# Baryogenesis in the extended Standard Model with a Complex Singlet and Iso-Doublet Vector Quarks

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# Outline

▶ Introduction to baryogenesis at the electroweak phase transitions

#### ► The Model

- \* Potential
- \* Constrained Potential
- \* Mass matrix
- \* Extremum conditions
- \* Region for possible CP violation
- \* Numerical analysis of the parameter space of the model

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- ▶ The electroweak Phase Transition
- ▶ Baryon Generation
- ▶ Summery and Conclusion

# A Universe of matter

The observations from WMAP experiment:

$$\frac{\eta}{7.04} = \frac{n_B - n_{\bar{B}}}{s} \approx \frac{n_B}{s} = 8.7 \pm 0.3 \times 10^{-11}$$

Sakharov conditions:

- Baryon number violation
- ▶ C and CP violation
- ▶ Departure from thermal equilibrium



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### Baryogenesis at the electroweak phase transitions

▶ Baryon-number can be generated during the first order EW phase transition





- Sphaleron rate at high temperature  $(T > T_{EW}), \Gamma_{sph} = K(\alpha_W T)^4$
- ► Baryon violation- reactions must be strongly suppressed in  $T < T_{EW} \sim 100 GeV$ ,  $\Gamma_{sph} = K(\alpha_W T)^4 exp(\frac{-E_{sph}}{T})$
- ► The out of thermal equilibrium → strong enough first order phase transition
- ▶ Strong enough first order  $\rightarrow \frac{E_{sph}(T_c)}{T_c} \ge 46$  or  $\frac{v(T_c)}{T_c} \ge 1$
- ▶ SM problem:  $\frac{v(T_c)}{T_c} \ge 1$  requires  $m_{Higgs} < 72 GeV$  while  $m_{Higgs} \simeq 125 GeV$

• Baryon generation dominated over the conjugate process through CP-violation.



- ▶ In the SM the amount of CP violation  $\rightarrow J_{CP}$  invariant from CKM matrix
- $\blacktriangleright$  The magnitude of the baryon asymmetry of the Universe in SM related to  $J_{CP}.$

$$\eta = \frac{n_B - n_{\bar{B}}}{n_{\gamma}} \approx \frac{n_B}{n_{\gamma}} \sim \frac{\prod_u^{i>j} (m^i - m^j) \prod_d^{i>j} (m^i - m^j) J_{CP}}{T_{EW}^{12}}$$

- ►  $J_{CP} \sim O(10^{-5})$
- ▶ SM problem: This gives a baryon asymmetry  $O(10^{-17})$  much much below than the observed value of  $O(10^{-11})$ .

- The CP violation of SM is insufficient to explain the baryon asymmetry in the universe.
- ▶ Additional sources of CP-violation is needed.
  - Explicit : like in the SM, CP is not the symmetry of the original Lagrangian.
  - Spontaneous : via VEV
    - CP is the symmetry of the original Lagrangian
    - CP is not the symmetry of the vacuum
- $\blacktriangleright$  Extension of the SM with a complex singlet and complex VEV  $\rightarrow$  spontaneous CP violation.
- We will consider an extension of the SM with a complex singlet scalar field and Iso-doublet Vector quark (cSMCS).

# The Model

- ▶ The model contains SM-like doublet  $\Phi$  and a complex singlet  $\chi$  and Iso-doublet vector quark
- ▶ Lagrangian

$$\mathcal{L} = \mathcal{L}_{gf}^{SM} + \mathcal{L}_Y(\psi_f, \Phi) + \mathcal{L}_Y(V_q, \chi) + \mathcal{L}_{scalar}, \quad \mathcal{L}_{scalar} = T - V$$

The scalar fields

$$\begin{split} \Phi &= \begin{pmatrix} \phi^+ \\ \frac{1}{\sqrt{2}} \left(\phi_1 + i\phi_4\right) \end{pmatrix}, \qquad \chi = \frac{1}{\sqrt{2}} (\phi_2 + i\phi_3). \\ &\langle \Phi \rangle = \frac{1}{\sqrt{2}} v \quad \text{and} \quad \langle \chi \rangle = \frac{1}{\sqrt{2}} w e^{i\xi} \end{split}$$

# Potential

► Singlet

$$V = V_{\Phi} + V_{\chi} + V_{\Phi\chi}$$

► SM term  

$$V_{\Phi} = -\frac{1}{2}m_{11}^{2}\Phi^{\dagger}\Phi + \frac{1}{2}\lambda\left(\Phi^{\dagger}\Phi\right)^{2}.$$
► Singlet term  

$$V_{S} = -\frac{m_{s}^{2}}{2}\chi^{*}\chi - \frac{m_{4}^{2}}{2}(\chi^{*2} + \chi^{2}) + \lambda_{s1}(\chi^{*}\chi)^{2} + \lambda_{s2}(\chi^{*}\chi)(\chi^{*2} + \chi^{2}) + \lambda_{s3}(\chi^{4} + \chi^{*4})$$

$$+\kappa_1(\chi+\chi^*)+\kappa_2(\chi^3+\chi^{*3})+\kappa_3(\chi(\chi^*\chi)+\chi^*(\chi^*\chi))$$

▶ Singlet and Doublet interaction

$$V_{\Phi\chi} = \Lambda_1(\Phi^{\dagger}\Phi)(\chi^*\chi) + \Lambda_2(\Phi^{\dagger}\Phi)(\chi^{*2} + \chi^2)$$
$$+ \kappa_4(\Phi^{\dagger}\Phi)(\chi + \chi^*).$$

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### **Constrained Potential**

To simplify the model we use U(1) symmetry

$$U(1): \Phi \to \Phi, \chi \to e^{i\alpha}\chi.$$

 $<\chi>$  spontaneous breaking symmetry  $\rightarrow$  To avoid having massless Nambu-Goldstone scalar keep U(1)-soft-breaking terms

- 1. U(1)-symmetric terms:  $m_{11}^2, m_s^2, \lambda, \lambda_{s1} \to \lambda_s, \Lambda_1 \to \Lambda$
- 2. U(1)-soft-breaking terms:  $m_4^2, \kappa_{2,3}, \kappa_4$
- 3. U(1)-hard-breaking terms:  $\lambda_{s2}, \lambda_{s3}, \Lambda_2$

$$V = -\frac{1}{2}m_{11}^2\Phi^{\dagger}\Phi + \frac{1}{2}\lambda\left(\Phi^{\dagger}\Phi\right)^2 - \frac{m_s^2}{2}\chi^*\chi + \lambda_s(\chi^*\chi)^2 + \Lambda(\Phi^{\dagger}\Phi)(\chi^*\chi)$$

$$-\frac{m_4^2}{2}(\chi^{*2}+\chi^2)+\kappa_2(\chi^3+\chi^{*3})+\kappa_3[\chi(\chi^*\chi)+\chi^*(\chi^*\chi)]+\kappa_4\Phi^{\dagger}\Phi(\chi+\chi)$$

# **Constrained Potential**

$$V = -\frac{1}{2}m_{11}^2\Phi^{\dagger}\Phi + \frac{1}{2}\lambda\left(\Phi^{\dagger}\Phi\right)^2 - \frac{m_s^2}{2}\chi^*\chi + \lambda_s(\chi^*\chi)^2 + \Lambda(\Phi^{\dagger}\Phi)(\chi^*\chi)$$

$$-\frac{m_4^2}{2}(\chi^{*2}+\chi^2)+\kappa_2(\chi^3+\chi^{*3})+\kappa_3[\chi(\chi^*\chi)+\chi^*(\chi^*\chi)]+\kappa_4\Phi^{\dagger}\Phi(\chi+\chi)$$

- ▶ Parameters →  $m_{11}^2$ ,  $m_s^2$ ,  $m_4^2$ ,  $\lambda_s$ ,  $\lambda$ ,  $\Lambda$ ,  $\kappa_2$ ,  $\kappa_3$ ,  $\kappa_4$ . If real → No explicit CP violation
- ► Vacua with spontaneous CP violation  $\langle \Phi \rangle = \frac{1}{\sqrt{2}} v$  and  $\langle \chi \rangle = \frac{1}{\sqrt{2}} w e^{i\xi} = \frac{1}{\sqrt{2}} (w_1 + iw_2)$
- ▶ Spontaneous CP violation → relevant parameters  $m_4^2$ ,  $\kappa_2$ ,  $\kappa_3$ ,  $\kappa_4$
- Positivity conditions

$$\lambda > 0, \qquad \lambda_s > 0, \qquad \Lambda + \sqrt{2\lambda\lambda_s} > 0.$$

### Mass matrix

The mass matrix that describes the singlet-doublet mixing, in the basis of natural fields  $\phi_1, \phi_2, \phi_3$ :

$$M = \begin{pmatrix} M_{11} & M_{12} & M_{13} \\ M_{21} & M_{22} & M_{23} \\ M_{31} & M_{32} & M_{33} \end{pmatrix}$$

The mass elements for mass matrix

$$\begin{split} M_{11} &= v^2 \lambda, \\ M_{12} &= v(w_1 \Lambda + \sqrt{2} \kappa_4), \\ M_{13} &= vw_2 \Lambda, \\ M_{22} &= \frac{w^2}{\sqrt{2} w_1} \left( 3 \kappa_2 + \kappa_3 (1 + 2(w_1^2 - w_2^2) / w^2) \right. \\ \left. - \kappa_4 v^2 / w^2 \right) + 2 w_1^2 \lambda_s, \\ M_{23} &= w_2 (2 w_1 \lambda_s + \sqrt{2} (-3 \kappa_2 + \kappa_3)), \\ M_{33} &= 2 w_2^2 \lambda_s. \end{split}$$

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# Mass eigenstate

▶ Diagonalization of  $M_{mix}^2$  gives the physical mass

$$RM_{mix}^2R^T = diag(M_{h_1}^2, M_{h_2}^2, M_{h_3}^2)$$

• The rotation matrix R :

$$R = \begin{pmatrix} c_1c_2 & c_3s_1 - c_1s_2s_3 & c_1c_3s_2 + s_1s_3 \\ -c_2s_1 & c_1c_3 + s_1s_2s_3 & -c_3s_1s_2 + c_1s_3 \\ -s_2 & -c_2s_3 & c_2c_3 \end{pmatrix}$$

R depends on the mixing of  $\alpha_1,\alpha_2$  and  $\alpha_3$  angles vary over an interval of length  $\pi,\,c_i=\cos\alpha_i,s_i=\sin\alpha_i$ 

Physical higgs masses

 $M^2_{h_1}\simeq v^2\lambda ~~ 
ightarrow$  The SM-like Higgs boson mass, 125 GeV

### Extremum conditions

From the potential we extract three extremum conditions:

$$-m_{11}^2 + v^2 \lambda + 2\sqrt{2}w_1 \kappa_4 + \Lambda w^2 = 0,$$

#### 2.

1.

 $w_1(-\mu_1^2 + v^2\Lambda + 2w^2\lambda_s) + \sqrt{2}[3(w_1^2 - w_2^2)\kappa_2 + (3w_1^2 + w_2^2)\kappa_3] + v^2\sqrt{2}\kappa_4 = 0,$ 

3.

$$w_2[-\mu_2^2 + v^2\Lambda + 2w^2\lambda_s + 2\sqrt{2}w_1(-3\kappa_2 + \kappa_3)] = 0,$$

When  $w_1, w_2 \neq 0$  (i.e. CP violating vacuum)  $\rightarrow$ 

 $-8m_4^2\cos^2\xi + 6R_2\cos\xi(1+2\cos 2\xi) + 2R_3\cos\xi + R_4 = 0,$ 

**CP**-violation condition

The parameters  $(R_2, R_3, R_4)$  with dimension  $[M]^2$  are:

$$R_2 = \sqrt{2}w\kappa_2, R_3 = \sqrt{2}w\kappa_3, R_4 = \frac{2\sqrt{2}v^2\kappa_4}{w}\cos\xi.$$

# Region for possible CP violation

$$-8m_4^2\cos^2\xi + 6R_2\cos\xi(1+2\cos 2\xi) + 2R_3\cos\xi + R_4 = 0$$



Figure:  $(R_2, R_3, R_4, \xi)$ , CP violation region for fix value of  $m_4^2$ 

# Region for possible CP violation

If  $R_4 = 0 \rightarrow m_4^2, R_2, R_3, \xi$   $-4m_4^2 \cos \xi + 3R_2(1 + 2\cos 2\xi) + R_3 = 0$ If  $R_4 = R_2 = 0 \rightarrow -4m_4^2 \cos \xi + R_3 = 0$ 





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A numerical analysis of the parameters of the cSMCS model

- ▶ On the allowed regions for CP violation
- ▶ In accordance with the extremum conditions
- Under the positivity and the pertubativity conditions
- ▶ In accordance with mass constraints

 $m_{h_1} \sim 125 \text{GeV}, \quad M_{h_3} \ge M_{h_2} > 150 \text{GeV}$ 

The relevant ranges of scanning:

 $-1 < \Lambda < 1, \quad 0 < \lambda_s < 1, \quad 0.2 < \lambda < 0.3, \quad 0 < \xi < \pi, \quad -1 < \rho_{2,3} < 1,$ 

 $\rho$ 's are dimensionless parameters of  $\kappa$ 's

$$-90000 \text{GeV}^2 < \mu_1^2, \ \mu_2^2 < 90000 \text{GeV}^2, \mu_1^2 = m_s^2 + 2m_4^2, \ \ \mu_2^2 = m_s^2 - 2m_4^2.$$



 $246 {\rm GeV} < v < 247 {\rm GeV}$ 



Figure:  $-0.2 < \Lambda < 0.2$ .



Figure: The correlation between masses  $M_{h_2}$  and  $M_{h_3}$  with  $\xi$ 

 $M_{h_3} > M_{h_2} > 150$ 

в	α1	$\alpha_2$	<i>α</i> 3	$M_{h_1}$	$M_{h_2}$	$M_{h_3}$	S	Т	$R^{h_1}_{\gamma\gamma}$
A1	-0.047	-0.053	1.294	124.64	652.375	759.984	-0.072	-0.094	0.98
A2	-0.048	0.084	0.084	124.26	512.511	712.407	-0.001	-0.039	0.98
A3	0.078	0.297	0.364	124.27	582.895	650.531	0.003	-0.046	0.98
A4	0.006	-0.276	0.188	125.86	466.439	568.059	-0.013	-0.169	0.92
A5	0.062	-0.436	0.808	125.21	303.545	582.496	0.002	-0.409	0.81
A6	-0.210	0.358	0.056	124.92	181.032	188.82	0.003	-0.010	0.82
A7	-0.205	0.403	0.057	125.01	175.45	178.52	0.002	-0.020	0.81

Table: Benchmark points A1 - A7. Masses are given in GeV and  $\alpha$  are the mixing angles

These are in agreement with current data for S and T parameters,

$$S = 0.05 \pm 0.11$$
,  $T = 0.09 \pm 0.13$ ,

and in agreement with  $R_{\gamma\gamma}$  from ATLAS and CMS,

$$R_{\gamma\gamma} = 1.14^{+0.27}_{-0.25}(\text{ATLAS})$$
 and  $R_{\gamma\gamma} = 1.14^{+0.27}_{-0.25}(\text{CMS})$ 

M. Krawczyk and ND,"CP violation in the Standard Model with a complex singlet", arXiv:1603.00598

 $\rm M.R.$  Masouminia and ND,"A phenomenological study on the production of Higgs bosons in the cSMCS model at the LHC," arXiv:1611.03312v1

# The Electroweak Phase Transition

T

$$V(T_0) = -\frac{1}{2}m_{11}^2\Phi^{\dagger}\Phi + \frac{1}{2}\lambda_1\left(\Phi^{\dagger}\Phi\right)^2 - \frac{\mu_1^2}{4}\phi_2^2 - \frac{\mu_2^2}{4}\phi_3^2 + \frac{1}{2}\Lambda(\Phi^{\dagger}\Phi)(\phi_2^2 + \phi_3^2) + \frac{1}{4}\lambda_s(\phi_2^2 + \phi_3^2)^2 + \frac{1}{\sqrt{2}}\kappa_2(\phi_2^3 - 3\phi_2\phi_3^2) + \frac{1}{\sqrt{2}}\kappa_3(\phi_2^3 + \phi_2\phi_3^2) + \sqrt{2}\kappa_4(\Phi^{\dagger}\Phi)\phi_2.$$

The one-loop thermal corrections to the effective potential at finite temperature T

$$\Delta V_{thermal} = \sum_{i} \frac{n_i T^4}{2\pi^2} I_{B,F} \left(\frac{m_i^2}{T^2}\right),$$

 $\operatorname{with}$ 

$$I_{B,F}(y) = \int_0^\infty dx \ x^2 \ ln \left[ 1 \mp e^{-\sqrt{x^2 + y}} \right]$$

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The potential at finite temperature T

$$V(v(T),T) = V(v_0) + \Delta V_{thermal}$$

The high temperature approximation

V

$$\begin{aligned} (T) = & \frac{1}{2}\overline{m}_{11}^2 \Phi^{\dagger} \Phi + \frac{1}{2}\lambda(\Phi^{\dagger} \Phi) + \frac{\overline{\mu}_1^2}{4}\phi_2^2 + \frac{\overline{\mu}_2^2}{4}\phi_3^2 \\ & \frac{1}{2}\Lambda(\Phi^{\dagger} \Phi)(\phi_2^2 + \phi_3^2) + \frac{1}{4}\lambda_s(\phi_2^2 + \phi_3^2)^2 \\ & + \kappa_2 \frac{1}{\sqrt{2}}(\phi_2^3 - 3\phi_2\phi_3^2) + \kappa_3 \frac{1}{\sqrt{2}}(\phi_2^3 + \phi_2\phi_3^2) \\ & + \sqrt{2}\kappa_4(\Phi^{\dagger} \Phi)\phi_2 + \overline{\kappa}_{34}\frac{T^2}{3}\phi_2, \end{aligned}$$

$$\begin{split} \overline{m}_{11}^2 &= -m_{11}^2 + (3\lambda \ + \Lambda + \frac{2m_W^2 + m_Z^2 + 2m_t^2}{2v^2})\frac{T^2}{3}, \\ \frac{1}{2}\overline{\mu}_1^2 &= -\frac{1}{2}\mu_1^2 + (\Lambda + 2\lambda_s)\frac{T^2}{3}, \\ \frac{1}{2}\overline{\mu}_2^2 &= -\frac{1}{2}\mu_2^2 + (\Lambda + 2\lambda_s)\frac{T^2}{3}, \\ \overline{\kappa}_{34} &= \sqrt{2}(\kappa_3 + \kappa_4). \end{split}$$

Extermum conditions of the effective potential

$$\begin{split} V(T) = & \frac{1}{2}\overline{m}_{11}^2 \Phi^{\dagger} \Phi + \frac{1}{2}\lambda(\Phi^{\dagger} \Phi) + \frac{\overline{\mu}_1^2}{4}\phi_2^2 + \frac{\overline{\mu}_2^2}{4}\phi_3^2 \\ & \frac{1}{2}\Lambda(\Phi^{\dagger} \Phi)(\phi_2^2 + \phi_3^2) + \frac{1}{4}\lambda_s(\phi_2^2 + \phi_3^2)^2 \\ & + \kappa_2 \frac{1}{\sqrt{2}}(\phi_2^3 - 3\phi_2\phi_3^2) + \kappa_3 \frac{1}{\sqrt{2}}(\phi_2^3 + \phi_2\phi_3^2) \\ & + \sqrt{2}\kappa_4(\Phi^{\dagger} \Phi)\phi_2 + \overline{\kappa}_{34}\frac{T^2}{3}\phi_2, \end{split}$$

 $\bullet \ \overline{m}_{11}^2 + \lambda v^2 + \Lambda w^2 + 2\sqrt{2}\kappa_4 w_1 = 0,$ 

- $\blacktriangleright \ w_1(\overline{\mu}_1^2 + \Lambda v^2 + 2\lambda_s w^2) + \sqrt{2} \left[ 3\kappa_2(w_1^2 w_2^2) + \kappa_3(3w_1^2 + w_2^2) + \kappa_4 v^2 \right] + \frac{2}{3}\overline{\kappa}_{34}T^2 = 0,$
- $w_2[\overline{\mu}_2^2 + \Lambda v^2 + 2\lambda_s w^2 + 2\sqrt{2}(-3\kappa_2 + \kappa_3)w_1] = 0.$

# Numerical solution for strong enough first order EWPT

Conditions for strong enough first order EWPT:

$$\blacktriangleright V_{eff}(v(T_c), T_c) = V_{eff}(0, T_c)$$

 $\blacktriangleright \ \frac{v(T_c)}{T_c} \ge 1$ 



$$\begin{split} -0.2 < \Lambda < 0.2, \\ 0.2 < \lambda_s < 1, \\ -1 < \rho_{2,3,4} < 1, \\ 0 < \xi < \pi, \\ -90000 \text{GeV}^2 < \mu_i^2 < 90000 \text{GeV}^2, \\ M_{h_1}^2 \approx m_{11}^2 \approx \lambda v^2 (M_{h_1} \approx 125 \text{GeV}), \\ 0.2 < \lambda < 0.3. \end{split}$$

• This model provides a strong enough first order EWPT.

## **Baryon** Generation

Baryon asymmetry resulting from a mixing of the SM quarks and heavy vector quarks at high temperature [McDonald 1995] (low temperature [Branco 1998])

• 
$$\mathcal{L}_Y(V_q, \chi) = \lambda_v \chi \bar{Q_L} V_R + M \bar{V_L} V_R + h.c.$$

$$\begin{bmatrix} Q'_L \\ V'_L \end{bmatrix} \qquad = \begin{bmatrix} a & b \\ -b^* & a^* \end{bmatrix} \begin{bmatrix} Q_L \\ V_L \end{bmatrix}$$

$$a \qquad = \begin{bmatrix} 1 + \left(\frac{\lambda_v w}{M}\right)^2 \end{bmatrix}^{-1/2}$$

$$b \qquad = \left(\frac{\lambda_v w}{M}\right) \left[ 1 + \left(\frac{\lambda_v w}{M}\right)^2 \right]^{-1/2} e^{-i\xi}$$

Diagonalizing the mass terms results in some non-diagonalized kinetic terms.

$$\begin{split} \overline{Q}_L i \gamma^\mu \partial_\mu Q_L + \overline{V}_L i \gamma^\mu \partial_\mu V_L \rightarrow \overline{Q'}_L i \gamma^\mu \partial_\mu Q'_L + \overline{V'}_L i \gamma^\mu \partial_\mu V'_L + \Delta \mathcal{L}_k + const \\ \Delta \mathcal{L}_k &= -\frac{\lambda_V^2 w^2}{M^2} \dot{\xi} (\overline{Q'}_L \gamma^0 Q'_L - \overline{V'}_L \gamma^0 V'_L) \end{split}$$

The phase of singlet should be time-dependent.

### **Baryon** Generation

The amount of baryon asymmetry:

$$n_B = -N_f \int \frac{\Gamma_{sph}(T)}{2T} \mu_B dt$$

 $N_f$  is the number of flavors in the model The sphaleron rate  $\Gamma_{sph} \to \Gamma_{sph} = K(\alpha_W T)^4$  at high temperature

$$\mu_B = -\frac{5}{6} \frac{\lambda_V^2 w^2}{M^2} \dot{\xi}$$

$$s = \frac{2\pi^2}{45}g^*T^3$$

 $g^{\,\ast}\,\sim\,100~$  the effective number of degrees of freedom therefore,

$$\frac{n_B}{s} = \frac{225K\alpha_W^4}{4\pi^2 g^*} \frac{\lambda_V^2 w^2}{M^2} \delta\xi$$

The observations from WMAP  $\rightarrow \frac{n_B}{s} = 8.7 \pm 0.3 \times 10^{-11}$ 

**Baryon** Generation

$$\frac{n_B}{s} = \frac{225K\alpha_W^4}{4\pi^2 g^*} \frac{\lambda_V^2 w^2}{M^2} \delta\xi = 8.7 \pm 0.3 \times 10^{-11}$$
$$K \frac{\lambda_V^2 w^2}{M^2} \delta\xi = 1.14 \pm 0.3 \times 10^{-3}$$



0.3 TeV < M < 13 TeV 2GeV < w < 400 GeV 0  $< \lambda_V <$  1 0  $< \delta\xi < \pi$ 

• This model provides an acceptable number for the generation of baryons in the Universe.

ND, "Baryogenesis of the Universe in cSMCS Model plus Iso-Doublet Vector Quark", arXiv:1608.02820

### Summery and Conclusion

- ▶ This model contains a SU(2) doublet as in the SM and a complex singlet with a complex VEV and iso-doublet vector quark.
- ▶ This model provides a source of spontaneous CP violation.
- This model provides a strong first order phase transition and an acceptable number for the generation of baryons in the Universe.
- The analysis of this model was performed as a part of full analysis of the cIDMS (Inert Doublet Model plus Complex Singlet) which was confronted with LHC data for 125 GeV, precision data STU as well as astro data on dark matter. Bonilla, Sokolowska, Diaz-Cruz, Krawczyk and ND, "IDMS: Inert Dark Matter Model with a complex singlet," arXiv:1412.8730.