

Baryogenesis from L-violating Higgs doublet decay

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Work in collaboration with Daniele Teresi (ULB): arXiv:1606.00017 (PRL 2016)
arXiv:1705.00016 (PRD 2017)

Baryogenesis via leptogenesis

 very highly motivated: same origin as neutrino masses



very natural at high scale: a series of numerical coincidences which make it particularly efficient but very difficult to test



clearly possible at low scale: if seesaw seesaw states have a quasi-degenerate mass spectrum and/or if large cancellation among Yukawa couplings

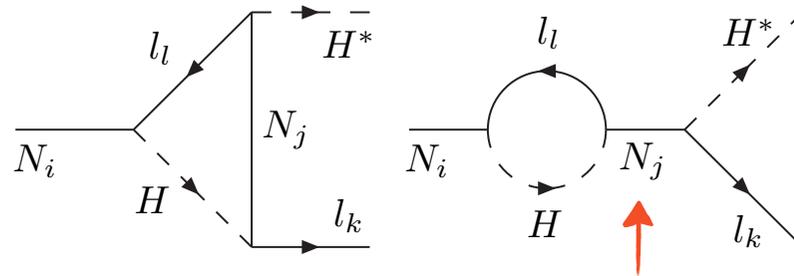
 this talk: new way at low scale: total lepton number violating Higgs doublet decay into ~ 0.1 - 100 GeV right-handed neutrinos

Leptogenesis relevant scales for low m_N

$$T_{Sphaler.} \sim 135 \text{ GeV}$$

usual leptogenesis: $m_N \gg T_{Sphaler.} > m_{H,L}$: leptogenesis from $N \rightarrow LH$ decay

↪ creation of L asymmetry at $T \sim m_N \gg T_{Sphaler.} \Rightarrow$ B asymmetry



resonant propagator if $m_{N_j} \sim m_{N_i} \Rightarrow$ \sim TeV scale leptogenesis

very low scale leptogenesis: $T_{Sphaler.} > m_H \gg m_{N,L}$

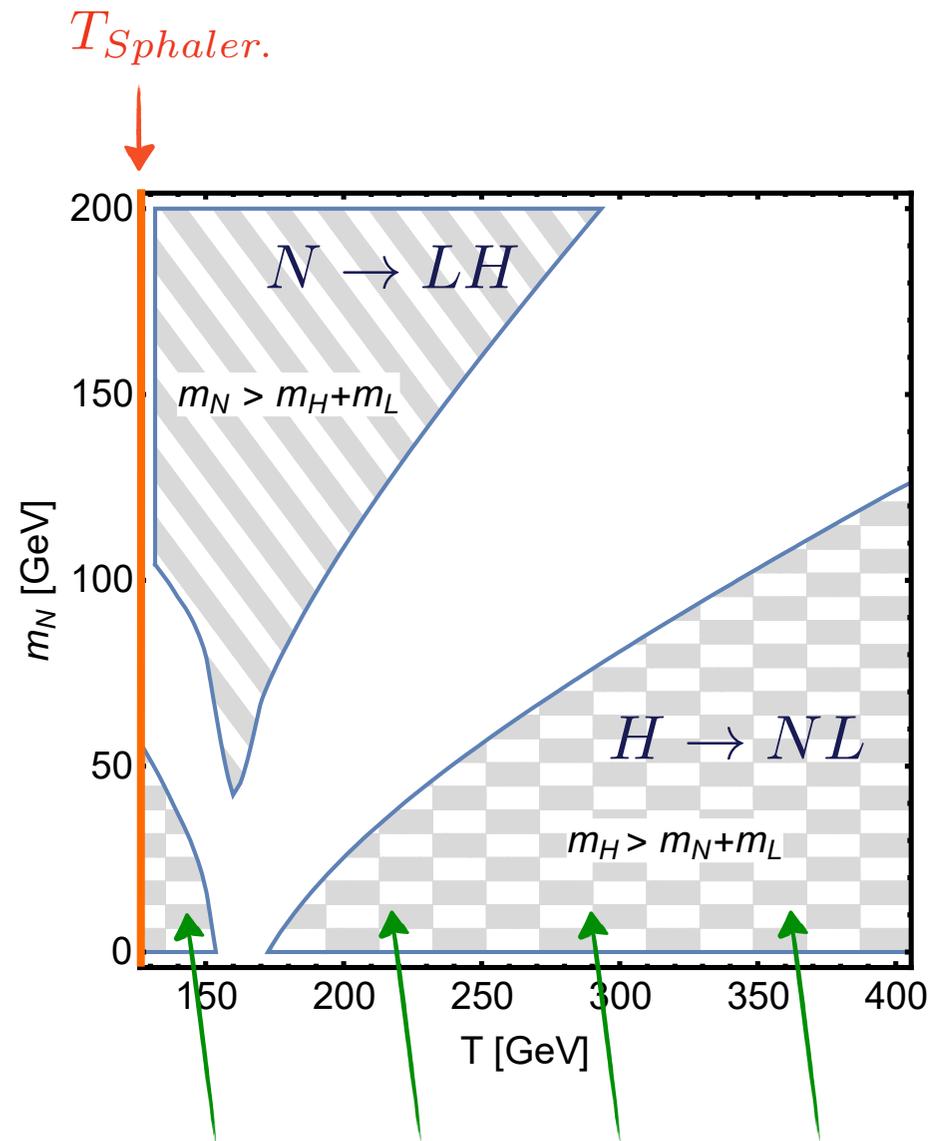
↪ creation of L asymmetry at $T > T_{Sphaler.} \gg m_N \Rightarrow \neq$ regime

↪ thermal effects are fully relevant: $T > T_{Sphaler.} > m_H \gg m_{N,L}$

$$m_H^2(T) = m_H^2 + c_H \cdot T^2 \quad m_L^2(T) = m_L^2 + c_L \cdot T^2 \quad m_N^2(T) = m_N^2 + c_N \cdot T^2$$

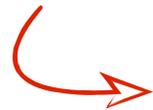
↪ $N \rightarrow LH$ forbidden but $H \rightarrow NL$ allowed

Temperatures allowing the $N \rightarrow LH$ and $H \rightarrow NL$ decays



$H \rightarrow NL$ leptogenesis from this region?

L asymmetry production from $H \rightarrow NL$ decay

 2 issues at first sight:

1) out-of-equilibrium decay?  3rd Sakharov condition

 H decaying particle is in deep thermal equilibrium at $T > T_{Sphaler.}$
but N in decay product is not necessarily in thermal equilibr.

$$\frac{dn_N}{dt} \propto (n_N^{eq} - n_N) \cdot \Gamma_{H \rightarrow NL}$$

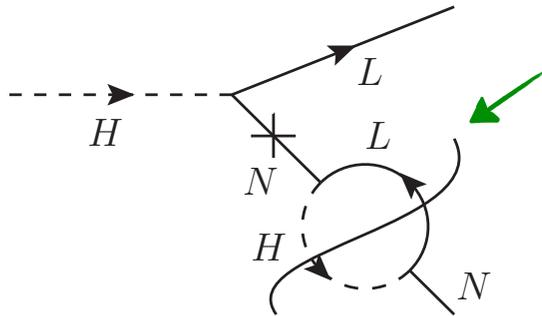
 

$H \rightarrow NL \quad NL \rightarrow H$

L asymmetry production from $H \rightarrow NL$ decay

T.H., Teresi 16'

2) Absorptive part for CP violation?



$m_H + m_L > m_N \Rightarrow$ no absorptive part?

but only for $T = 0!$

finite T corrections: thermal cut: if H or L comes from the thermal bath the cut is kinematically allowed

Giudice, Notari, Raidal, Riotto, Strumia 03'

Frossard, Garny, Hohenegger, Kartavtsev, Mitrouskas 12'

\Rightarrow absorptive part $\Gamma_N(T)$ (calculated in Kadanoff Baym formalism)

Frossard, Garny, Hohenegger, Kartavtsev, Mitrouskas 12'

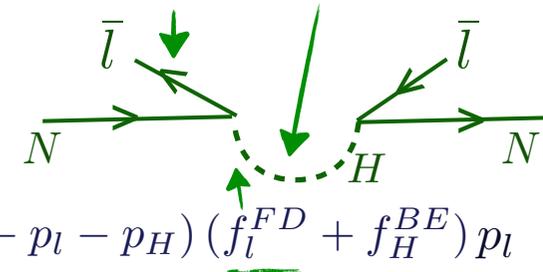
$$\Pi^{\alpha\beta}(q_N) = \Pi_{>}^{\alpha\beta}(q_N) + \Pi_{<}^{\alpha\beta}(q_N)$$

$$\Rightarrow -2 \int \frac{d^4 p_l}{(2\pi)^4} \frac{d^4 p_H}{(2\pi)^4} (2\pi)^4 \delta^4(q_N - p_l - p_H) (Y_N^\dagger Y_N)_{\alpha\beta} [P_L S_{>}^l(t, p_l) P_R \Delta_{>}^H(t, p_H)]$$

$$\propto f_{\bar{l}}^{FD} \propto 1 + f_H^{BE}$$

$$\Rightarrow \Gamma_N(T) = \frac{1}{8\pi} m_{N_2} (Y_N Y_N^\dagger)_{22} \cdot \frac{p \cdot L_N}{q_N \cdot p_l}$$

$$\Rightarrow L_N = 16\pi \int \frac{d^3 p_l}{(2\pi)^3 2E_l} \frac{d^3 p_H}{(2\pi)^3 2E_H} (2\pi)^4 \delta^4(q_N - p_l - p_H) (f_{\bar{l}}^{FD} + f_H^{BE}) p_l$$



Total L number violating CP asymmetry

$$\epsilon_{CP} = \frac{\text{Im}[(Y_N Y_N^\dagger)_{12}^2]}{(Y_N Y_N^\dagger)_{11} (Y_N Y_N^\dagger)_{22}} \cdot \frac{2 \Delta m_N^0 \Gamma_N(T)}{4 \Delta m_N(T)^2 + \Gamma_N(T)^2}$$

↪ with thermal mass splitting: $\Delta m_N(T) \simeq \Delta m_N^0 + \frac{\pi T^2}{4 m_N^2} \Gamma_{22} \sqrt{\left(1 - \frac{\Gamma_{11}}{\Gamma_{22}}\right)^2 + 4 \frac{|\Gamma_{12}|^2}{\Gamma_{22}^2}}$

$$\Gamma_{ij} \equiv m_N (Y_N Y_N^\dagger)_{ij} / (8\pi)$$

Boltzmann equations:

$$\frac{n_\gamma H_N}{z} \frac{d\eta_N}{dz} = \left(1 - \frac{\eta^N}{\eta_N^{\text{eq}}}\right) \left[\gamma_D + 2(\gamma_{Hs} + \gamma_{As}) + 4(\gamma_{Ht} + \gamma_{At}) \right],$$

$$\frac{n_\gamma H_N}{z} \frac{d\eta_L}{dz} = \gamma_D \left[\left(\frac{\eta^N}{\eta_N^{\text{eq}}} - 1 \right) \epsilon_{CP}(z) - \frac{2}{3} \eta_L \right] - \frac{4}{3} \eta_L \left[2(\gamma_{Ht} + \gamma_{At}) + \frac{\eta^N}{\eta_N^{\text{eq}}} (\gamma_{Hs} + \gamma_{As}) \right]$$

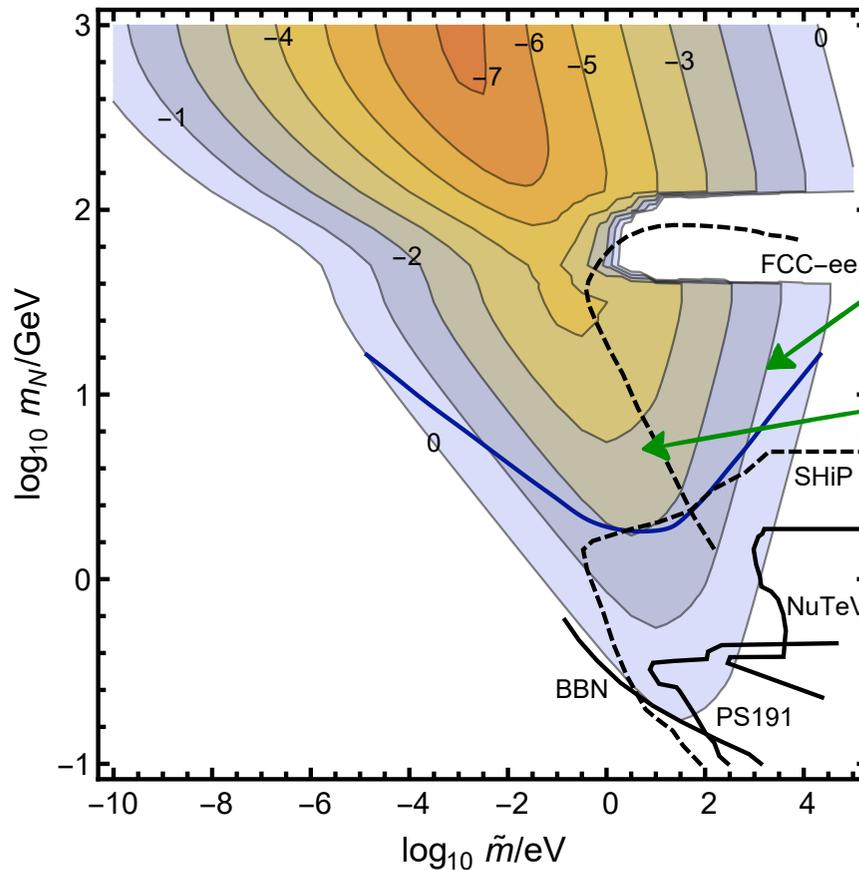
$$\eta_N \equiv n_N / n_\gamma$$

$$z \equiv m_N / T$$

Results for the case where the N have thermalized

if N thermalized by large Y_N Yukawas or other interaction (e.g. a W_R) before an asymmetry is produced

CP-asymmetry needed for successful leptog.



the lower is m_N , the later it goes out-of-equilibrium, the more it will be in equilibrium at $T > T_{Sphaler}$.



lower bound on m_N

$$m_N > 2.2 \text{ GeV}$$

if only $N \rightarrow LH$ decay we get: $m_N > 50 \text{ GeV}$

$$\tilde{m} \equiv \frac{Y_N Y_N^\dagger v^2}{2m_N}$$

requires that at least 2 of the N have quasi-degenerate masses

Results for the case where the N have not thermalized

- if no extra interaction thermalizing N , no thermalization is much more natural than in ordinary leptogenesis: thermalization at $T > T_{Sphaler.} \gg m_N$ requires much larger Y_N Yukawas than in ordinary leptogenesis at $T \sim m_N$

$$\tilde{m} \equiv \frac{Y_N Y_N^\dagger v^2}{2m_N}$$

$$\tilde{m} \gg 10^{-3} \text{ eV}$$

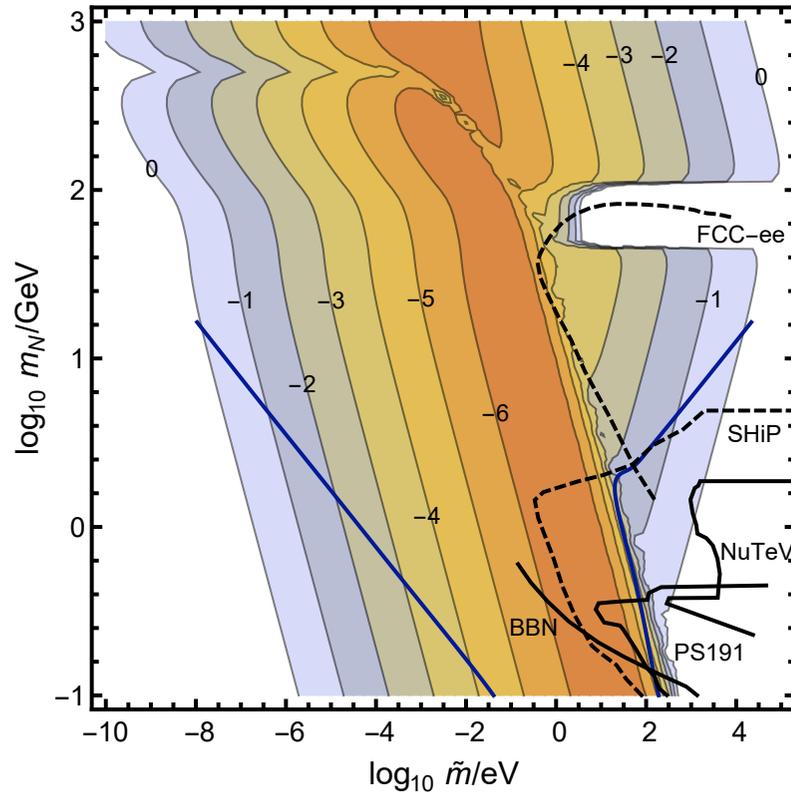
$$\tilde{m} \gtrsim 10^{-3} \text{ eV}$$

- for $H \rightarrow NL$ decay, to start from no N in the thermal bath boosts the asymmetry production, unlike for ordinary $N \rightarrow LH$ leptogenesis

$H \rightarrow NL$: many H to decay and produce the asymmetry but few N to $NL \rightarrow H$ inverse decay

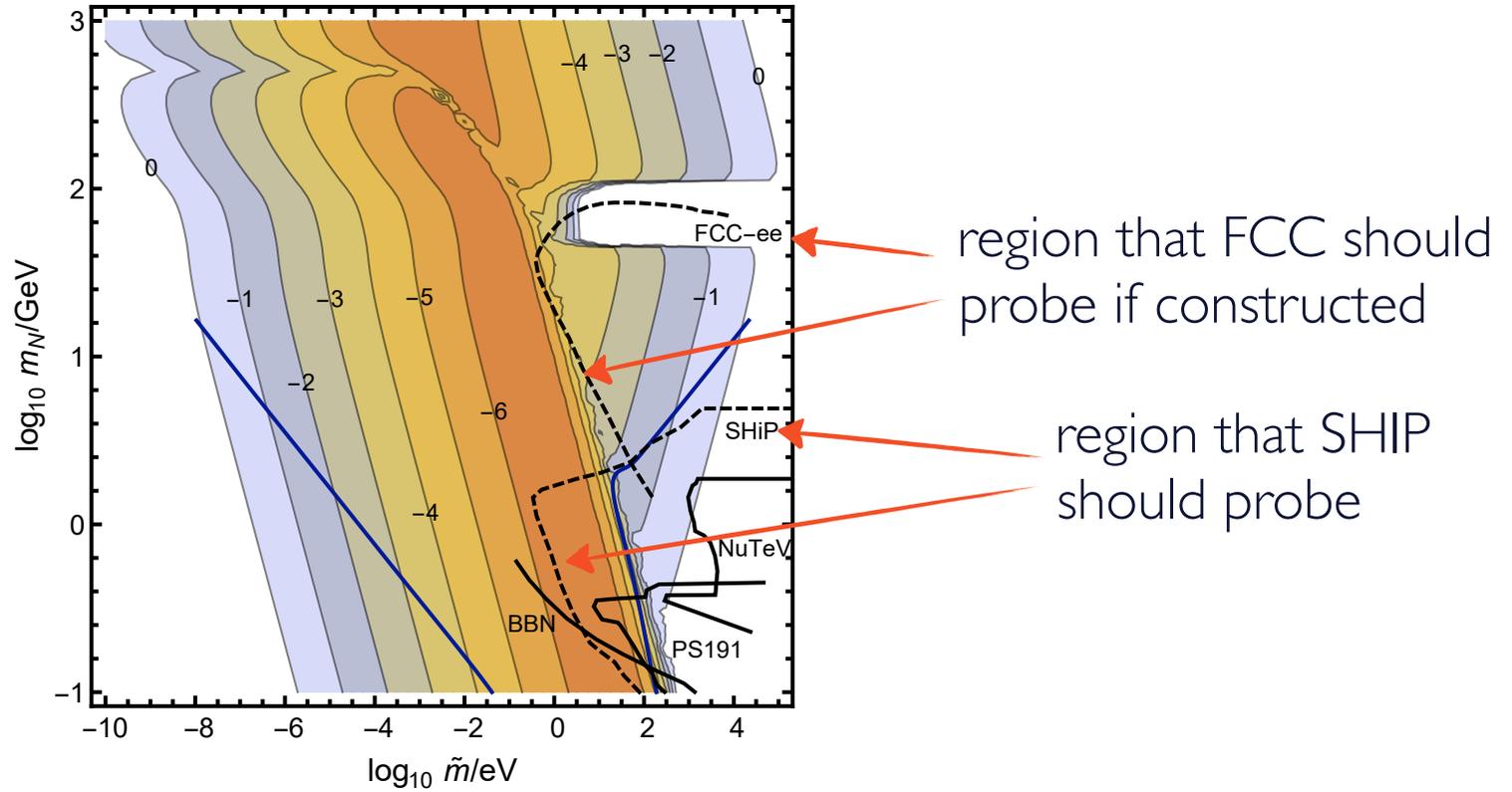
$$n_N^{eq} - n_N \sim n_N^{eq} \gg n_N$$

Results for the case where the N have not thermalized



- ↪ for example for $m_N \sim 10 \text{ GeV}$ and $\tilde{m} \sim 0.1 \text{ eV}$ one needs $\Delta m_N^0/m_N \lesssim 10^{-5}$
- ↪ leptogenesis for m_N as low as $\sim 20 \text{ MeV}$ is possible (but BBN concerns)
- ↪ in all cases: asymmetry production at T just above $T_{Sphaler}$. \Rightarrow no dependence on UV physics!

Testability!



Two important comparisons to do

- for $m_N \sim \text{GeV}$: well-known baryogenesis mechanism in seesaw model:
baryogenesis from right-handed neutrino oscillations: ``ARS'' mechanism

Akhmedov, Rubakov, Smirnov 98'

Asaka, Shaposhnikov 05'; Shaposhnikov 08'

Drewes, Garbrecht 11'

Canetti, Drewes, Frossard, Shaposhnikov 13'

Hernandez, Kekic, Lopez-Pavon, Racker, Rius 15'

.....

 comparison of ARS with L-violating Higgs decay setup???

- to compute evolution of asymmetries with thermal effects: another well-known formalism: density matrix formalism

 comparison of results of decay formalism and density matrix formalism???

Density matrix formalism

N_{R_α} quantum system is described by density matrix : $n_{\alpha\beta}^N \equiv \langle a_\beta^{+\dagger} a_\alpha^+ \rangle = Tr(\rho a_\beta^{+\dagger} a_\alpha^+)$

$\overline{N_{R_\alpha}}$ quantum system is described by density matrix : $n_{\alpha\beta}^{\bar{N}} \equiv \langle a_\beta^{-\dagger} a_\alpha^- \rangle = Tr(\rho a_\beta^{-\dagger} a_\alpha^-)$

$n_{\alpha\alpha}^N = n_\alpha^N =$ number density of N_α states

$n_{\alpha\beta}^N =$ coherence between N_α and N_β states

\Rightarrow evolution of density matrix:

$$\frac{d}{dt} n_{\alpha\beta}^N(\mathbf{k}, t) = i \langle [H_0^N, n_{\alpha\beta}^N(\mathbf{k}, t)] \rangle - \frac{1}{2} \int_{-\infty}^{\infty} dt' \langle [H_{int}(t'), [H_{int}(t), n_{\alpha\beta}^N(\mathbf{k}, t)]] \rangle_t$$

oscillation term

interaction term

$$H_{int} = h_{l\alpha} \bar{L}_l \tilde{H} P_R N_\alpha + h.c.$$

$\hookrightarrow H_{int} \cdot H_{int}$

- terms in $a_\beta^{+\dagger} a_\alpha^+ \rightarrow n_{\alpha\beta}^N$
- terms in $a_\alpha^- a_\beta^{-\dagger} \rightarrow 1 - n_{\alpha\beta}^N$

$\Rightarrow \frac{d}{dt} n_{\alpha\beta}^N(\mathbf{k}) = -i [E_N, n^N(\mathbf{k})]_{\alpha\beta} - \frac{1}{2E_N} \left(\frac{1}{2} \{ \Gamma^>(\mathbf{k}), n^N(\mathbf{k}) \} - \frac{1}{2} \{ \Gamma^<(\mathbf{k}), 1 - n^N(\mathbf{k}) \} \right)_{\alpha\beta}$

$N + \bar{L} \rightarrow H$ $H \rightarrow N + \bar{L}$

ARS contribution in density matrix formalism

 keeping only the transitions where there is no m_N mass insertions because the asymmetry is produced at $T \gg T_{\text{sphaler.}} \gg m_N$
 if mass insertion: m_N^2/T^2 suppression

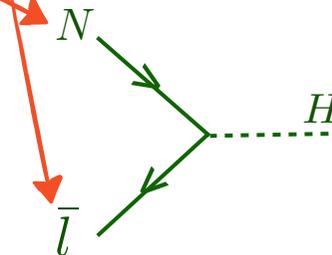
$$\Gamma_{\alpha\beta}^{\lessgtr}(\mathbf{k}) = -i \text{tr} \{ P_R u_+(\mathbf{k}) \bar{u}_+(\mathbf{k}) P_L \Sigma_{\alpha\beta}^{\lessgtr}(k) \}$$

with:

$$-i \Sigma_{\alpha\beta}^{\lessgtr}(k) = \int \frac{d^4 p}{(2\pi)^4} \int \frac{d^4 q}{(2\pi)^4} (2\pi)^4 \delta^{(4)}(q - k - p) \underbrace{i S_l^{\lessgtr}(-p)}_{\text{Wightman propagator of } L} i \underbrace{\Delta^{\lessgtr}(-q)}_{\text{Wightman propagator of } H} h_{l\alpha}^* h_{l\beta}$$

for example:

$$n_N \Gamma^> \sim n_N \Sigma^> \propto n_N \cdot S_l^> \cdot \Delta^< \cdot h_{l\alpha}^* h_{l\beta} \propto n_N \cdot n_{\bar{l}} \cdot (1 + n_H) \cdot h_{l\alpha}^* h_{l\beta}$$



Density matrix formalism: evolution equations and rates

ARS scenario

writing
$$\begin{aligned} n_l &= n_l^{Eq} + \frac{\delta n_l}{2} \\ n_{\bar{l}} &= n_l^{Eq} - \frac{\delta n_l}{2} \end{aligned}$$
 a term proportional to δn_l shows up: washout term: W

$$\frac{dn_{\alpha\beta}^N}{dt} = -i [\mathcal{E}_N, n^N(k)]_{\alpha\beta} - \frac{1}{2} \left\{ \gamma^{LC}, \frac{n^N}{n_{eq}^N} - \mathbf{I} \right\}_{\alpha\beta} + \frac{\delta n_l^L}{2n_{eq}^L} \left((\gamma_{WQ,l}^{LC}) + \frac{1}{2} \left\{ \gamma_{WC,l}^{LC}, \frac{n^N}{n_{eq}^N} \right\} \right)_{\alpha\beta}$$

and similarly for \bar{N}

$$\begin{aligned} \frac{d\delta n_l^L}{dt} &= \frac{1}{n_{eq}^N} \text{tr} \{ (\gamma_l^{LC}) n^N \} - \frac{1}{n_{eq}^N} \text{tr} \{ (\gamma_l^{LC*}) \bar{n}^N \} \\ &\quad - \frac{\delta n_l^L}{n_{eq}^L} \text{tr} \{ \gamma_{WQ,l}^{LC} \} - \frac{\delta n_l^L}{2n_{eq}^L} \frac{1}{n_{eq}^N} \text{tr} \{ n^N (\gamma_{WC,l}^{LC}) \} - \frac{\delta n_l^L}{2n_{eq}^L} \frac{1}{n_{eq}^N} \text{tr} \{ \bar{n}^N (\gamma_{WC,l}^{LC*}) \} \end{aligned}$$

$$\begin{aligned} \gamma_{\alpha\beta}^{LC} &\equiv \int d\Pi_{PS} n_{eq}^N(\mathbf{k}) (n_{eq}^L(\mathbf{p}) + n_{eq}^H(\mathbf{q})) \times \text{tr} \{ \mathbf{P}_R u_+(\mathbf{k}) \bar{u}_+(\mathbf{k}) \mathbf{P}_L \not{p} \} h_{l\alpha}^* h_{l\beta} \\ &= \frac{g_{iso}}{4\pi^2} \int_0^\infty dk \frac{k}{e^{k/T} + 1} \frac{M_H^2 - M_L^2}{8\pi k} \times \int_{E^*}^\infty dE \left(\frac{1}{e^{\frac{E}{T}} + 1} + \frac{1}{e^{\frac{E+k}{T}} - 1} \right) h_{l\alpha}^* h_{l\beta} \\ &\simeq 3.26 \times 10^{-4} T^4 h_{l\alpha}^* h_{l\beta} \end{aligned}$$

$$\gamma_{WQ}^{LC} = \frac{g_{iso}}{4\pi^2} \int_0^\infty dk k \frac{M_H^2 - M_L^2}{8\pi k} \times \int_{E^*}^\infty dE \frac{1}{e^{\frac{E}{T}} + 1} \frac{1}{e^{\frac{E+k}{T}} - 1} h_{l\alpha}^* h_{l\beta} \simeq 1.05 \times 10^{-4} T^4 h_{l\alpha}^* h_{l\beta}$$

$$\gamma_{WC}^{LC} = \frac{g_{iso}}{4\pi^2} \int_0^\infty dk \frac{k}{e^{k/T} + 1} \frac{M_H^2 - M_L^2}{8\pi k} \int_{E^*}^\infty dE \frac{1}{e^{\frac{E}{T}} + 1} h_{l\alpha}^* h_{l\beta} \simeq 1.86 \times 10^{-4} T^4 h_{l\alpha}^* h_{l\beta}$$

Density matrix formalism: ARS final result

first non-vanishing term in Yukawa coupling expansion

$$Y_{LC} \simeq -18.5 \times (\alpha^{LC})^2 \alpha_W^{LC} \frac{M_0^{7/3}}{T_c (\Delta m_N^2)^{2/3}} \times (h^\dagger h)_{11} (h^\dagger h)_{22} \sum_l \delta_l^{LC} (hh^\dagger)_{ll}$$

$$\delta_l^{LC} = \frac{\text{Im}[h_{l1}^* h_{l2} (h^\dagger h)_{21}]}{(h^\dagger h)_{11} (h^\dagger h)_{22}}$$

ARS is based on processes where there is no m_N mass insertions

only $N \rightarrow L$ and $\bar{N} \rightarrow \bar{L}$ transitions

no $N \rightarrow \bar{L}$

no $\bar{N} \rightarrow L$

assigning $L = 1$ to N and $L = -1$ to \bar{N} , all processes conserve L :

$$\sum_{i=e,\mu,\tau} (n_{L_i} - n_{\bar{L}_i}) + \sum_{\alpha} (n_{N_\alpha} - n_{\bar{N}_\alpha}) = 0$$

at $\mathcal{O}(h^4)$: SM lepton number and N lepton number are separately conserved

$$\sum_{i=e,\mu,\tau} (n_{L_i} - n_{\bar{L}_i}) = 0 \quad \sum_{\alpha} (n_{N_\alpha} - n_{\bar{N}_\alpha}) = 0 \quad \text{but flavor lepton number is not conserved:}$$

$$\delta n_{l_i} \equiv n_{L_i} - n_{\bar{L}_i} \neq 0$$

at $\mathcal{O}(h^6)$: if for example Yukawa for electron much smaller than for muon:

$n_{L_\mu} - n_{\bar{L}_\mu}$ strongly washed-out
 $n_{L_e} - n_{\bar{L}_e}$ much less washed-out

$$\sum_{i=e,\mu,\tau} (n_{L_i} - n_{\bar{L}_i}) = - \sum_{\alpha} (n_{N_\alpha} - n_{\bar{N}_\alpha}) \neq 0$$

converted to B asym by sphalerons

not converted to B asym. by sphalerons

baryon asymmetry!

Total lepton number violating density matrix contribution

T.H., Teresi 17'

→ the L-violating Higgs decay contribution to baryogenesis is clearly \neq from the ARS one since it is a $\mathcal{O}(h^4)$ contribution based on processes which do involve a Majorana mass insertion, i.e. which do violate total lepton number, unlike ARS

→ where to find this contribution in density matrix formalism??

⇓
the density matrix commutators lead also to contributions $\propto m_N^2$
which corresponds to processes with a Majorana mass insertion

↑
 N to \bar{L} transition instead of N to L transition

$$\Gamma_{\alpha\beta}^{\leq}(\mathbf{k}) \ni + i \text{tr} \{ P_R v_+(\mathbf{k}) \bar{v}_+(\mathbf{k}) P_L \Sigma_{\beta\alpha}^{\leq}(-k) \}$$

Full set of density matrix equation with LC and LV contributions

T.H., Teresi 17'

$$\begin{aligned}
 \frac{dn_{\alpha\beta}^N}{dt} &= -i [\mathcal{E}_N, n^N(\mathbf{k})]_{\alpha\beta} - \frac{1}{2} \left\{ \gamma^{LC} + \gamma^{LV}, \frac{n^N}{n_{eq}^N} - \mathbf{I} \right\}_{\alpha\beta} & \frac{d\delta n_l^L}{dt} &= \frac{1}{n_{eq}^N} \text{tr} \{ (\gamma_l^{LC} - \gamma_l^{LV}) n^N \} \\
 &+ \frac{\delta n_l^L}{2n_{eq}^L} \left((\gamma_{WQ,l}^{LC} - \gamma_{WQ,l}^{LV}) + \frac{1}{2} \left\{ \gamma_{WC,l}^{LC} - \gamma_{WC,l}^{LV}, \frac{n^N}{n_{eq}^N} \right\} \right)_{\alpha\beta}, & & - \frac{1}{n_{eq}^N} \text{tr} \{ (\gamma_l^{LC*} - \gamma_l^{LV*}) \bar{n}^N \} \\
 & & & - \frac{\delta n_l^L}{n_{eq}^L} \text{tr} \{ \gamma_{WQ,l}^{LC} + \gamma_{WQ,l}^{LV} \} \\
 \frac{d\bar{n}_{\alpha\beta}^N}{dt} &= -i [\mathcal{E}_N, \bar{n}^N(\mathbf{k})]_{\alpha\beta} & & - \frac{\delta n_l^L}{2n_{eq}^L} \frac{1}{n_{eq}^N} \text{tr} \{ n^N (\gamma_{WC,l}^{LC} + \gamma_{WC,l}^{LV}) \} \\
 &- \frac{1}{2} \left\{ \gamma^{LC*} + \gamma^{LV*}, \frac{\bar{n}^N}{n_{eq}^N} - \mathbf{I} \right\}_{\alpha\beta} & & - \frac{\delta n_l^L}{2n_{eq}^L} \frac{1}{n_{eq}^N} \text{tr} \{ \bar{n}^N (\gamma_{WC,l}^{LC*} + \gamma_{WC,l}^{LV*}) \} \\
 &- \frac{\delta n_l^L}{2n_{eq}^L} \left((\gamma_{WQ,l}^{LC*} - \gamma_{WQ,l}^{LV*}) + \frac{1}{2} \left\{ \gamma_{WC,l}^{LC*} - \gamma_{WC,l}^{LV*}, \frac{\bar{n}^N}{n_{eq}^N} \right\} \right)
 \end{aligned}$$

$$\begin{aligned}
 \gamma^{LV} &= \frac{g_{iso}}{4\pi^2} \int_0^\infty dk \frac{k}{e^{k/T} + 1} \frac{m_N^2}{32\pi k^3} \times \int_{E^*}^\infty dE (4Ek + M_L^2 - M_H^2) \left(\frac{1}{e^{\frac{E}{T}} + 1} + \frac{1}{e^{\frac{E+k}{T}} - 1} \right) h_{l\alpha}^* h_{l\beta} \\
 &\simeq 3.35 \times 10^{-3} m_N^2 T^2 h_{l\alpha}^* h_{l\beta}
 \end{aligned}$$

$$\gamma_{WQ}^{LV} = \frac{g_{iso}}{4\pi^2} \int_0^\infty dk k \frac{m_N^2}{32\pi k^3} \times \int_{E^*}^\infty dE (4Ek + M_L^2 - M_H^2) \left(\frac{1}{e^{\frac{E}{T}} + 1} + \frac{1}{e^{\frac{E+k}{T}} - 1} \right) h_{l\alpha}^* h_{l\beta} \simeq 5.49 \times 10^{-4} m_N^2 T^2 h_{l\alpha}^* h_{l\beta}$$

$$\gamma_{WC}^{LV} = \frac{g_{iso}}{4\pi^2} \int_0^\infty dk \frac{k}{e^{k/T} + 1} \frac{m_N^2}{32\pi k^3} \times \int_{E^*}^\infty dE (4Ek + M_L^2 - M_H^2) \frac{1}{e^{\frac{E}{T}} + 1} h_{l\alpha}^* h_{l\beta} \simeq 1.79 \times 10^{-3} m_N^2 T^2 h_{l\alpha}^* h_{l\beta}$$

Analytical solution for the LV contribution for weak washout

$$Y_{LV} \simeq 7.9 \times \alpha^{LC} \alpha^{LV} \frac{M_0}{T_c} \frac{m_N^2}{\Delta m_N^2} (h^\dagger h)_{11} (h^\dagger h)_{22} \delta^{LV} \leftarrow \delta^{LV} \equiv \sum_l \delta_l^{LV} \neq 0 \quad \text{T.H., Teresi 17'}$$

CP-violating Yukawa combination which breaks total lepton number $\rightarrow \delta_l^{LV} = \frac{\text{Im}[h_{l1}^* h_{l2} (h^\dagger h)_{12}]}{(h^\dagger h)_{11} (h^\dagger h)_{22}}$

$$Y_{LC} \simeq -18.5 \times (\alpha^{LC})^2 \alpha_W^{LC} \frac{M_0^{7/3}}{T_c (\Delta m_N^2)^{2/3}} \times (h^\dagger h)_{11} (h^\dagger h)_{22} \sum_l \delta_l^{LC} (hh^\dagger)_{ll} \leftarrow \delta_l^{LC} = \frac{\text{Im}[h_{l1}^* h_{l2} (h^\dagger h)_{21}]}{(h^\dagger h)_{11} (h^\dagger h)_{22}}$$

CP-violating Yukawa combination which leaves the SM total lepton number unchanged

\Rightarrow LV vs LC contributions:

- $\mathcal{O}(h^4)$ instead of $\mathcal{O}(h^6)$ for the LC contribution
- suppressed by 2 rates instead of 3 rates for the LC contribution

$$\alpha^{LC} \alpha^{LV} \quad (\alpha^{LC})^2 \alpha_W^{LC} \quad \sim \text{Planck mass}$$

- but m_N^2 suppression with different Δm_N^2 and M_0 dependence

\Rightarrow all in all the various factors compensate each other more or less with dominance of one or the other contribution depending on the parameters

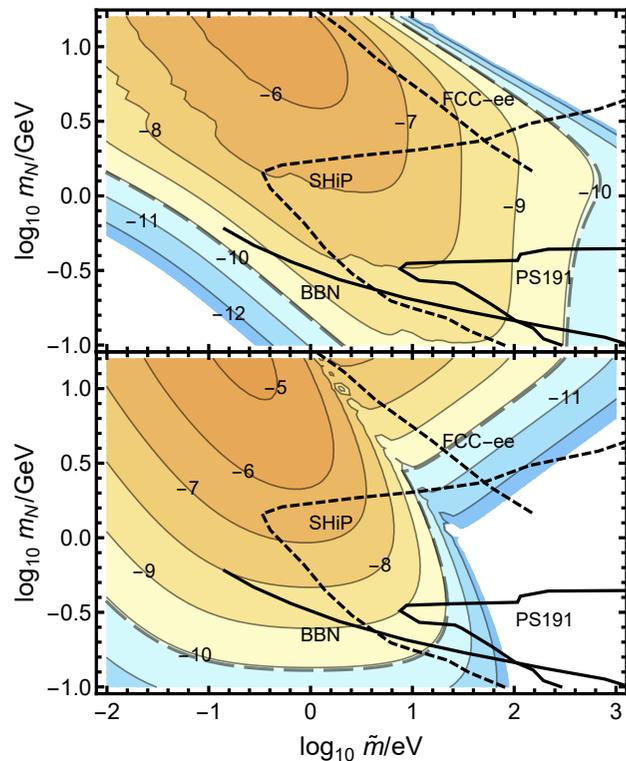
Numerical results: comparison of decay and density matrix formalisms for the LV contribution

T.H., Teresi 17'

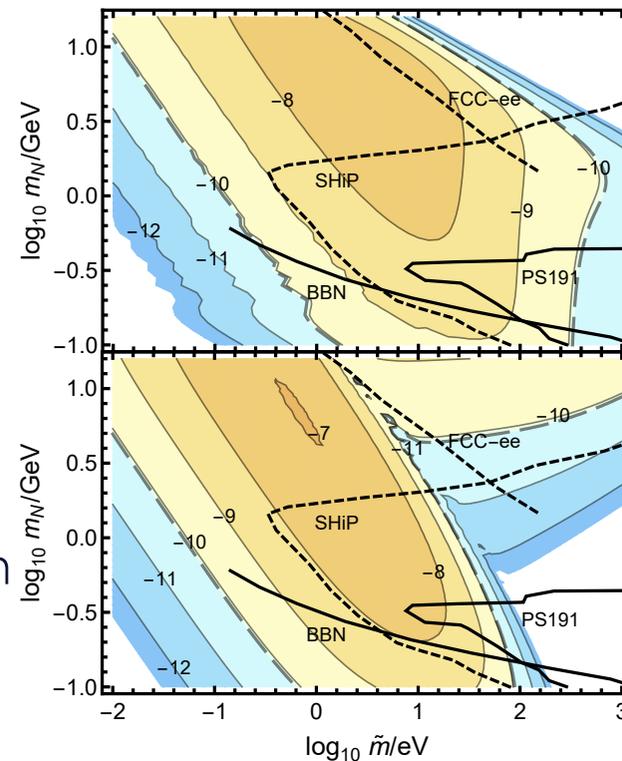
↪ with only one lepton flavour: no ARS, only LV contribution

$Y_B = \frac{n_B}{s}$ contour plot $\Delta m_N/m_N = 10^{-10}$

$Y_B = \frac{n_B}{s}$ contour plot $\Delta m_N/m_N = 10^{-8.5}$



density matrix



decay formalism

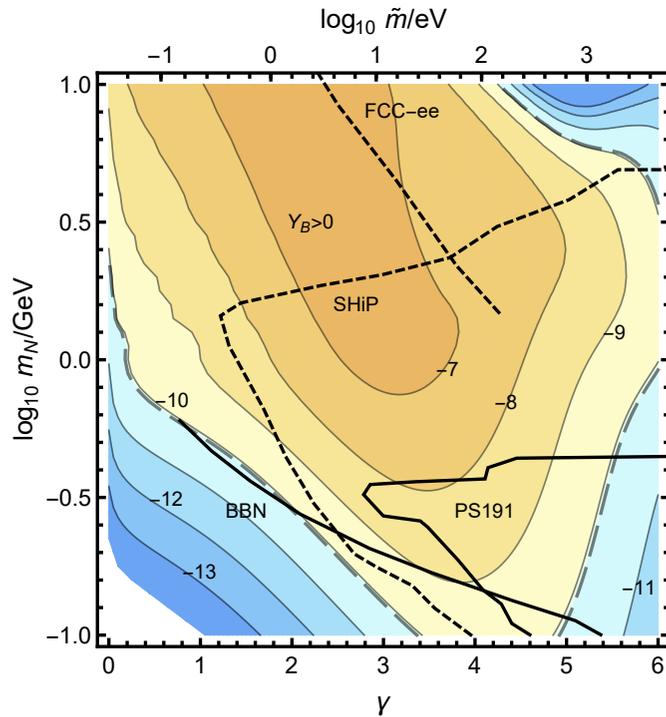
⇒ qualitative or even quantitative agreement:

- except for small m_N : different thermal masses taken
- except for large \tilde{m} : washout suppression too big in decay formalism because doesn't take into account formation of $N - \bar{N}$ asymmetries
- in decay formalism the H is decaying "at rest" unlike in density matrix formalism

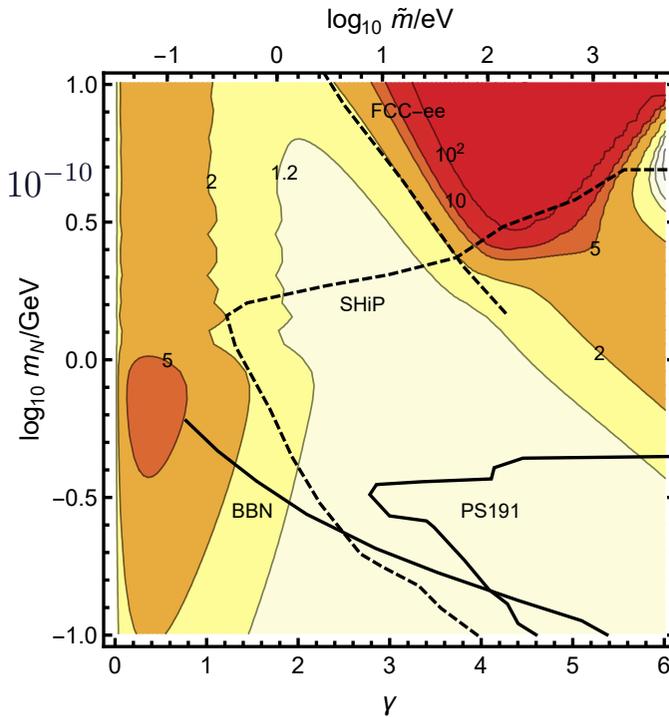
Numerical results: comparison of LC and LV contributions in matrix density formalism

T.H., Teresi 17'

$Y_B = \frac{n_B}{s}$ contour plot: full LC+LV result



ratio of LV+LC over LC



dominance of LV = - for "seesaw" expected Yukawa couplings

- for very large Yukawas: less washout for LV than for LC!

$$\propto m_N^2 T^2 \ll \propto T^4$$

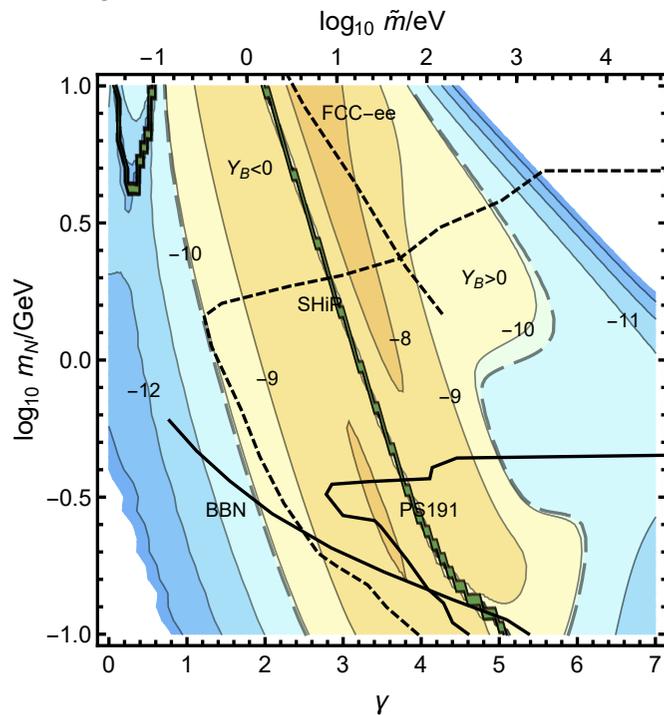
- the smaller $\Delta m_N/m_N$ the more LV dominates

- the larger m_N the more LV dominates

Numerical results: comparison of LC and LV contributions in matrix density formalism

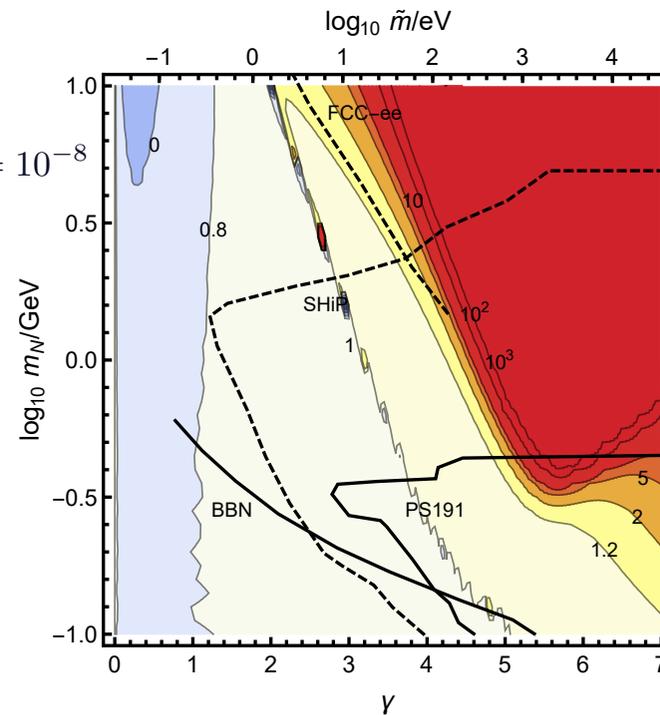
T.H., Teresi 17'

$Y_B = \frac{n_B}{s}$ contour plot: full LC+LV result



$\Delta m_N / m_N = 10^{-8}$

ratio of LV+LC over LC



dominance of LV = - for "seesaw" expected Yukawa couplings

- for very large Yukawas: less washout for LV than for LC !

$$\propto m_N^2 T^2 \ll \propto T^4$$

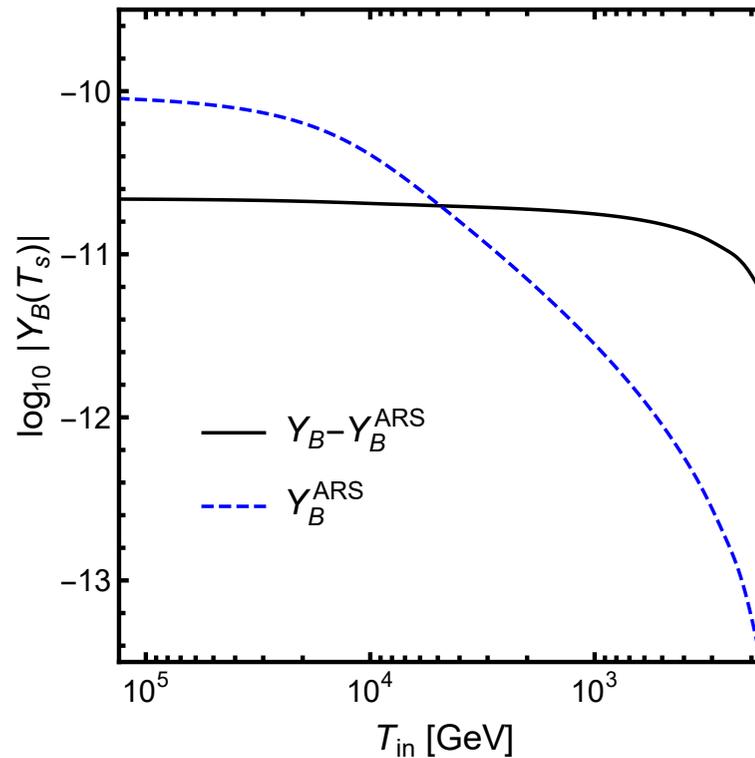
- the smaller $\Delta m_N / m_N$ the more LV dominates

- the larger m_N the more LV dominates

Dominance of the LV contribution for low reheating temperatures

T.H., Teresi 17'

↪ LV contribution produced at lower temperature than the ARS-LC contribution due to the m_N^2/T^2 relative factor



Need to incorporate other processes for a full quantitative asymmetry calculation

→ top quark scattering processes, gauge scattering processes, ... Besak, Bodeker 12'
see also Ghiglieri, Laine 17'

→ have all a $H \rightarrow LN$ transition as building block ⇒ same mechanism
expected to be
operative

→ additional effect found to be small

Drewes, Garbrecht, Güter, Klaric, 16'

Summary

In usual leptogenesis decay formalism the L violating $H \rightarrow NL$ decay can easily lead to enough baryon asymmetry for $m_N < m_H$

↳ in type-I seesaw model with nothing else

↳ thanks to thermal effect leading to N self-energy thermal cut

↳ from total L number violating CP asymmetries: no need for flavour interplay

↳ at electroweak scale temperatures: $T \gtrsim T_{Sphaler}$.

↳ with boosted production if no N to begin with

↳ in a testable way (SHIP,...) for part of the parameter space

We have confirmed these results in density matrix formalism...

↳ both ARS-LC and LV contributions can dominate baryogenesis depending on parameters

baryon asymmetries obtained for 3 values of $\Delta m_N^0/m_N$

