_s$:

$$Y_{\chi_1} = \frac{270 M_{Pl}}{(1.66) 8\pi^3 g_*^{3/2}} \left(\sum_{B=Z,h} \frac{\Gamma[\chi_3 \rightarrow B\chi_1]}{m_{\chi_3}^2} + \sum_{B=Z,h} \frac{\Gamma[\chi_2 \rightarrow B\chi_1]}{m_{\chi_2}^2} + g_\psi \frac{\Gamma[\psi^+ \rightarrow W^+\chi_1]}{m_\psi^2} \right)$$



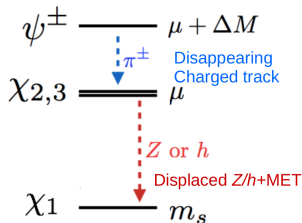
$\rightsquigarrow$ relevant H -coupling for freeze-in in 1M2D: $y < \mathcal{O}(10^{-8})$ for $m_{DM} > \text{few keV}$

Results confirmed using micrOMEGAS5.0 [Belanger'18]

Disappearing track or displaced h/Z+ MET

Copious production of $\chi_{2,3}, \psi^\pm$ at colliders through EW processes

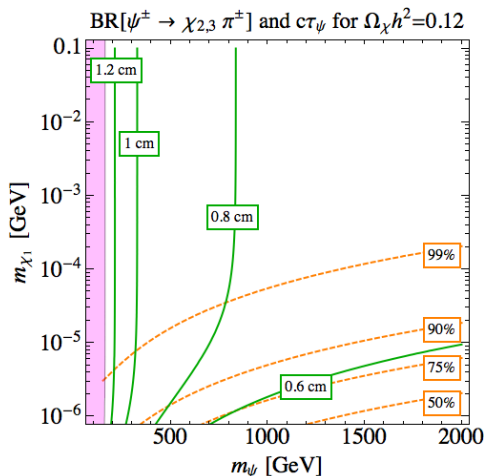
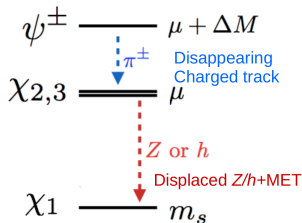
- $\psi^\pm \rightarrow \pi^\pm \chi_{2,3}$
- $\chi_2 \chi_3 \rightarrow ZZ/hh/Zh + \text{MET}$



Disappearing track or displaced h/Z+ MET

Copious production of $\chi_{2,3}, \psi^\pm$ at colliders through EW processes

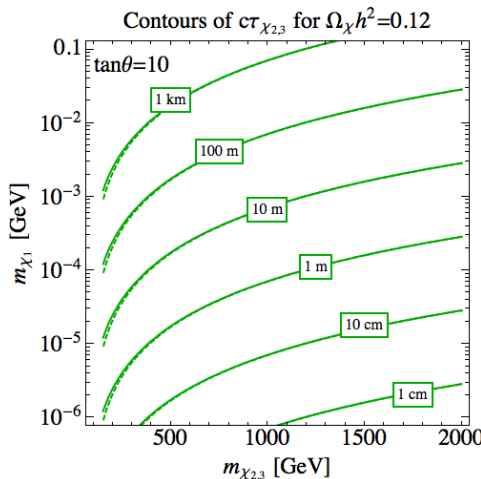
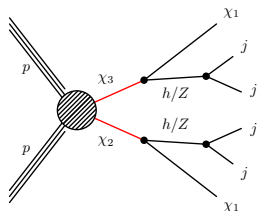
- $\psi^\pm \rightarrow \pi^\pm \chi_{2,3}$
 $\rightsquigarrow$ disappearing charged track
 with $c\tau \sim \mathcal{O}(\text{cm})$
 $\rightsquigarrow \mu > 150 \text{ GeV}$ from [ATLAS DT'17]
- $\chi_2 \chi_3 \rightarrow ZZ/hh/Zh + \text{MET}$



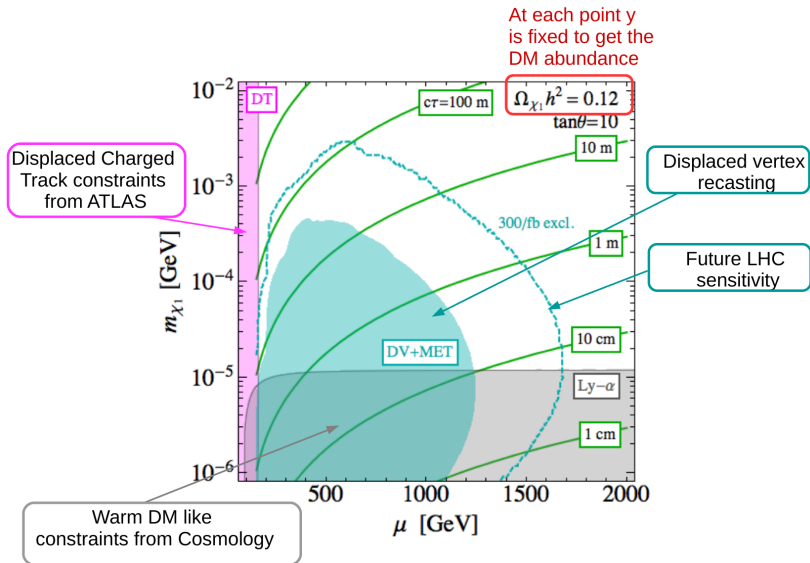
Disappearing track or displaced h/Z+ MET

Copious production of $\chi_{2,3}, \psi^\pm$ at colliders through EW processes

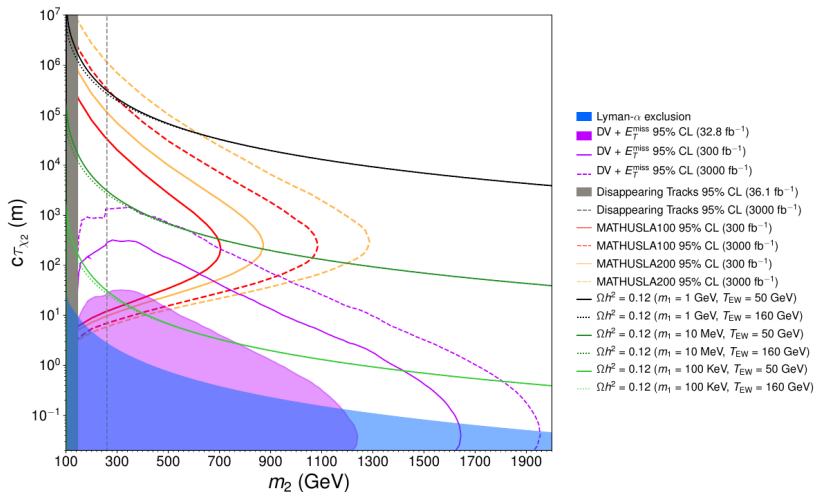
- $\psi^\pm \rightarrow \pi^\pm \chi_{2,3}$
 $\rightsquigarrow$ **disappearing charged track**
 with $c\tau \sim \mathcal{O}(\text{cm})$
 $\rightsquigarrow \mu > 150 \text{ GeV}$ from [ATLAS DT'17]
- $\chi_2 \chi_3 \rightarrow ZZ/hh/Zh + \text{MET}$
 $\rightsquigarrow$ **displaced vertex** with jet and MET with possibly $c\tau > \mathcal{O}(\text{cm})$
 $\rightsquigarrow$ **probe $\mu \leq 1.2 \text{ TeV}$** [ATLAS DV'17]



LHC & Cosmo complementarity



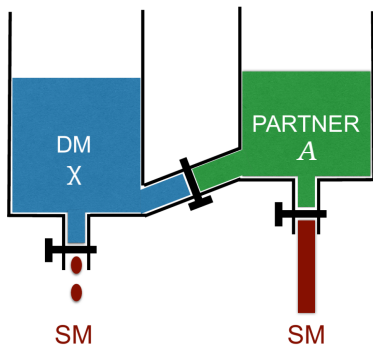
LHC & Cosmo complementarity



[No, Tunney & Zaldivar'0919]

Conversion driven freeze-out Leptophilic DM

Annihilations vs coannihilations: going to $\Delta m \ll m$



Freeze-out
Partner annihilation driven

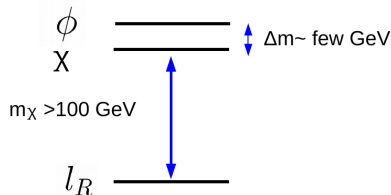
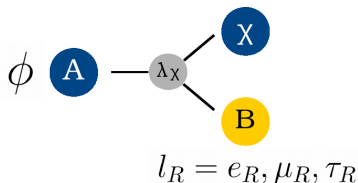
$$\Omega h^2 \propto 1/\langle\sigma v\rangle_{eff}$$

$$\langle\sigma v\rangle_{eff} \propto e^{-2x_f \frac{\Delta m}{m_\chi}} \langle\sigma v\rangle_{AA}$$

$$\Delta m = m_A - m_\chi$$

Framework: Leptophilic DM with compressed spectrum

see also [Berstrom '89+, Bringmann '08+, Ciafaloni '11, Garny '11+, Toma '13, Giacchino '13+, Ibarra '14, Belanger '18, Calibbi '18...]

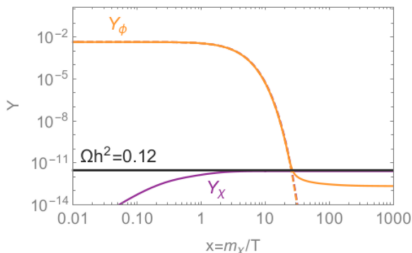
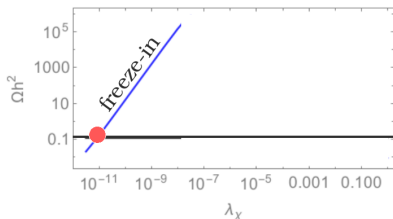


$$\mathcal{L} \subset \mathcal{L}_K - \frac{m_\chi}{2} \bar{\chi} \chi - m_\phi \phi^\dagger \phi - \lambda_\chi \phi \bar{\chi} l_R + \lambda_H |\phi|^2 |H|^2 + h.c.$$

- SM + 1 charged dark scalar ϕ + 1 Majorana dark fermions χ (Z_2 symmetry for DM stability)
- $m_\phi > m_\chi$ and we are mostly interested in $\lambda_\chi \ll 10^{-4}$
 $\rightsquigarrow \phi$ LLP at colliders
- We work in the compressed limit: $\Delta m = m_\phi - m_\chi < m_\chi$

From freeze-in to Freeze-out through coannihilations

$$\mathcal{L} \subset -\lambda_\chi \phi \bar{\chi} l_R, m_\chi=150\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=8 \cdot 10^{-12}, \lambda_H=0.1$$



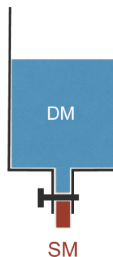
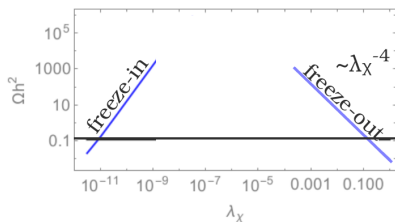
Freeze-in: feeble couplings $\lambda_\chi < 10^{-8}$

$$Y_\chi \propto \Gamma_{\phi \rightarrow \chi \ell} \propto \lambda_\chi^2$$

- **Annihilations:** $\phi\phi^\dagger \rightarrow \text{SM SM}$
keep ϕ in eq. until ϕ fo (rate $\propto g^4$)
- **Conversions:** $\phi \rightarrow \ell\chi$ and $\phi\text{SM} \rightarrow \chi\text{SM}$
DM production (rate $\propto \lambda_\chi^2, \lambda_\chi^2 g^2$)

From freeze-in to Freeze-out through coannihilations

$$\mathcal{L} \subset -\lambda_\chi \phi \bar{\chi} l_R$$

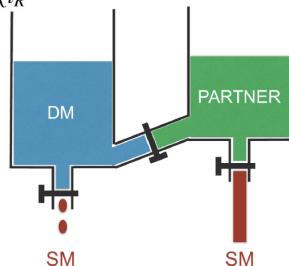
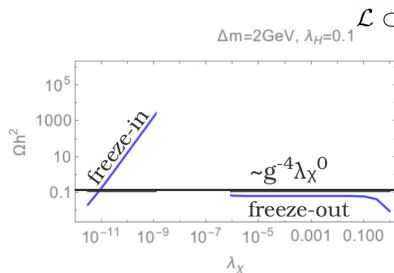


DM annihilations driven Freeze-out: weak couplings $\lambda_\chi \gtrsim 0.1$

$$Y_\chi \propto 1/\sigma v_{\chi\chi} \propto \lambda_\chi^{-4}$$

- **Annihilations:** $\chi\chi \rightarrow \ell\ell$
keep χ in eq. until χ fo (rate $\propto \lambda_\chi^4$)

From freeze-in to Freeze-out through coannihilations



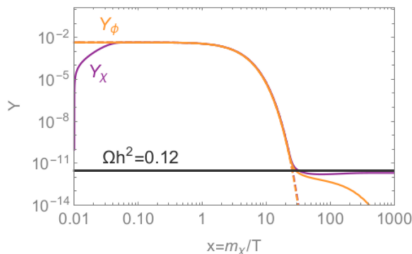
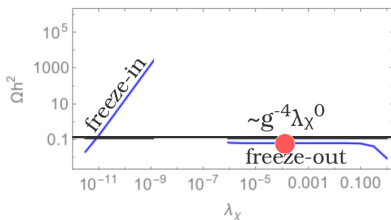
Mediator annihilation driven Freeze-out: weak couplings $\lambda_\chi \lesssim 0.1$

$$Y_\chi \propto 1/\sigma v_{\phi\phi^\dagger} \propto \lambda_\chi^0$$

- small mediator-DM splitting $\Delta m/m_\chi \lesssim 0.1$
- **Annihilations:** $\phi\phi^\dagger \rightarrow \text{SM SM}$
keep ϕ in eq. until ϕ fo (rate $\propto g^4$)
- **Conversions:** $\phi \rightarrow \ell\chi$ and $\phi\text{SM} \rightarrow \chi\text{SM}$
keep ϕ and χ in chemical equ. (rate $\propto \lambda_\chi^2, \lambda_\chi^2 g^2$)

From freeze-in to Freeze-out through coannihilations

$$\Delta m = 2\text{GeV}, \lambda_H = 0.1 \quad \mathcal{L} \subset -\lambda_\chi \phi \bar{\chi} l_R \quad m_\chi = 150\text{GeV}, \Delta m = 2\text{GeV}, \lambda_\chi = 10^{-4}, \lambda_H = 0.1$$

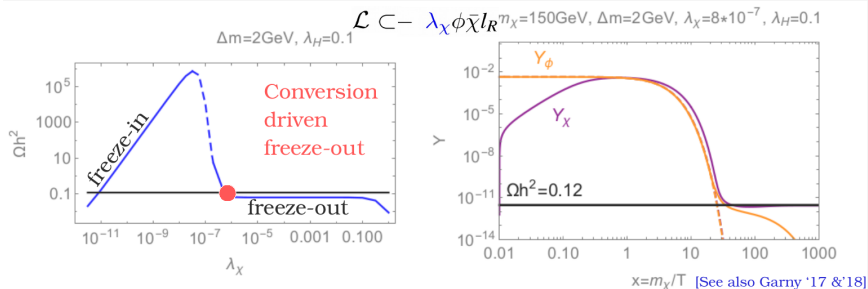


Mediator annihilation driven Freeze-out: weak couplings $\lambda_\chi \lesssim 0.1$

$$Y_\chi \propto 1/\sigma v_{\phi\phi^\dagger} \propto \lambda_\chi^0$$

- small mediator-DM splitting $\Delta m/m_\chi \lesssim 0.1$
- **Annihilations:** $\phi\phi^\dagger \rightarrow \text{SM SM}$
keep ϕ in eq. until ϕ fo (rate $\propto g^4$)
- **Conversions:** $\phi \rightarrow \ell\chi$ and $\phi\text{SM} \rightarrow \chi\text{SM}$
keep ϕ and χ in chemical equ. (rate $\propto \lambda_\chi^2, \lambda_\chi^2 g^2$)

From freeze-in to Freeze-out through coannihilations



Conversion driven Freeze-out: feeble couplings $\lambda_\chi \lesssim 10^{-7}$

$Y_\chi \nearrow$ when $\Gamma_{\phi \leftrightarrow \chi} \searrow$

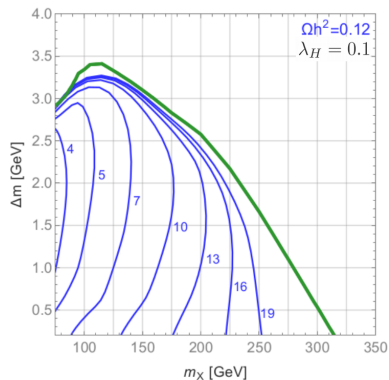
- small mediator-DM splitting $\Delta m/m_\chi \lesssim 0.1$

- Annihilations: $\phi\phi^\dagger \rightarrow \text{SM SM}$
keep ϕ in eq. until χ fo (rate $\propto g^4$)

- **Conversions:** $\phi \rightarrow \ell\chi$ and $\phi\text{SM} \rightarrow \chi\text{SM}$
barely keep ϕ and χ in chemical eq. (rate $\propto \lambda_\chi^2$) $\rightsquigarrow n_\chi(T) > n_\chi^{\text{eq}}(T)$

Conversion driven FO: Viable parameter space

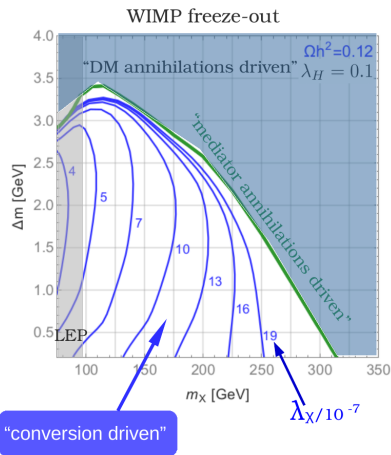
$\lambda_\chi \lesssim \text{few} \times 10^{-6}$, $\Delta m < \text{few GeV}$ and $m_\chi \sim \text{few} \times 100 \text{ GeV}$ for $\lambda_H = 0.1$



Conversion driven FO: Viable parameter space

$\lambda_\chi \lesssim \text{few} \times 10^{-6}$, $\Delta m < \text{few GeV}$ and $m_\chi \sim \text{few} \times 100 \text{ GeV}$ for $\lambda_H = 0.1$

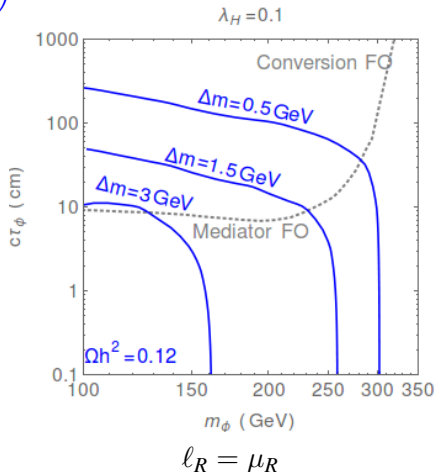
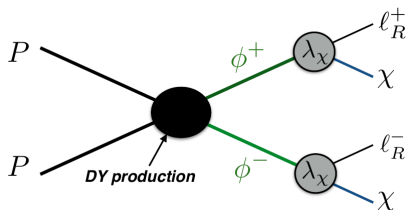
- right limit: freeze-out driven by **mediator annihilations**
- left limit: LEP constraints on charged mediators
- bottom limit: $\Delta m > m_l$ (or beyond $2 \rightarrow 2$)



Collider constraints on charged (disappearing) track

Production of charged ϕ at colliders through EW processes (for small Δm):

$$\Gamma_{\phi \rightarrow \chi l} \approx \frac{1}{25 \text{ cm}} \left(\frac{\lambda_\chi}{10^{-6}} \right)^2 \left(\frac{\Delta m}{1 \text{ GeV}} \right)^2 \left(\frac{100 \text{ GeV}}{m_\chi} \right)$$

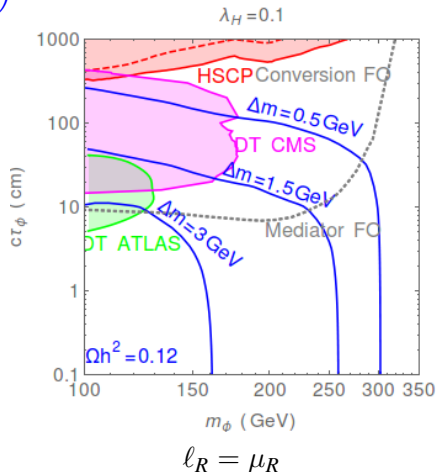


Collider constraints on charged (disappearing) track

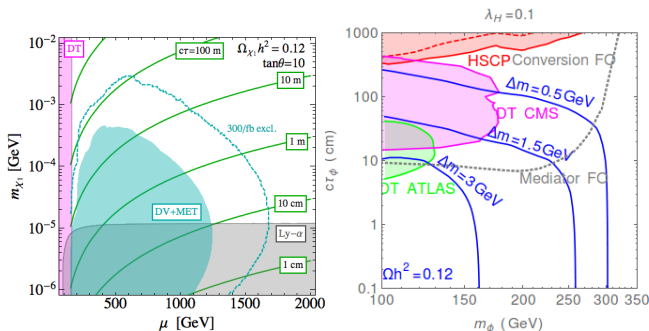
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- **Heavy Stable Charged Particles :**
charged track with
 $c\tau \geq \mathcal{O}(\text{detector size})$
 $\rightsquigarrow c\tau < \text{few m for } m_\phi < 350 \text{ GeV}$
from [CMS HSCP'16] using SModelS
- **Disappearing track:** charged track with
 $c\tau < (\text{detector size})$
exclusion region around
 - $c\tau \sim (5, 50) \text{ cm}, m_\phi < 130 \text{ GeV}$
from [ATLAS DT'18]
 - $c\tau \sim (0.2, 4) \text{ m}, m_\phi < 170 \text{ GeV}$
from [CMS DT'18]



Conclusion



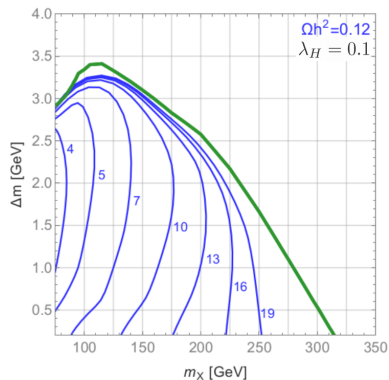
- **H-coupled FIMP** fermionic DM: freeze-in through decays
 - small couplings $y \sim 10^{-8} - 10^{-9}$ long lived mediator at colliders tested with **displaced vertices and disappearing charged track**
 - **complementary constraints from cosmology** with WDM-like behaviour of light FIMP.
- **Leptophilic FIMP** fermionic DM: Conversion driven freeze-out
 - small couplings $y \sim 10^{-6} - 10^{-8}$ but long lived mediator at colliders partially tested with **charged tracks**

Thank you for your attention!

Backup

Conversion driven FO: Viable parameter space

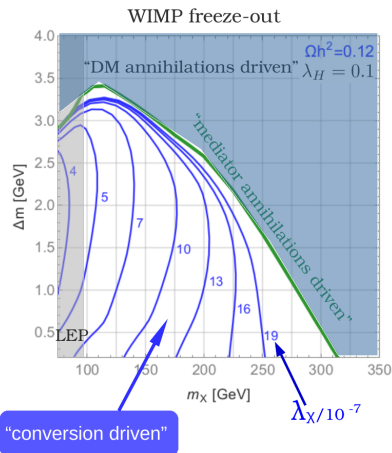
$\lambda_\chi \lesssim \text{few} \times 10^{-6}$, $\Delta m < \text{few GeV}$ and $m_\chi \sim \text{few} \times 100 \text{ GeV}$ for $\lambda_H = 0.1$



Conversion driven FO: Viable parameter space

$$\lambda_\chi \lesssim \text{few} \times 10^{-6}, \Delta m < \text{few GeV} \text{ and } m_\chi \sim \text{few} \times 100 \text{ GeV} \text{ for } \lambda_H = 0.1$$

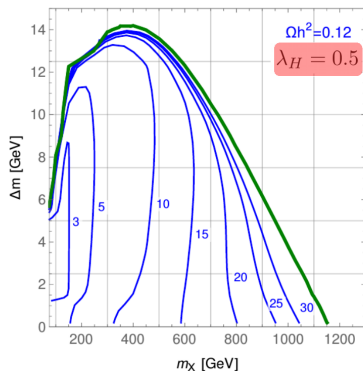
- right limit: freeze-out driven by mediator annihilations
- left limit: LEP constraints on charged mediators
- bottom limit: $\Delta m > m_l$ (or beyond $2 \rightarrow 2$)



Conversion driven FO: Viable parameter space

$\lambda_\chi \lesssim \text{few} \times 10^{-6}$, $\Delta m < \text{few GeV}$ and $m_\chi \sim \text{few} \times 100 \text{ GeV}$ for $\lambda_H = 0.1$

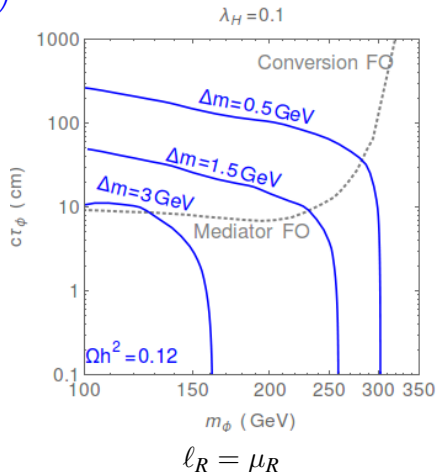
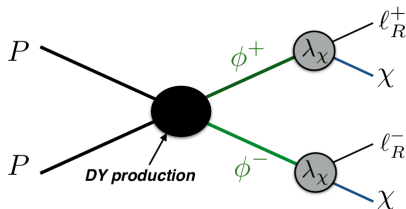
- right limit: freeze-out driven by **mediator annihilations**
- left limit: LEP constraints on **charged mediators**
- bottom limit: $\Delta m > m_l$ (or beyond $2 \rightarrow 2$)
- Larger $\langle \sigma v \rangle_{\phi\phi}$ for **larger λ_H** increase the viable parameter space



Collider constraints on charged (disappearing) track

Production of charged ϕ at colliders through EW processes (for small Δm):

$$\Gamma_{\phi \rightarrow \chi l} \approx \frac{1}{25 \text{ cm}} \left(\frac{\lambda_\chi}{10^{-6}} \right)^2 \left(\frac{\Delta m}{1 \text{ GeV}} \right)^2 \left(\frac{100 \text{ GeV}}{m_\chi} \right)$$

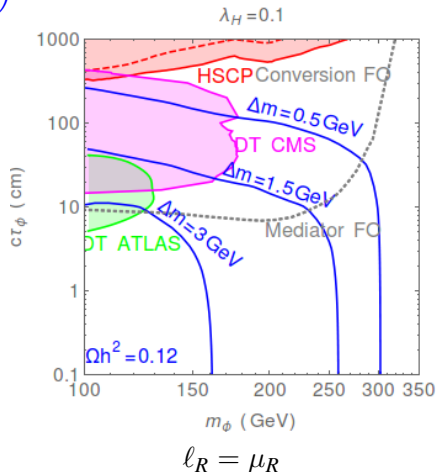


Collider constraints on charged (disappearing) track

Production of charged ϕ at colliders through EW processes (for small Δm):

$$\Gamma_{\phi \rightarrow \chi l} \approx \frac{1}{25 \text{ cm}} \left(\frac{\lambda_\chi}{10^{-6}} \right)^2 \left(\frac{\Delta m}{1 \text{ GeV}} \right)^2 \left(\frac{100 \text{ GeV}}{m_\chi} \right)$$

- **Heavy Stable Charged Particles :**
charged track with
 $c\tau \geq \mathcal{O}(\text{detector size})$
 $\rightsquigarrow c\tau < \text{few m for } m_\phi < 350 \text{ GeV}$
from [CMS HSCP'16] using SModelS
- **Disappearing track:** charged track with
 $c\tau < (\text{detector size})$
exclusion region around
 - $c\tau \sim (5, 50) \text{ cm}, m_\phi < 130 \text{ GeV}$
from [ATLAS DT'18]
 - $c\tau \sim (0.2, 4) \text{ m}, m_\phi < 170 \text{ GeV}$
from [CMS DT'18]

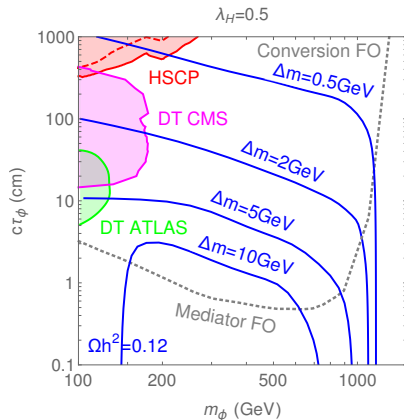


Collider constraints on charged (disappearing) track

Production of charged ϕ at colliders through EW processes (for small Δm):

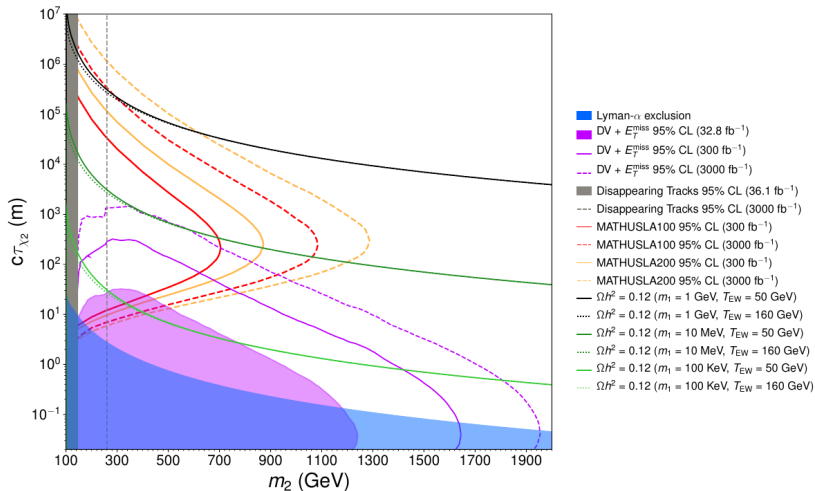
$$\Gamma_{\phi \rightarrow \chi l} \approx \frac{1}{25 \text{ cm}} \left(\frac{\lambda_\chi}{10^{-6}} \right)^2 \left(\frac{\Delta m}{1 \text{ GeV}} \right)^2 \left(\frac{100 \text{ GeV}}{m_\chi} \right)$$

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from [ATLAS DT'18]
 - $c\tau \sim (0.2, 4) \text{ m}, m_\phi < 170 \text{ GeV}$
from [CMS DT'18]



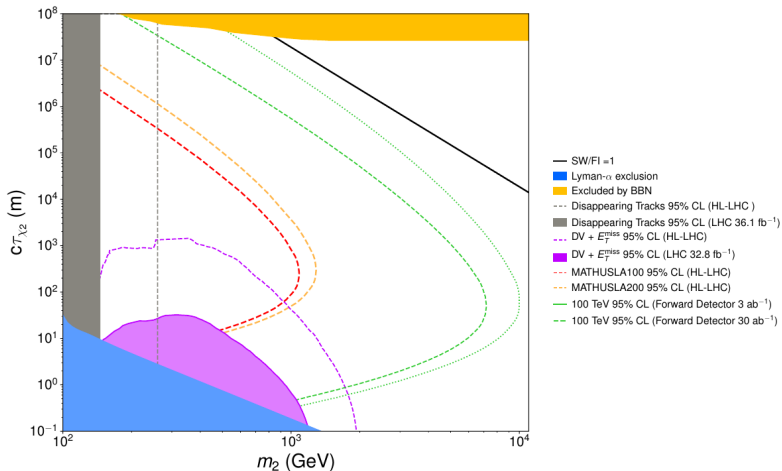
$$\ell_R = \mu_R$$

LHC & Cosmo complementarity



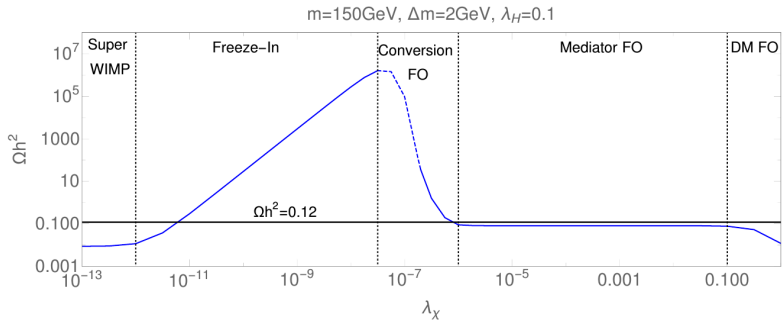
[No, Tunney & Zaldívar'0919]

LHC & Cosmo complementarity



[No'19]

From SuperWIMP to freeze-out



Relative scattering contribution to FI

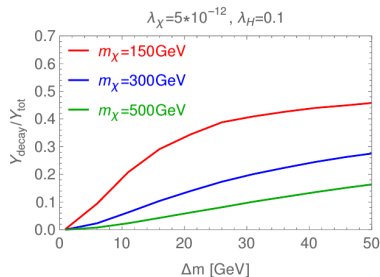


Figure 4: The ratio of the freeze-in yields $Y_{\text{decay}}^{fi}/Y_{\text{tot}}^{fi}$ as a function of Δm when taking only decay processes into account for Y_{decay}^{fi} and including all the relevant scattering and decay processes for Y_{tot}^{fi} . We show the results for three different values of the DM mass $m_\chi \in \{150, 300, 500\}$ GeV while keeping $\lambda_\chi = 5 \cdot 10^{-12}$ and $\lambda_H = 0.1$ fixed.

Displaced lepton searches inappropriate

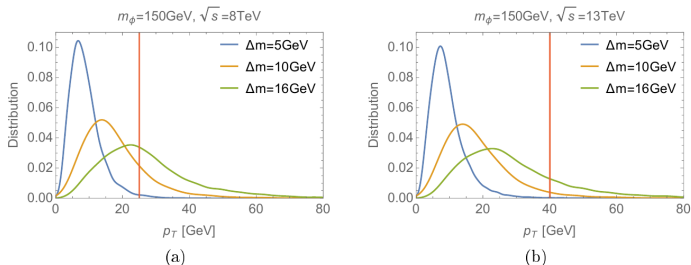
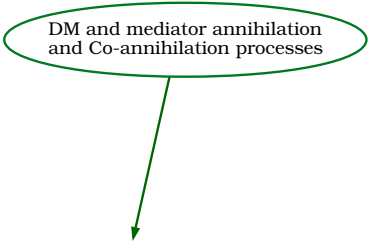


Figure 7: The p_T -distribution of the leading muons for the case when $m_\phi = 150 \text{ GeV}$ and different values of the mass splitting. The red vertical line denotes the cut on the p_T that is made in the search [67, 68] at 25 GeV for $\sqrt{s} = 8 \text{ TeV}$ (left) and 40 GeV for $\sqrt{s} = 13 \text{ TeV}$ (right).

In the small mass splitting case, most events do not pass current p_T cuts

Conversion processes can affect the DM abundance

DM and mediator annihilation
and Co-annihilation processes



$$\frac{dY_{DM}}{dx} = \frac{s\langle\sigma v_{\text{eff}}\rangle}{Hx} (Y_{DM}^2 - Y_{DM,eq}^2)$$

Conversion processes can affect the DM abundance

DM and mediator annihilation
and Co-annihilation processes

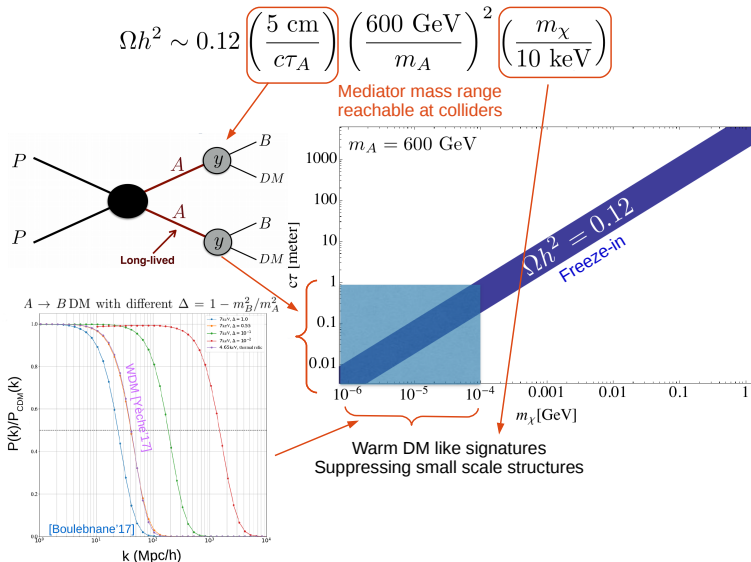
$$\frac{dY_\chi}{dx} = \frac{-2}{Hxs} \left[\gamma_{\chi\chi} \left(\frac{Y_\chi^2}{Y_{\chi,eq}^2} - 1 \right) + \gamma_{\chi\phi} \left(\frac{Y_\chi Y_\phi}{Y_{\chi,eq} Y_{\phi,eq}} - 1 \right) \right. \\ \left. + \gamma_{\chi \rightarrow \phi} \left(\frac{Y_\chi}{Y_{\chi,eq}} - \frac{Y_\phi}{Y_{\phi,eq}} \right) + \gamma_{\chi\chi \rightarrow \phi\phi^\dagger} \left(\frac{Y_\chi^2}{Y_{\chi,eq}^2} - \frac{Y_\phi^2}{Y_{\phi,eq}^2} \right) \right],$$

$$\frac{dY_\phi}{dx} = \frac{-2}{Hxs} \left[\gamma_{\phi\phi^\dagger} \left(\frac{Y_\phi^2}{Y_{\phi,eq}^2} - 1 \right) + \gamma_{\chi\phi} \left(\frac{Y_\chi Y_\phi}{Y_{\chi,eq} Y_{\phi,eq}} - 1 \right) \right. \\ \left. - \gamma_{\chi \rightarrow \phi} \left(\frac{Y_\chi}{Y_{\chi,eq}} - \frac{Y_\phi}{Y_{\phi,eq}} \right) - \gamma_{\chi\chi \rightarrow \phi\phi^\dagger} \left(\frac{Y_\chi^2}{Y_{\chi,eq}^2} - \frac{Y_\phi^2}{Y_{\phi,eq}^2} \right) \right],$$

DM and mediator
Conversion processes

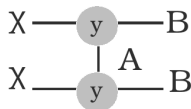
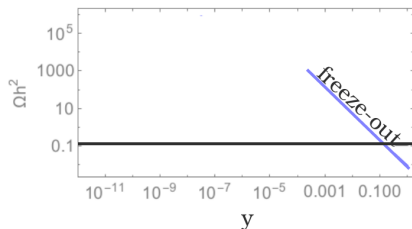
FIMP: displaced vertices and cosmology interplay

e.g. [Hall'09, Co'15, d'Eramo'17, Heeck'17, Boulebnane'17, Brooijmans'18, Garny'18, Calibbi'18, Belanger'18 ...]



Freeze-in versus Freeze-out: feeble vs weak

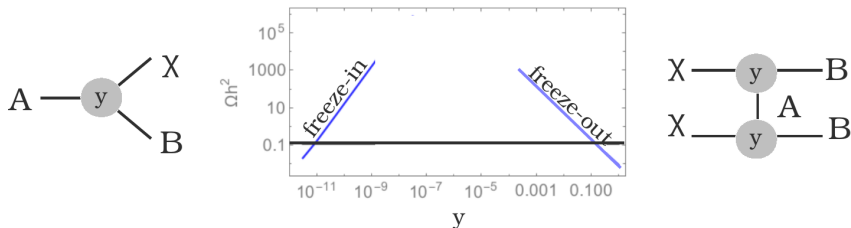
Assuming $\mathcal{L} \subset y\chi AB$ and $m_A \gg m_\chi$:



- **Freeze-out** ($y \gtrsim 0.01$):
 - $Y_{DM} \propto 1/\langle\sigma v\rangle_{\chi\chi\rightarrow BB} \sim 1/y^4$
 - A role is irrelevant
 - $Y_{DM}(t_i) = Y_{DM}^{eq}$
independent of init. cond.

Freeze-in versus Freeze-out: feeble vs weak

Assuming $\mathcal{L} \subset y\chi AB$ and $m_A \gg m_\chi$:



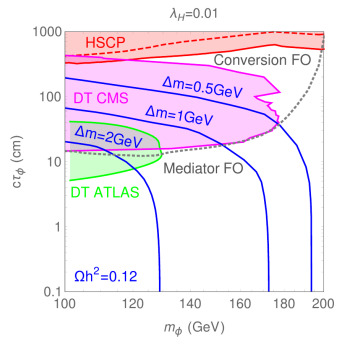
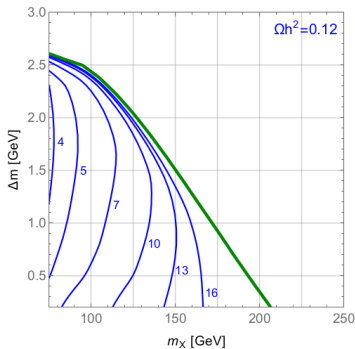
• **Freeze-in** ($y \lesssim 10^{-8}$):

- $Y_{DM} \propto \Gamma_{A \rightarrow \chi B} \sim y^2$
- A plays the leading role
- We assume $Y_{DM}(t_i) = 0$
(dependence in init. cond.)

• **Freeze-out** ($y \gtrsim 0.01$):

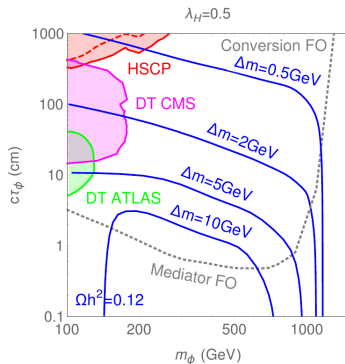
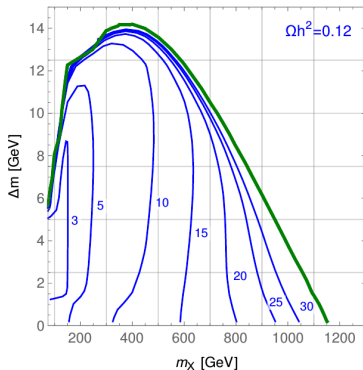
- $Y_{DM} \propto 1/\langle \sigma v \rangle_{\chi\chi \rightarrow BB} \sim 1/y^4$
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- $Y_{DM}(t_i) = Y_{DM}^{eq}$
independent of init. cond.

Conversion driven freeze-out

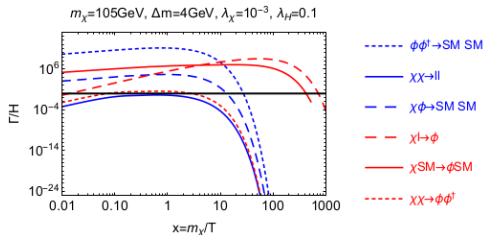


$$\lambda_H = 0.01$$

Conversion driven freeze-out



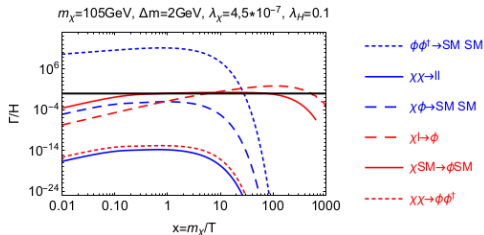
$$\lambda_H = 0.5$$



$$\frac{dY_{DM}}{dx} = \frac{s\langle\sigma v_{\text{eff}}\rangle}{Hx} (Y_{DM}^2 - Y_{DM,eq}^2)$$

$$\langle\sigma v_{\text{eff}}\rangle \simeq \frac{1}{g_{\text{eff}}^2} \sum_{ij} r_i r_j \langle\sigma v_{ij}\rangle \quad g_{\text{eff}} = \sum_i g_i$$

and $r_i = g_i(1 + \Delta_i)^{3/2} \exp(-x_f \Delta_i)$

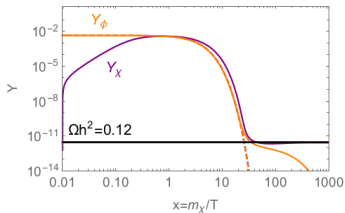


$$\frac{dY_\chi}{dx} = \frac{-2}{Hxs} \left[\gamma_{\chi\chi} \left(\frac{Y_\chi^2}{Y_{\chi,eq}^2} - 1 \right) + \gamma_{\chi\phi} \left(\frac{Y_\chi Y_\phi}{Y_{\chi,eq} Y_{\phi,eq}} - 1 \right) + \gamma_{\chi \rightarrow \phi} \left(\frac{Y_\chi}{Y_{\chi,eq}} - \frac{Y_\phi}{Y_{\phi,eq}} \right) + \gamma_{\chi\chi \rightarrow \phi\phi^\dagger} \left(\frac{Y_\chi^2}{Y_{\chi,eq}^2} - \frac{Y_\phi^2}{Y_{\phi,eq}^2} \right) \right],$$

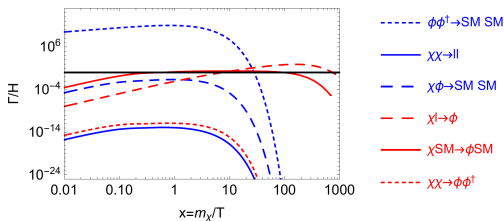
$$\frac{dY_\phi}{dx} = \frac{-2}{Hxs} \left[\gamma_{\phi\phi^\dagger} \left(\frac{Y_\phi^2}{Y_{\phi,eq}^2} - 1 \right) + \gamma_{\chi\phi} \left(\frac{Y_\chi Y_\phi}{Y_{\chi,eq} Y_{\phi,eq}} - 1 \right) - \gamma_{\chi \rightarrow \phi} \left(\frac{Y_\chi}{Y_{\chi,eq}} - \frac{Y_\phi}{Y_{\phi,eq}} \right) - \gamma_{\chi\chi \rightarrow \phi\phi^\dagger} \left(\frac{Y_\chi^2}{Y_{\chi,eq}^2} - \frac{Y_\phi^2}{Y_{\phi,eq}^2} \right) \right],$$

Conversion driven

$$m_\chi=150\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=8\cdot 10^{-7}, \lambda_H=0.1$$

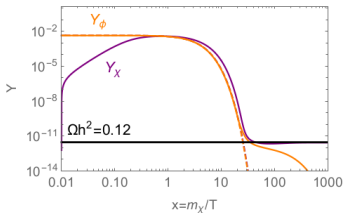


$$m_\chi=150\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=8\cdot 10^{-7}, \lambda_H=0.1$$



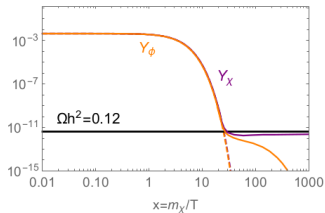
Conversion driven

$$m_\chi=150\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=8\cdot 10^{-7}, \lambda_H=0.1$$

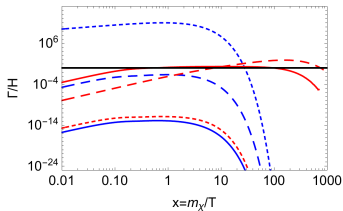


Co-annihilation driven

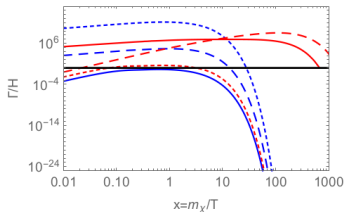
$$m_\chi=105\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=10^{-3}, \lambda_H=0.1$$



$$m_\chi=150\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=8\cdot 10^{-7}, \lambda_H=0.1$$

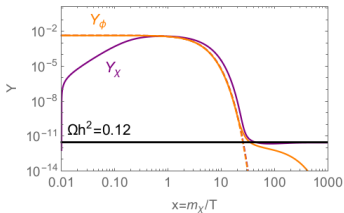


$$m_\chi=105\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=10^{-3}, \lambda_H=0.1$$



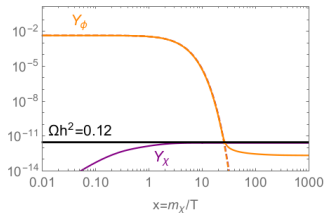
Conversion driven

$$m_\chi=150\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=8\cdot 10^{-7}, \lambda_H=0.1$$

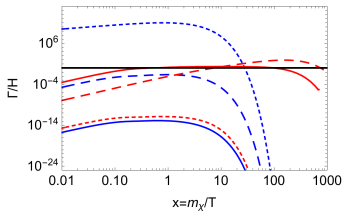


Freeze-in

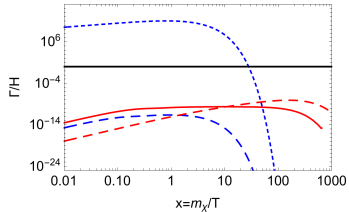
$$m_\chi=150\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=8\cdot 10^{-12}, \lambda_H=0.1$$



$$m_\chi=150\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=8\cdot 10^{-7}, \lambda_H=0.1$$



$$m_\chi=150\text{GeV}, \Delta m=2\text{GeV}, \lambda_\chi=8\cdot 10^{-12}, \lambda_H=0.1$$



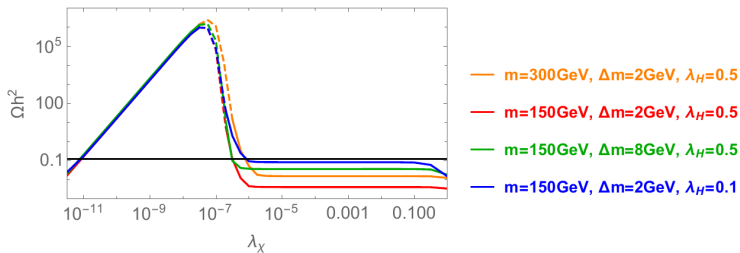


Figure 8: DM abundance as a function of the Yukawa coupling (as in Fig. 2) for different values of the parameters m_χ , Δm and λ_H

Relevant processes in Leptophilic DM

(co-)annihilations

initial state		final state		scaling
χ	χ	l^-	l^+	λ_χ^4
χ	ϕ	l^-	γ, Z, H	λ_χ^2
		W^-	ν_l	
ϕ	$\phi^\dagger$	γ, Z, W^+	γ, Z, W^-	λ_χ^0
		q	$\bar{q}$	
		H	Z	
		l^-	l^+	
ϕ	ϕ	l^-	l^-	λ_χ^4

conversions

initial state		final state		scaling
χ	l^-	ϕ	γ, Z, H	λ_χ^2
	γ, Z, H		l^+	
	W^-		$\bar{\nu}_l$	
	ν_l		W^+	
ϕ		χ	l^-	λ_χ^2
χ	χ	ϕ	$\phi^\dagger$	λ_χ^4

Conversion driven freeze-out coupling to q

$$\mathcal{L}_{\text{int}} = |D_\mu \tilde{q}|^2 - \lambda_\chi \tilde{q} \tilde{q} \frac{1 - \gamma_5}{2} \chi + \text{h.c.}$$

[Heisig garmy et al '17 & '18]

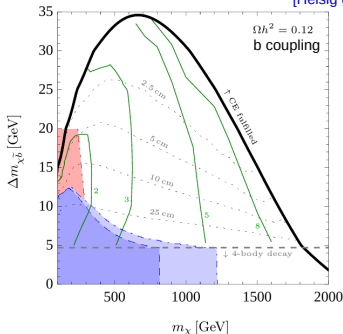


FIG. 4. Viable parameter space in the plane spanned by m_χ and $\Delta m_{\chi_b} = m_{\tilde{b}} - m_\chi$. We adjust λ_χ such that $\Omega h^2 = 0.12$. Above the thick black curve CE holds, while below this curve CE breaks down and the freeze-out is conversion-driven. The corresponding coupling $\lambda_\chi/10^{-7}$ (decay length $c\tau$) of the mediator is denoted by the thin green (gray) dotted lines. The blue dashed (dot-dashed) curve shows our estimates for the limits from R -hadron searches at 8 (13) TeV. The constraint from monojet searches is shown as the red dot-dot-dashed curve.

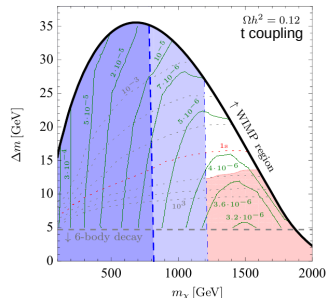
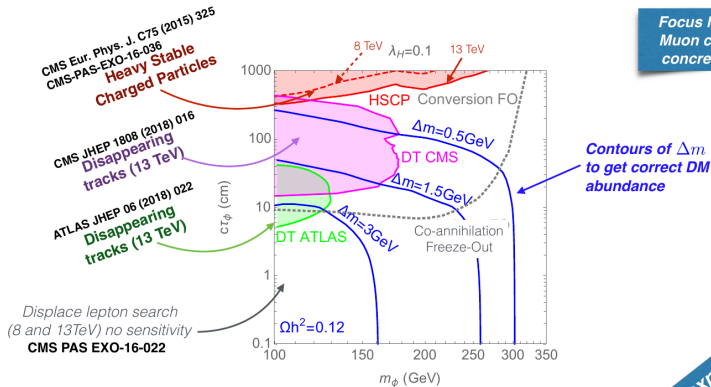


FIG. 9. Cosmologically viable parameter space ($\Omega h^2 = 0.12$) in the conversion-driven freeze-out region (below black thick curve). Contours of constant λ_χ are shown in green while contours of constant τ lifetimes are shown as gray dotted curves spanning from 10^{-3} s to 10^{-6} s in steps of an order of magnitude (the curve for 1 s is highlighted in red for better readability). 95% C.L. exclusion regions from R -hadron searches at the 8 and 13 TeV LHC are shown in dark and light blue, respectively. The red shaded region denotes the constraint from BBN. Below the horizontal gray dashed line (~ 5 GeV) the 4-body decay of the mediator is kinematically forbidden rendering the 6-body decay to be dominant.

Conversion driven freeze-out: Collider probes



Focus here on
Muon case for
concreteness

Contours of Δm
to get correct DM
abundance

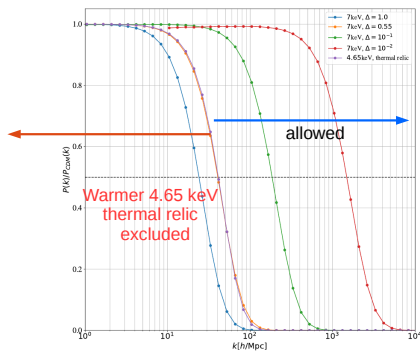
- ✦ Weak constraints from LHC
- ✦ Large regions unconstrained
- ✦ No signal expected in direct/indirect detection

Still much to explore!

WDM constraints on FIMPs

See Heeck et al '1706 and '1709

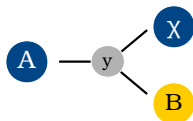
- **WDM: free-streaming** (collision-less damping): collisionless particles can stream out of overdense to underdense regions



smooth out inhomogeneities for $\lambda < \lambda_{FS} = \int_0^{t_0} \frac{v}{a} dt$
 $\rightsquigarrow$ particles relativistic at the time of decoupling can
 give substantial λ_{FS}

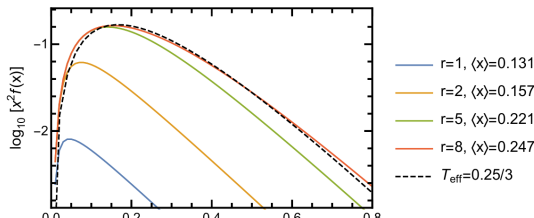
WDM constraints on FIMPs

See Heeck et al '1706 and '1709

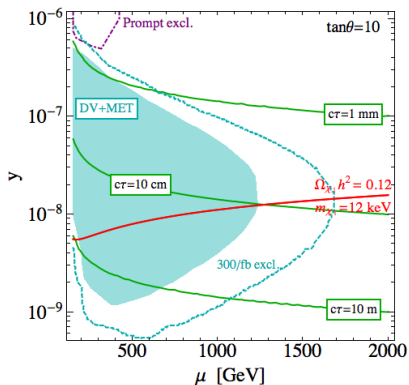


$$\frac{\partial f_{\text{DM}}}{\partial t} - H p \frac{\partial f_{\text{DM}}}{\partial p} = \mathcal{C}(p)$$

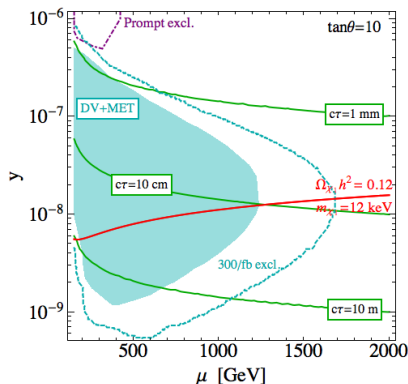
$$\frac{\partial f_{\text{DM}}}{\partial r} = \frac{g_A S \Gamma M_0 r^2}{2 p_{\text{DM}} x \sqrt{m_A^2 x^2 + m_{\text{DM}}^2 r^2}} \int_{\xi^-}^{\xi^+} d\xi f_A(\xi)$$



LHC & Cosmology complementarity



LHC & Cosmology complementarity



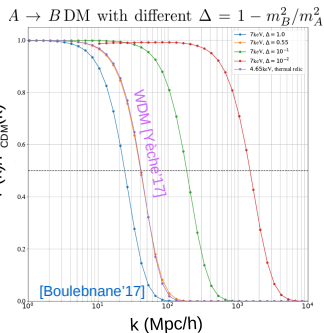
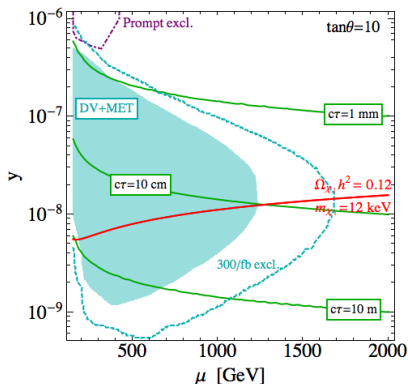
Fixed DM mass: **overabundant**
 Fixed DM abundance: **need lower mass**

Fixed DM mass: **underabundant**
 Fixed DM abundance: **need higher mass**

LHC & Cosmology complementarity

Light FIMP can behave as WDM:

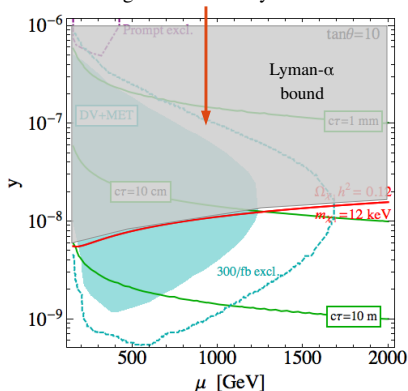
Freestreams from overdense to underdense regions in the early universe



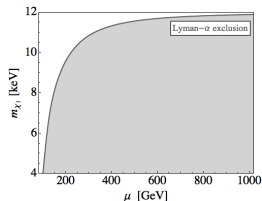
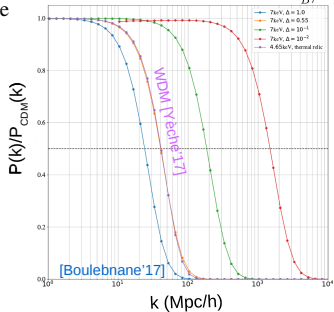
LHC & Cosmology complementarity

Light FIMP can behave as WDM:

Freestreams from overdense to underdense regions in the early universe



$A \rightarrow B$ DM with different $\Delta = 1 - m_B^2/m_A^2$



Displaced vertices recasting

Calibbi, Lopez-Honorez, Lowette, AM

Recasting ATLAS DV+MET

ATLAS arXiv: 1710.04901
CERN-EP-2017-202

- Follow object selection of auxiliary materials

E_T^{truth} , d_0 , n_{tracks}^{DV} , m_{DV} , R_{decay} , z_{decay}

- Apply the efficiency grids

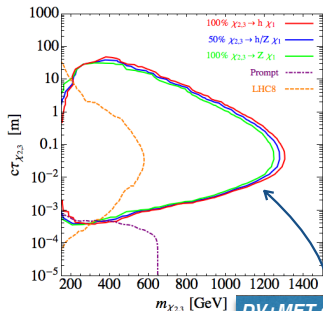
- Validate recasting with model in ATLAS paper as advocated in Les Houches 2017

G. Cottin, N. Desai, J. Heisig, A. Lessa

See Nishita Desai
talk of yesterday

See Hideyuki Oide
talk of yesterday

Simplified models with
fixed BR into $h(+\text{MET})$ or $Z(+\text{MET})$



DV+MET search
has strong
reach on EW
states in SD-FI

Displaced vertices recasting

Recasting a DV+MET search by ATLAS

EUROPEAN ORGANISATION FOR NUCLEAR RESEARCH (CERN)



Submitted to: Phys. Rev. D.



CERN-EP-2017-202
October 16, 2017

Search for long-lived, massive particles in events with displaced vertices and missing transverse momentum in $\sqrt{s} = 13$ TeV pp collisions with the ATLAS detector

The ATLAS Collaboration

A search for long-lived, massive particles predicted by many theories beyond the Standard Model is presented. The search targets final states with large missing transverse momentum and at least one high-mass displaced vertex with five or more tracks, and uses 32.8 fb⁻¹ of $\sqrt{s} = 13$ TeV pp collision data collected by the ATLAS detector at the LHC. The observed yield is consistent with the expected background. The results are used to extract 95% CL exclusion limits on the production of long-lived gluinos with masses up to 2.37 TeV and lifetimes of $O(10^{-2})$ – $O(10)$ ns in a simplified model inspired by Split Supersymmetry.

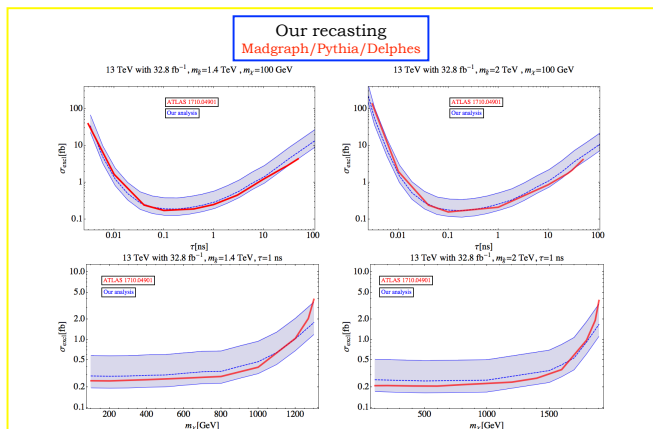
arXiv:1710.04901

Freeze-in DM at the LHC

Lorenzo Calibbi (ITP)

Displaced vertices recasting

Recasting a DV+MET search by ATLAS

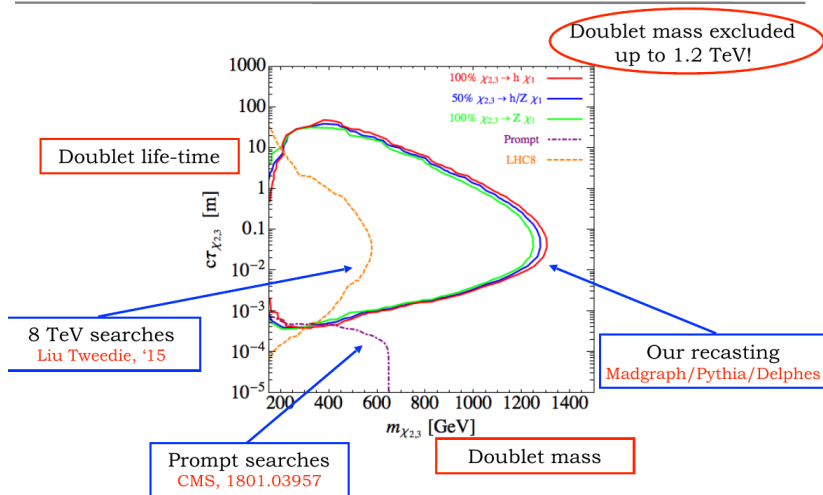


Freeze-in DM at the LHC

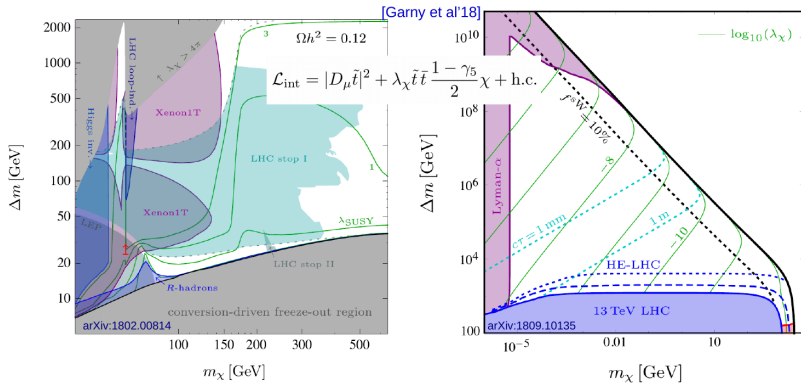
Lorenzo Calibbi (ITP)

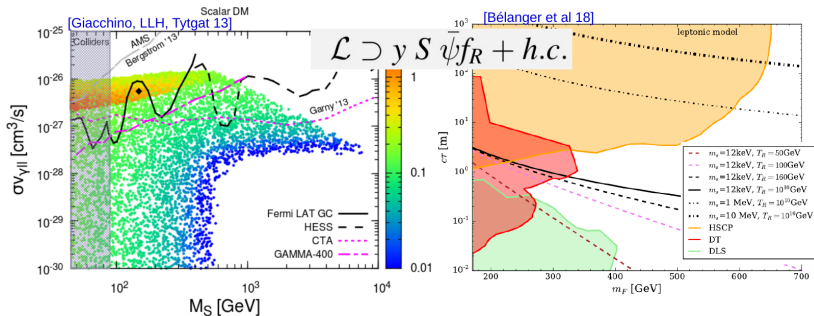
Displaced vertices recasting

Recasting a DV+MET search by ATLAS



Rather general result: it also applies e.g. to Higgsino decaying to gravitino





bla

This is really the end