

Boltzmann equation for relativistic species and Hot Dark Matter

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Scalars 2017

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Motivation

Hubble constant measurement discrepancy:

Planck CMB data: $67.8 \pm 0.9 \text{ km s}^{-1} \text{ Mpc}^{-1}$ [arXiv:1502.01589]

Direct measurement: $73.24 \pm 1.74 \text{ km s}^{-1} \text{ Mpc}^{-1}$ [arXiv:1604.01424]

Possible solution: additional effective number of relativistic degrees of freedom ΔN_{eff} :

$$\rho_{\text{HDM}} = \Delta N_{\text{eff}} \frac{7}{8} \left(\frac{4}{11} \right)^{4/3} \rho_{\gamma}$$

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Therefore, it is crucial to estimate ΔN_{eff} with better accuracy in models with HDM component

Boltzmann equation for relativistic particles

Dolgov & Kainulainen, Nucl. Phys. B 402 (1993) 349 [hep-ph/9211231]

$$E(\partial_t - pH\partial_p)f(p, t) = C_E(p, t) + C_I(p, t)$$

Pseudopotential method ($\bar{\psi}\psi \rightarrow \bar{N}N$, $x = \frac{m}{T}$, $y = \frac{E}{T}$, z):

$$f(p, t) \simeq (e^{xy+z} \mp 1)^{-1}$$

$$f_N(p, t) \simeq (e^{xy_N} \mp 1)^{-1}$$

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where:

$$A(z, x) = \frac{g}{2\pi^2} m^3 e^z J_2(z, x)$$

$$B(z, x) \equiv \frac{g}{2\pi^2} m^3 e^z x J_3(z, x) \left(H(T) + \frac{1}{T} \frac{dT}{dt} \right)$$

$$J_n(z, x) \equiv \int_0^\infty dy y^n \frac{e^{xy}}{(1 \mp e^{xy+z})^2}$$

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$$S_I(z, x) = \frac{m^4}{512\pi^6} (e^{2z} - 1) \int \mathcal{D}\Phi \int_0^{2\pi} d\phi \sum_{\text{spin}} |M(u, v, t, \cos\theta, \phi)|^2$$

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For $|M|^2 = |M(s)|^2$:

$$\begin{aligned} \mathcal{D}\Phi &= \frac{1}{x^4} \int_{2x}^{\infty} dp \int_0^{\sqrt{p^2 - 4x^2}} dq \frac{1}{(1 - e^{-p})(e^{p+2z} - 1)} \\ &\times \ln \left[\frac{\cosh(\frac{1}{2}(p+q) + z) \mp 1}{\cosh(\frac{1}{2}(p-q) + z) \mp 1} \right] \ln \left[\frac{\cosh(\frac{1}{2}(p+qV(p, q)) + z) \mp 1}{\cosh(\frac{1}{2}(p-qV(p, q)) + z) \mp 1} \right] \end{aligned}$$

$$\text{where } V = \left(1 - \frac{4x^2}{p^2 - q^2}\right)^{1/2}, s = m^2 \frac{p^2 - q^2}{x^2}$$

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Finally:

$$n(x) \equiv n_\psi(x) = g \frac{m^3}{2\pi^2} \int_0^\infty \frac{y^2 dy}{e^{xy+z} \mp 1}$$

Approximations

$$Y'(x) = -\sqrt{\frac{\pi}{45}} \frac{g(x)}{\sqrt{g_s(x)}} \frac{M_{\text{Pl}} m}{x^2} \left(Y^2(x) - Y_{\text{eq}}^2(x) \right) \langle \sigma v \rangle_{\text{MB}} \frac{1}{\zeta^{\mp}}$$

$$Y_{\text{eq}}(x) = g \frac{45}{2\pi^4} \frac{\zeta^{\mp}}{g_s(x)} \quad \zeta^{\mp} \equiv \zeta(3) \begin{cases} 1 & \text{BE } (-) \\ 3/4 & \text{FD } (+) \end{cases}$$

$$\langle \sigma v \rangle_{\text{MB}} = \frac{1}{512\pi} \frac{x^5}{m^5} \int_{4m^2}^{\infty} ds \sqrt{s} \sqrt{1 - \frac{4m^2}{s}} |M(s)|^2 K_1 \left(\frac{x\sqrt{s}}{m} \right)$$

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- Pure MB: $\langle \sigma v \rangle_{\text{MB}}$ with $\zeta^{\mp} = 1$ artificial MB
- Fractional BE/FD: $\langle \sigma v \rangle_{\text{MB}} \times \zeta^{\mp}$ used fBE/FD
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Weinberg's Higgs portal model

S. Weinberg, Phys. Rev. Lett. **110** (2013) no.24, 241301 [arXiv:1305.1971]

$$\begin{aligned}\mathcal{L}_{H,\phi} = & (D_\mu H)^\dagger (D^\mu H) + \mu_H^2 H^\dagger H - \lambda_H (H^\dagger H)^2 \\ & + \partial_\mu \phi^* \partial^\mu \phi + \mu_\phi (\phi^* \phi)^2 - \lambda_\phi (\phi^* \phi)^2 + \kappa (H^\dagger H)(\phi^* \phi) \\ H_0 = & v_H + \frac{\tilde{h} + iG^0}{\sqrt{2}} \quad \phi = v_\phi + \frac{\rho + i\sigma}{\sqrt{2}}\end{aligned}$$

Diagonalization from $(\tilde{h} \tilde{\rho})$ to $(h \rho)$: $\tan 2\theta = \frac{\kappa v_H v_\phi}{\lambda_H v_H^2 - \lambda_\phi v_\phi^2}$

Convenient set of independent parameters: $m_\rho, \kappa, \lambda_\phi$

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Let's consider HDM annihilation into muons: $\sigma\sigma \rightarrow \bar{\mu}\mu$.

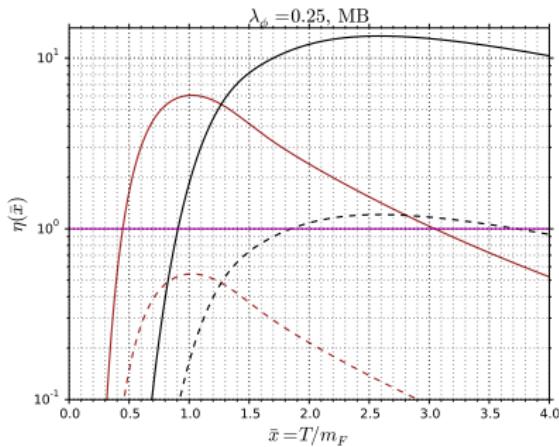
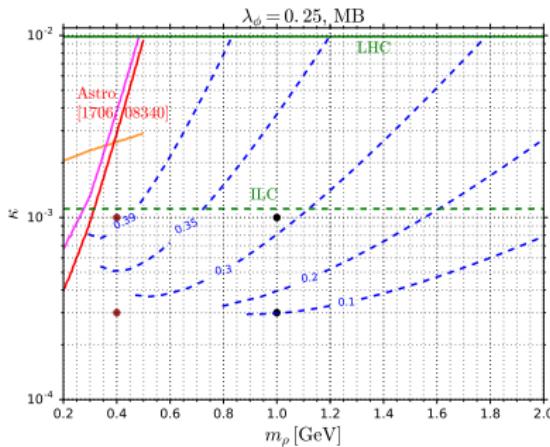
In narrow resonance approximation $(\Gamma_\rho/m_\rho)^2 \ll 1$:

$$|M|^2 = \frac{2\pi\kappa^2}{m_\rho\Gamma_\rho} (m_\rho^2 - 4m_F^2) \frac{m_\rho^4 + \frac{m_\rho^2 m_h^2}{(m_\rho^2 - m_h^2)^2} (m_\rho\Gamma_h - m_h\Gamma_\rho)^2}{[(m_\rho^2 - m_h^2)^2 + \Gamma_h^2 m_h^2]}$$

Freeze-out approximation

$$\eta(x) = \frac{\Gamma}{H} = \left. \frac{n\langle\sigma v\rangle}{H} \right|_{x=x_f} = 1$$

$$\eta(x) \approx \frac{\kappa^2}{\lambda_\phi} \frac{\sqrt{45} \zeta(3)}{32\pi^{5/2}} \frac{m_\rho^5 M_{\text{Pl}}}{m_h^4 m_F^2} \left(1 - \frac{4m_F^2}{m_\rho^2}\right)^{3/2} \frac{x^4}{\sqrt{g(x)}} K_1\left(\frac{x m_\rho}{m_F}\right) \Big|_{x=x_f} = 1$$



Comparison of different statistics & approximations

$$\mathcal{D}\Phi_{\text{fBE}} \approx \mathcal{D}\Phi_{\text{MB}} \times \zeta^{\mp}$$

$$\mathcal{D}\Phi_{\text{fFD}} \approx \mathcal{D}\Phi_{\text{MB}} \times \zeta^{\mp}$$

$$\mathcal{D}\Phi_{\text{pBE}} \approx \mathcal{D}\Phi_{\text{MB}} \times \coth\left(\frac{m_\rho x}{4m_F}\right)$$

$$\mathcal{D}\Phi_{\text{pFD}} \approx \mathcal{D}\Phi_{\text{MB}} \times \tanh\left(\frac{m_\rho x}{4m_F}\right)$$

$$\mathcal{D}\Phi_{\text{BEFD}} \approx \mathcal{D}\Phi_{\text{MB}}$$

$$\mathcal{D}\Phi_{\text{FDBE}} \approx \mathcal{D}\Phi_{\text{MB}}$$

$$\mathcal{D}\Phi_{\text{BEBE}} \approx \mathcal{D}\Phi_{\text{MB}} \times \coth^2\left(\frac{m_\rho x}{4m_F}\right)$$

$$\mathcal{D}\Phi_{\text{FDFD}} \approx \mathcal{D}\Phi_{\text{MB}} \times \tanh^2\left(\frac{m_\rho x}{4m_F}\right)$$

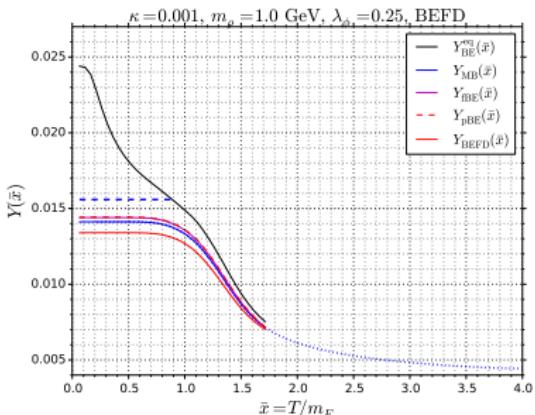
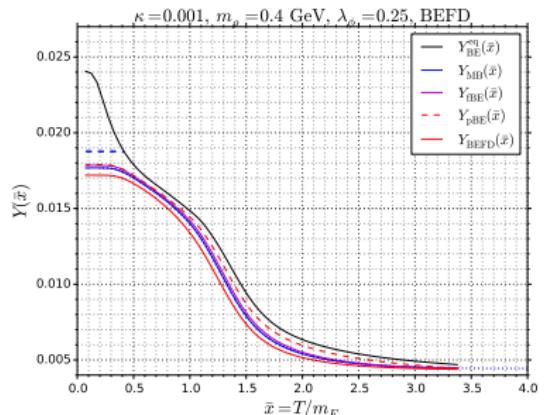
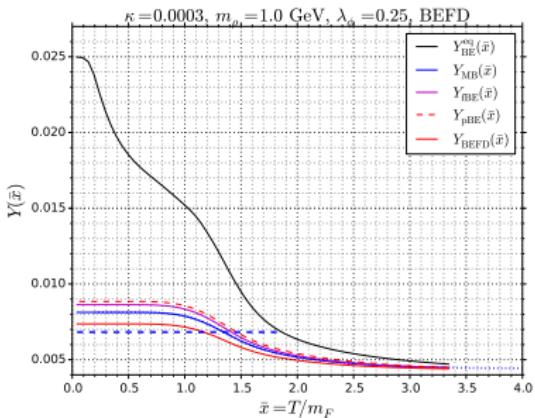
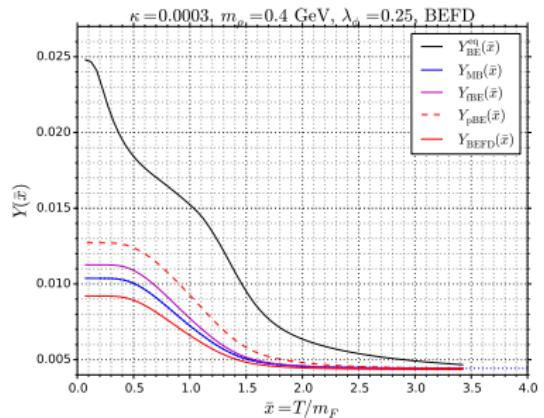
where

$$\mathcal{D}\Phi_{\text{MB}} = e^{-2x} \frac{|M|^2}{2x} \sqrt{\frac{m_\rho^2}{m_F^2} - 4} K_1\left(\frac{x m_\rho}{m_F}\right)$$

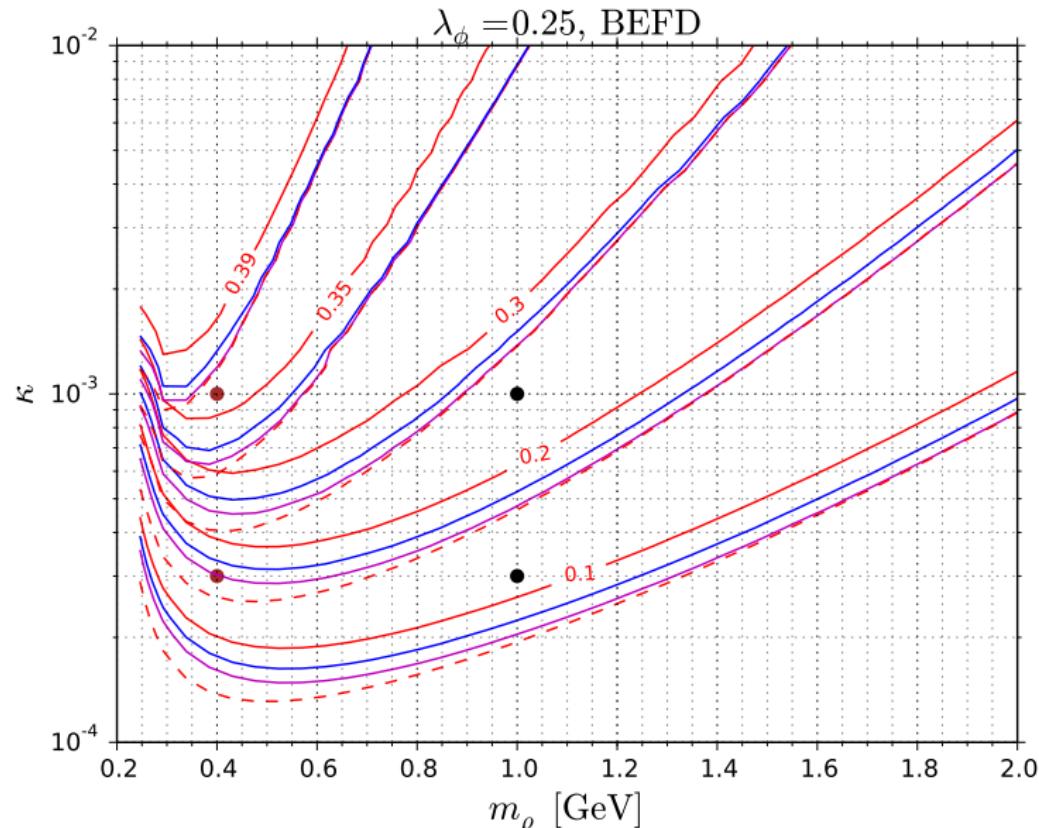
- ▶ Cancellation for mixed statistics
- ▶ Amplification for BE
- ▶ Suppression for FD

Example: $m_\rho = 0.4$ GeV

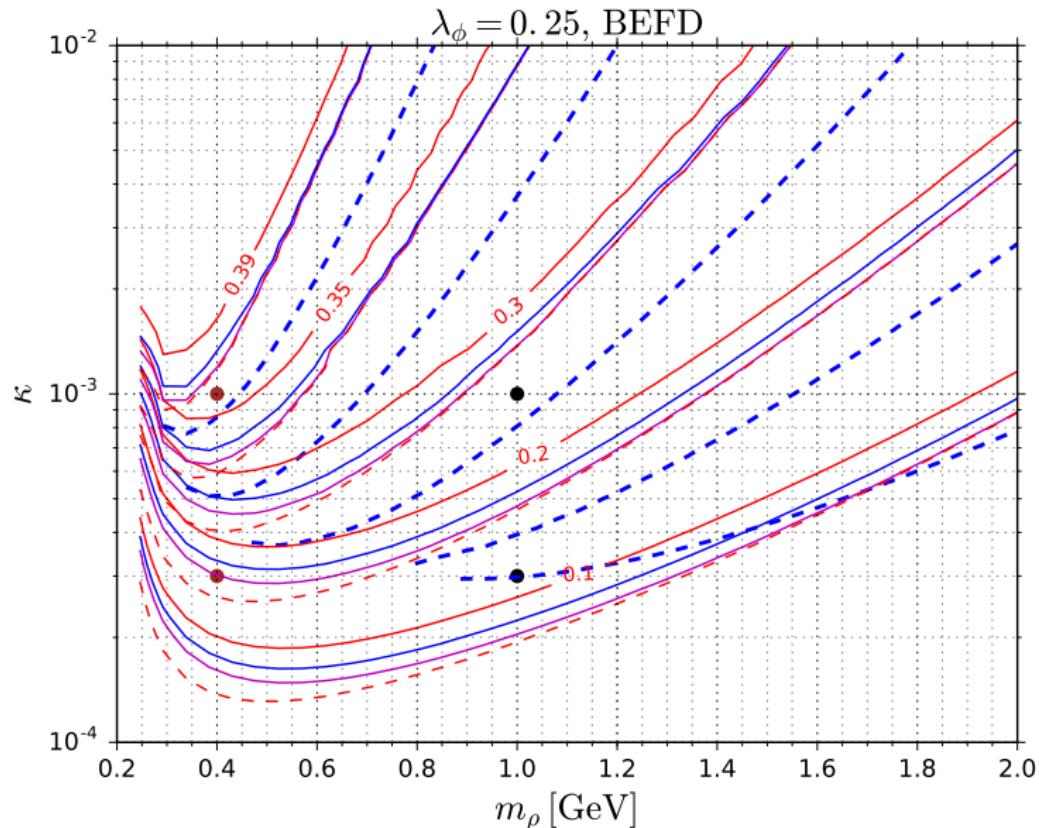
$m_\rho = 1$ GeV



Numerical scan



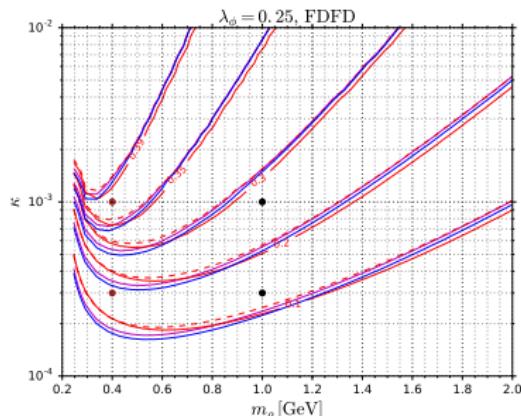
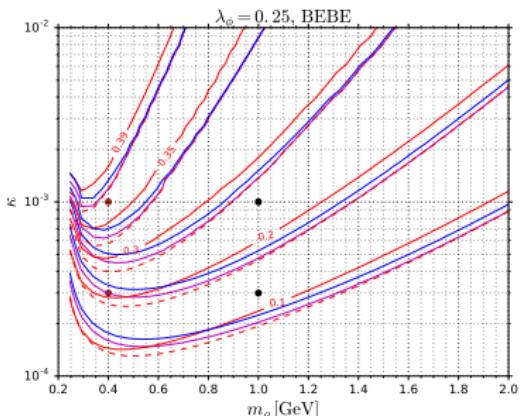
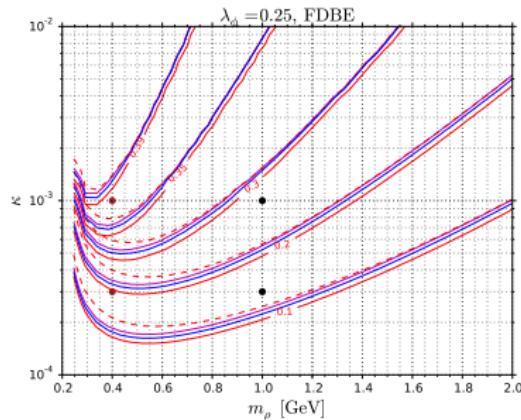
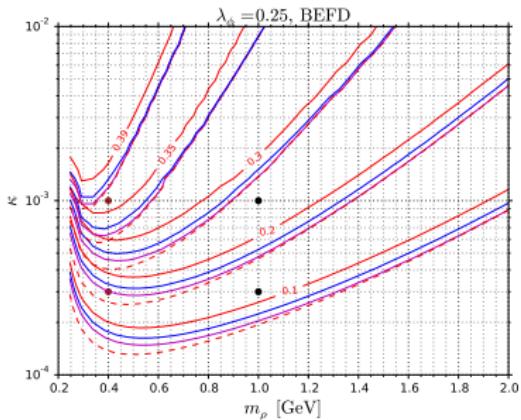
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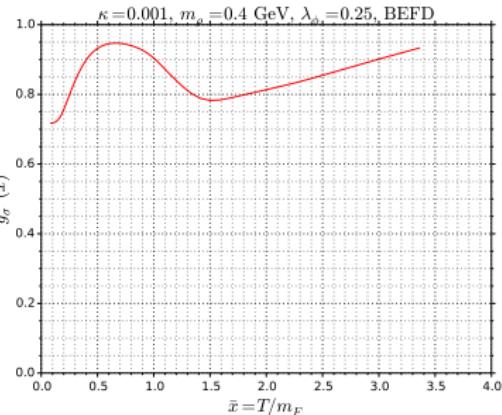
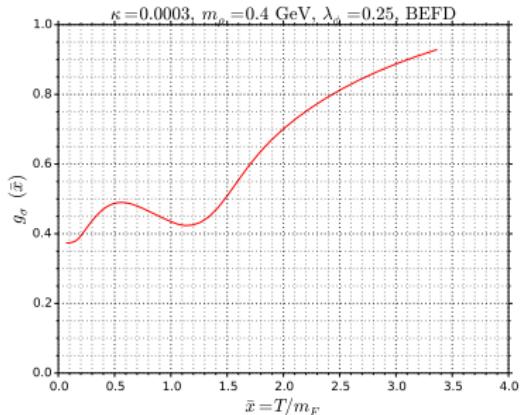
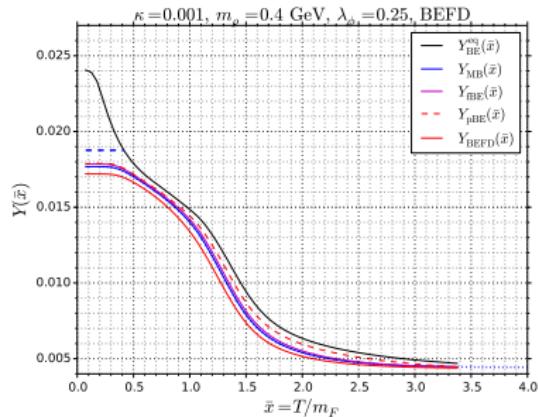
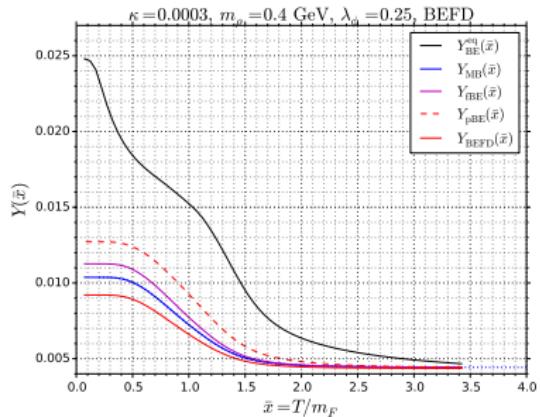
Conclusions

- ▶ We have compared a few approaches for estimating ΔN_{eff} of the Goldstone boson in Weinberg's model
- ▶ Freeze-out approximation breaks down for small Higgs portal coupling κ
- ▶ Pure Maxwell-Boltzmann approximation works better than others because of cancellation effect between mixed statistics
- ▶ Both the statistics of incoming and outgoing particles influence the results
- ▶ In order to obtain higher accuracy Boltzmann equation with full statistics should be considered
- ▶ Work in progress

Backup slides



Backup slides



Backup slides

